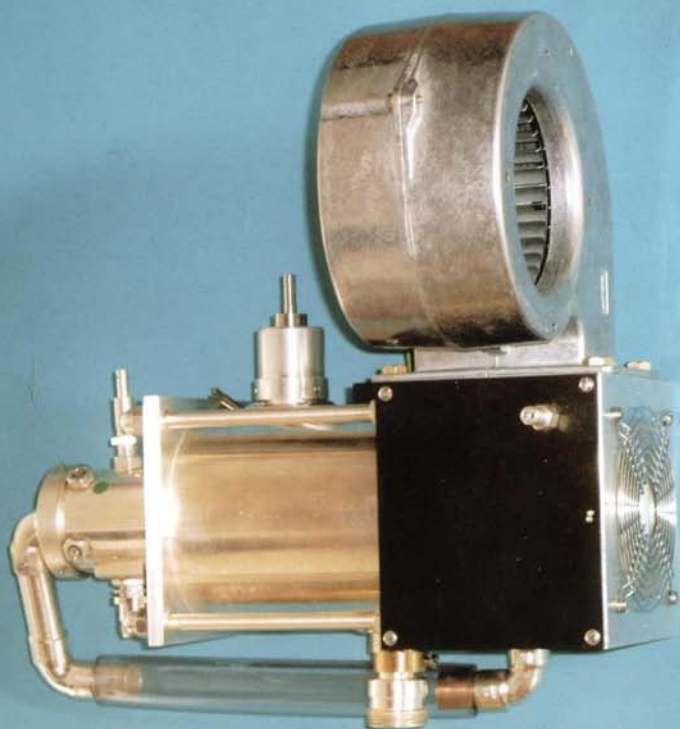


Vol. 33

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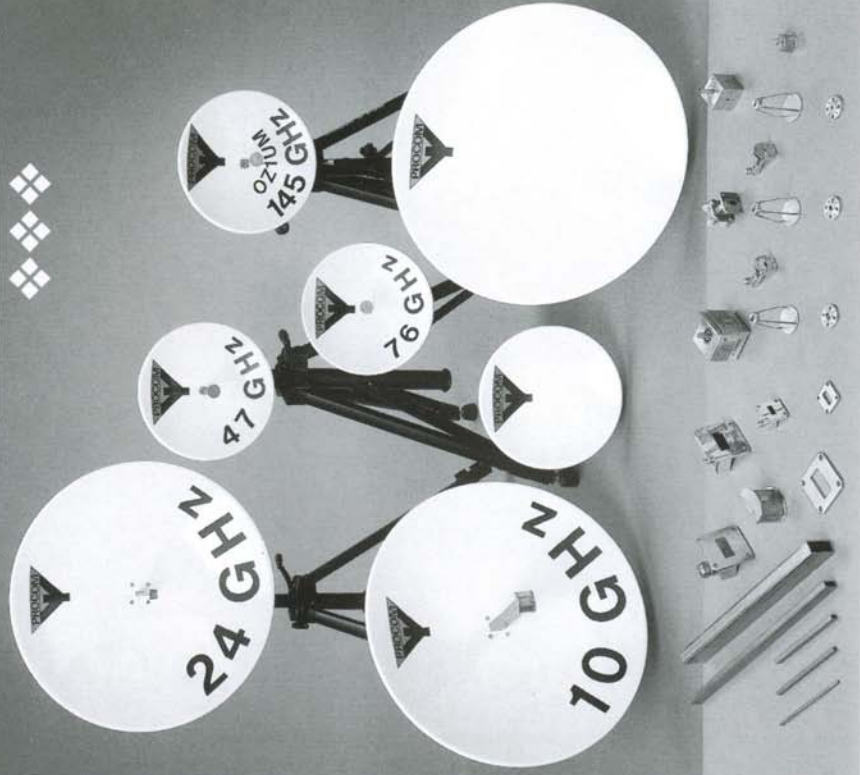
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600 Watt 23cm PA by DJ3FI

MICROWAVE



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e-mail: eisch-electronic@t-online.de

www.eisch-electronic.com

Inhalt / Content

1/04

Title: DJ3FI's 23cm Power Amplifier

DUBUS Order Form	5
Editorial	8
Technical Reports	
600 Watt 23cm Power Amplifier with YD1336 by Hubert Tiedmann, DJ3FI	9
Using MRF284 a Motorola LDMOSFET on our bands by Franck Rousseau, F4CIB	34
28 V Switched Power Supply for LDMOS-PA's, by Achim Vollhardt, DH2VA	38
L-Band Oscillator Chain, by Daniel Uppström, SM6VFZ	45
Practical and Cheap Solution of a Circular Waveguide for 10 GHz by Ivo Chladek, ZS6AXT, OK2WCG	49
New Software "UKW Tools", by Jonathan Naylor, G4KLX	54
Microwave Parts	56
Microwave Europe	60
Microwave Japan	62
Microwave USA	65
FAI News	66
2m - EME News	67
70cm + up - EME News	69
Tropo News	78
6m News	82
Meteor Scatter Reports	84
Meteor-Shower Calendar April - June 2004	88
Aurora Reports	89
New Products	99
News & Comments	100
DUBUS Content 2003	105
DUBUS Beaconlist 2004 for 6m, 4m, 2m and 70cms	106

Redaktion / Editorial Staff

Technical Reports / Team of authors

Ian White, G3SEK, 52 Abingdon Rd, Drayton, Abingdon, Oxon,
OX14 4HP email: G3SEK@ifwtech.co.uk
Michael Kuhne, DB6NT, Scheibenacker 3, D-95180 Berg, Germany
email: kuhne.db6nt@t-online.de
Luis Cupido, CT1DMK, R. S. Joao, Lote 5, St. Joana, 3810 Aveiro,
Portugal cupido@mail.ua.pt
Leif Åsbrink, SM5BSZ, Jäders Prästgård, S-635 05 Eskilstuna,
Sweden email: leif.asbrink@mbox300.tele2.se
Werner Rahe, DC8NR, Conrad-Schulz-Weg 5, D-82211 Herrsching

Microwave Parts

Hennig Weddig, DK5LV, Konsul-Lieder-Allee 11, D-24226 Heikendorf
email: Henning-Christof.Weddig@desy.de

EME-News, 70cm + up

Bernd Wilde, DL7APV, Nordhoferstr. 8, D-16833 Koenigshorst,
Germany. email: DL7APV@t-online.de

EME-News, 2m

DF2ZC, Bernd Mischlewski, Auf dem Scheid 36, D-53547 Burscheid
Germany. Email: BerndDF2ZC@t-online.de

EME-Email-List 70cm + up

Klaus Tiedemann, DL4EBY, Halskestr. 35, D-12167 Berlin, Germany
email: tklaus@snafu.de

Microwave Europe

Simon Lewis, GM4PLM, Creoch Farm, Ochiltree Ayrshire, KA18 2QH,
Scotland, U.K. simon@creoch.freeserve.co.uk simon.lewis@vodafone.com

Microwave Japan

JH6RTO (also 7L3TDU/M0RAA/AH0R), Seiji Fukushima, 1182-2505
Hase, Atsugi-city 243-0036 Japan. Email: jh6rto@m.ieice.org

Microwave USA

Kent Britain, WA5VJB, 1626 Vineyard Rd., Grand Prairie, TX 75052, USA
Email: wa5vjb@flash.net

Rain Scatter Reports

Jonathan Naylor, G4KLX/HB9DRD, 24 Castle View Drive, Cromford,
Matlock, DE4 3RL, England, U.K. NaylorJS@yahoo.com

FAI Reports

Dom Dehays, F6DRO, Le Fourtou, F-31450 Montesquieu Lauragais
France. email: f6dro@aol.com or dominique.dehays@enac.fr

Aurora Reports

Stefan Heck, LA0BY, Hildsvej 5, N-1349 Rykkinn, Norway
email: la0by@darcd.de

Meteor-Scatter Reports Es-Reports, Top List

Norbert Goettsche, DL8LAQ, Tulpenweg 3, D-24568 Kaltenkirchen
email: DL8LAQ@sh-home.de

Meteor Shower List

Guido Juenkersfeld, DL8EBW, Gustav-Freitag-Str. 1, D-42327
Wuppertal, Germany. Email: guy@dl8ebw.de

Tropo Reports

Wolfgang Schneider, DJ8ES, Lohweg 14, D-25591 Ottenbüttel
email: DJ8ES@GMX.de

6m-News, Beacon List New Products News & Comments

Joachim Kraft, DL8HCZ, Gruetzmuehlenweg 23, D-22339 Hamburg
email: funktelegramm@t-online.de info@dubus.de

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DUBUS-Magazine, Verlag Joachim Kraft,
DL8HCZ, Grüzmühlenweg 23,
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email: DUBUS@web.de info@dubus.de
Phone/FAX: ++49 40 538 31 86

Abo-Verwaltung / Circulation:

Norbert Goettsche, DL8LAQ, Tulpenweg 3,
D-24568 Kaltenkirchen, Germany.
Phone: ++49 4191 959913
email: DUBUS@web.de info@dubus.de

Abo-Konto / Subscription-Account

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DUBUS Distributoren

International DUBUS Representatives

Austria	Reinhilde Riml, OE9YTV, Marktstr. 35, A-6971 Hard, Austria. email: riml.reinhilde@aon.at
Australia	Doug McArthur, VK3UM, Tikaluna, 26 Old Murrindindi Rd., Glenburn, Victoria 3717. email: tikaluna@yccs.com.au
Belgium	InfoRad c/o Jan Galicia ON6JG, Oude Gendarmeriestraat 62, B-2220 Heist-op-den-Berg. email: inforad@pandora.be
Bulgaria	Christo Mintchev, LZ1DP, Ul Boruigrad 8 AP 10, 6004 Stara Zagora email: lz1dp@stz.orbitel.bg
Czech Republic	GES Electronics, OK1FM, Sady S. Kvetna, 30100 Plzen, Tel: +420 37 73 73 111 email: ges@ges.cz
Denmark	Gert Rahbek-Udengaard, OZ1FKZ, Vurringvej 60, Vurring, DK-8660 Skanderborg Tel.: 0045 8657 2660, email: oz1fkz@fritid.tele.dk
Eire	Albert E. Latham, EI6AS, 226 the Rise, Belgard Heights Tallaght Dublin 24
France	Fabienne Ansel, La Grave, F-31560 Seyre, France Tel.: 05 34 66 10 09 email: dubus.mag@free.fr
Finland	Jukka Hietaranta, OH1FF, Kuutie 25, FIN-36110 Ruutana Fax: 03 2619255, email: Jukka.Hietaranta@elektrobit.com
Italy	Tony Fumagalli, I2FUM, Via Masia 34, I-22100 Como Tel: 031571537, email: elcosrl@tin.it a.fumagalli@elcocomo.it
Japan	Ryuichiro Nagaoka, JA3MKS, 15-18-704, 2 chome, Higashi-izumioka, Toyonaka-shi, Osaka, 561-0864 Japan, Tel/Fax: +81-6-6850-1890, E-mail: mks@x.age.ne.jp
Malta	Joseph Falzon, 9H1VW, « Babs » 104, Sir William Reid Street, Gzira, GZR02, Malta danielg@maltanet.net
Netherlands	Veron Verkoop Bureau, P.O. Box 1166, NL-6801 BD Arnhem email: veroncb@worldonline.nl
Norway	Anders Thorrud, LA8LF, Jerpeveien 20, N-3030 Drammen email: LA8LF@tiscali.no
Portugal	Antonio Filipe S. Ferreira, CT1DDW, B S Eugenia 28-1 Post Viso, P-3500 Viseu, Portugal email: filipeferreira@net.sapo.pt
Poland	Andrzej Włodarczyk, SP1JX, P.O. Box 67, 76-270 Ustka
Romania	Folea Ion, YO5TE, P.O.Box 168, RO-3400, Cluj 1 email: yo5te@yo5kai.codec.ro
Russia	Sergej Joutiaev, RW3BP, P.O. Box 462, 111555 Moscow, email: rw3bp@co.ru
R.S.A.	Katti Venter, ZS6KVR, Postnet suite 534, Private bag X4, Menlo Park, 0102
+ Namibia	email: zs6kvr@ananzi.co.za phone: +27 12 9934848
Spain	J.R. Daglio Accunzi, EA2LU, Manual Irribarren 2-5 D, E-31006 Pamplona email: ea2lu@telefonica.net
Serbia	Nemethy Istvan, YU7EW, Branka Radicevica 63, YU-23217 Aleksandrovo email: yu7ew@ptt.yu
Slovenia	Miha Habic, S51FB, Vodnikova 8, SI-1000 Ljubljana email: Miha.Habic@mors.si
Sweden	Anders Pettersson, SM7ECM, Hökvägen 111, SE-24562 Hjärup email: Anders.Pettersson@sonyericsson.com
Switzerland	Ferdinand Stämpfli, HB9MIO, Pilatusstr. 1, CH-5703 Seon. Tel.: 079/4734650 stagra@bluewin.ch
Ukraine	Sergej Lysenko, UR5LX, Zolochov Reg., Iwaschki, Kharkov Obl, Ukr. 312211 email: ur5lx@sergej.kharkiv.com
United Kingdom	Roger Blackwell, G4PMK, 5 Tollgate Road, Culham, Nr. Abingdon, Oxon OX14 4NL email: dubus@marsport.demon.co.uk
USA	Kyle Britain, KB5UBE, 1626 Vineyard Road, Grand Prairie, TX 75052 email: dubususa@hotmail.com
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Editorial

Dear DUBUS Reader!

Here is the first issue for the new year 2004 with a couple of hopefully interesting technical articles. This issue focuses a bit on the 23cm band - a band that becomes more and more popular for DX and EME, probably also due to the higher noise and RFI problematic on 2m and 70cms.

But we have some more news for you! After a break of six years a new DUBUS TECHNIK book is available now. It is number 6 and in the same size and format like the previous ones (see also the back cover). Order your copy now if you want to be sure to get one. Technik books 1, 2, 4 and 5 are sold out and there is only a limited number of Technik 3 available.

We now have a new DUBUS distributor for France. Patrick, F6HYE, has "retired" after 16 years of successful work for DUBUS - thank you, Patrick! And much fun with your new hobby! Jose Balanca, F1EIT together with his YL Fabienne Ansel do the job now and made a good start already with many French subscribers for 2004. Welcome Fabienne and Jose!

After many years without any distributor for Malta now we have Joseph Falzon, 9H1VW, who will help us with the DUBUS representation in this country. Welcome, Joseph!

In the past we had several requests regarding the promised follow-up parts of the oscillator article from Gerhard, DK4XP. The agreement with DK4XP about this series of articles was made a long ago, but unfortunately Gerhard has many other commitments now, and is not able to continue. We apologize for this but it is beyond our control. Anyway the article from SM6VFZ in this issue brings some practical results based on DK4XP's articles.

From the DX operating side of the DUBUS activities we have positive news: with the kind help of DF7KF, DL8YHR and others, 5T5SN will erect a 6m and a 2m beacon in Mauretania very soon (Locator IK28ac). We think that this will be an interesting location for Tropo and Es towards Europe.

Please keep sending your technical articles, activity reports and info. DUBUS needs YOUR contribution. Many thanks!

Vy 73 from Joe, DL8HCZ / CT1HZE
and the DUBUS team!

Liebe DUBUS Leser!

Hier ist die erste Ausgabe für das neue Jahr 2004 mit einigen -hoffentlich- interessanten technischen Artikeln. In dieser Ausgabe liegt der Schwerpunkt etwas auf dem 23cm-Band. Dieses Band wird immer populärer für DX und EME. Wahrscheinlich liegt dies u.a. auch in der zunehmend problematischer werdenden Stör- und TVI-Situation auf 2m und 70cm.

Aber wir haben weitere Neuigkeiten: Nach einer sechsjährigen Pause gibt es jetzt ein neues DUBUS Technik Buch. Siehe auch Heftrückseite. Es ist die Nummer 6 und entspricht in Größe und Umfang den vorherigen Ausgaben. Wir empfehlen, jetzt ein Exemplar zu bestellen, wenn Sie sicher sein wollen, eines zu erhalten. Die Technik-Bücher 1, 2, 4 und 5 sind ja bereits ausverkauft. Von Technik 3 gibt es noch einige Restexemplare.

Wir haben einen neuen Distributor für Frankreich. Patrick, F6HYE, hat nun nach 16 Jahren erfolgreicher Arbeit für DUBUS aufgehört und widmet sich einem neuen Hobby. Danke, Patrick! Jose Balanca, F1EIT, ist nun zusammen mit seiner YL Fabienne Ansel neuer Distributor für Frankreich und ist bereits mit vielen Abonnenten für 2004 sehr gut gestartet. Willkommen im Team, Fabienne und Jose!

Nach vielen Jahren ohne Distributor in Malta konnten wir nun Joseph Falzon, 9H1VW, dafür gewinnen. Er wird DUBUS in diesem Land vertreten. Willkommen, Joseph!

In der Vergangenheit haben wir einige Anfragen nach der versprochenen Fortsetzung des Oszillator-Artikels von Gerhard, DK4XP, erhalten. Leider hat Gerhard zwischenzeitlich viele andere Verpflichtungen und kann den Artikel nicht fortsetzen. Wir möchten uns dafür entschuldigen, können es aber leider nicht ändern. Die Absprache mit DK4XP wurde vor langer Zeit seitens des ehemaligen Chefredakteurs der DUBUS gemacht. Immerhin setzt der Artikel von SM6VFZ in dieser Ausgabe einiges von DK4XP's Artikeln in die Praxis um.

Nun zur praktischen Seite der DUBUS-Aktivitäten. Hier gibt es Positives zu berichten: Dank der Hilfe von DF7KF, DL8YHR u.a. wird 5T5SN in Mauretanien (IK28ac) in Kürze eine 6m und 2m-Bake in Betrieb nehmen. Wir denken, dass dies ein interessanter Baken-Standort für Es und Tropo Richtung Europa ist.

Wir bitten, weiter so zahlreiche technische Artikel und Aktivitätsberichte einzusenden. Vielen Dank!

Vy 73 von Joachim, DL8HCZ / CT1HZE
und vom ganzen DUBUS-Team!

Verstärker für das 23cm-Band mit der Triode YD 1336

Hubert Tiedmann, DJ3FI

1. Einleitung

Die gewählte Röhre wurde für Fernsehsumsetzer entwickelt. Hier haben die Transistoren wegen ihrer unbegrenzten Lebensdauer die Röhren verdrängt. VHF-UHF Amateure bestreiten ihre Konteste und EME-QSO's überwiegend mit Röhren-Endstufen. Die begrenzte Lebensdauer der Röhre (10.000 – 20.000 Stunden) reicht auch bei gebrauchten Röhren noch für genügend Konteste. Es macht daher Sinn, einen Röhrenverstärker für das 23 cm-Band zu betreiben, da die Röhre bei diesen Leistungen auch einige Vorteile hat.

Der Sockel der Röhre ist baugleich mit den Typen TH 308, TH 329, YD 1333 bis YD 1337.

Die Anodenverlustleistung dieser Röhren reicht von 700 W bei der TH 308 bis 1900 W bei der YD 1335. Entsprechend sind die Baulänge und der Durchmesser des Anodenradiator sowie die inneren Kapazitäten unterschiedlich. Der Eingangskreis hat einen weiten Abstimmbereich, mit ihm lassen sich alle genannten Röhren anpassen. Soll die TH308 eingesetzt werden, ist lediglich der Anodentopf um 1 cm zu verlängern.

2. Aufbau des Verstärkers

Der Verstärker ist in coaxialer Topfkreis-Technik aufgebaut. Hierbei arbeitet die Triode in Gitterbasis-Schaltung. Zwei coaxiale Resonanzkreise sind ineinander geschachtelt. Es wurden Messingrohre mit dem auf den erforderlichen Wellenwiderstand abgestimmten Durchmesser verwendet. Die Kontaktierung der Koaxanschlüsse erfolgt an der Anode und am Gitter über Federkontaktstreifen. Beim Kathoden- und Heizungsring sind die Rohre zur Kontaktierung geschlitzt. Bodenplatte und Anodenplatte sind ebenfalls mit Fingerkontakten versehen.

Der Luftschacht über dem Anodenradiator besteht aus einem Teflonring und einem geschlitzten Aluminiumring. Über den Aluminiumring wird auch die Anodenspannung zugeführt. Der Luftschacht ist an der oberen Abdeckung beweglich angebracht damit kein seitlicher Druck auf die Röhre ausgeübt wird.

3. Der Ausgangskreis

Der coaxiale Schwingkreis zwischen Gitter (Masse) und Anode arbeitet im $3/4 \lambda$ Mod. Abgestimmt

wird kapazitiv mit einem Stempel im ersten Spannungsmaximum ca. 30 – 50 mm von der unteren Bodenplatte. Die Auskopplung der Sendeleistung erfolgt kapazitiv am oberen Teil des Topfes. Mit diesem kapazitiven Spannungsteiler wird die Transformation vom Arbeitswiderstand der Röhre (ca. 2,5 k Ω) auf den Wellenwiderstand der Leitung vorgenommen. Die Auskopplung am oberen Topfende hat den Vorteil, dass dort für alle Frequenzen Spannungsmaximum ist, so dass der Kreis im $1/4 \lambda$ Mod ebenfalls bedämpft wird und somit in der Nähe von 430 MHz nicht schwingen kann. Um Eigenschwingung zu vermeiden, ist es wichtig, dass die Kontaktfedern am Gitterrohr am gesamten Umfang fest anliegen damit keine Energie vom Ausgangskreis auf den Eingang gelangt. Der obere Teil des Topfes wird von der Abschlussplatte, der Röhre (Anode) und der mit 0,1 mm starken Kaptonfolie isolierten Messingplatte abgeschlossen. Diese Kondensatorplatte hat eine Kapazität von 600 – 800 pF. Sie wird mit einem Teflonring zentriert und mit einer 0,2 mm Teflonfolie und einer Messingplatte abgedeckt. Damit ist der Topf HF-dicht abgeschlossen.

4. Der Eingangskreis

Die Röhre wird zwischen Gitter (Masse) und Kathode angesteuert. Mit dem Koaxkreis im Eingang wird die Eingangsimpedanz der Kathode auf den Wellenwiderstand des Kabels transformiert. Weiterhin müssen die Heizleistung zugeführt und die Kathode sowie der Gitterring gekühlt werden. Das Koaxsystem (Gitter und Kathodenrohr) hat die elektrische Länge von 1 λ . Abgestimmt wird mit einem Alu-Ring, der isoliert und beweglich als Schieber ausgeführt ist. Befindet sich der Schieber in einem Spannungsmaximum, wirkt er überwiegend kapazitiv und damit verlängernd auf den Kreis. Im Strommaximum wirkt er als Stromverdrängungsschieber und vermindert die Induktivität und damit den Wellenwiderstand, was den Kreis verkürzt. So ergibt sich ein verlustarmer Resonator mit weitem Abstimmbereich, so dass sich alle Röhren dieser Baureihe anpassen lassen. Am offenen Ende des Koaxsystems wird mittels einer Scheibe an einem N Stecker kapazitiv eingespeist. Der Abstand der Scheibe zum Kathodenrohr bestimmt die Spannungsteilung und damit die Transformation von dem Widerstand der Röhre auf den Wellenwiderstand des Kabels (50 Ω). Voraussetzung ist, der Kreis ist mit dem Schieber auf Resonanz abgestimmt, sodass am Einspeisepunkt ein Spannungsmaximum entsteht.

Im Kathodenrohr befindet sich durch Teflonbuchsen isoliert das Rohr für den Heizkontakt. Die Heizspannung wird über Drosseln und Durchführungskondensatoren dem Kathodenrohr und dem Heizungsrohr zugeführt. Die Kühlluft für die

Kathodenseite wird durch das innere Rohr zugeführt und verlässt den Kathodenraum über seitliche Bohrungen in der Fassung. Dann strömt die Luft durch das Gitterrohr und tritt durch Bohrungen in der Bodenplatte aus.

5. Der Kathodenspannungsregler

Soll die Röhre bei einer Anodenspannung von 3 kV und einem Ruhestrom von 80 mA betrieben werden, so ist eine negative Gittervorspannung von ca. 38 V nötig. Um diesen Betrag muss die Kathode positiv vorgespannt werden. Zur Vermeidung von Überschlügen in der Röhre sollte auch in Ruhe ein kleiner Strom von 3 – 10 mA fließen. Dieser Strom wird durch den Widerstand R1 bestimmt.

Wird das TX Relais aktiviert, so stellt sich ein Strom von ca. 70 mA ein, der von den Widerständen R2 + R3 bestimmt wird. Soll ein höherer Strom fließen, so muss der Abgriff am Widerstand R2 in Richtung Kathodenanschluss verschoben werden bis die Zenerspannung von D1 überschritten ist und der Transistor angesteuert wird. So lässt sich der Arbeitspunkt auf 80 mA und höher einstellen. Der Widerstand R6 übernimmt einen Teil der Last des Reglers und begrenzt den Strom durch die Röhre, wenn der Transistor durch einen Überschlag in der Röhre einen Emitter-Kollektor-Schluss bekommt.

6. Einschalten und Abgleichen

Wenn Lüfter, Heizung und Hochspannung einschaltbereit sind, werden die Grundeinstellungen vorgenommen. Der Koppelstempel des Eingangs-kreises wird auf 1 mm Luftspalt zum Kathodenrohr eingestellt und der Auskoppelstempel auf 2 mm zum Gitterrohr. Dann wird die Schelle von R2 für den Spannungsteiler in die Mitte geschoben. In die Anodenspannungs-Zuführung wird ein Widerstand von ca. 1 kΩ geschaltet. Nun können der Lüfter und die Heizung eingeschaltet werden.

Da die Röhren alle älter als 10 Jahre sind, ist es nötig, sie ca. 30 Minuten zu heizen, um dann die Anodenspannung über einen Hochlastwiderstand aufzuschalten.

Dieses ist wegen der Erwärmung des Widerstandes nicht lange möglich, sollte dann unterbrochen und wiederholt werden, so werden Schlüsse festgestellt und Spannungen und Ströme überprüft.

Ist der Ausgang des Verstärkers mit Lastwiderstand und Leistungsmesser abgeschlossen, kann die volle Hochspannung eingeschaltet werden.

Zum ersten Abgleich steuert man mit 5-10 W an. Dann wird der Schieber bewegt, bis der Anodenstrom auf ein Maximum angestiegen ist. Nun lässt sich der Anodenkreis auf maximale Ausgangsleistung trimmen. Die Reflektionswerte des Einganges sollte mit Richtkoppler gemessen auf minus 16 – 20 dB abgeglichen werden. Dieses geschieht durch verändern der Kopplung. Da bei einer Gitterbasisschaltung der Ausgangswiderstand und der Eingangswiderstand mit der Röhre in Reihe liegen, hat die Auskopplung und Abstimmung auf der Anodenseite einen Einfluss auf die Anpassung des Einganges. Der Abgleich des Einganges ist nach der Optimierung der Ausgangsleistung erneut zu überprüfen.

Zur Optimierung der Ausgangsleistung folgender Hinweis: Nähert man den Koppelstempel dem Gitterrohr (festere Kopplung), so wird der Arbeitswiderstand kleiner. Die Verstärkung geht zurück, an der Anode ist die Spannung größer, dadurch sinkt der Gitterstrom. Bei lockerer Kopplung verhält es sich umgekehrt. Man hat höchste Verstärkung, muss aber die Ansteuerung zurücknehmen, da der Gitterstrom schnell unzulässig hoch wird. Durch eine zu lockere Kopplung wird die Betriebs-güte zu hoch und der Kreis wegen der geringen Bandbreite leichter verstimmbar durch thermische Veränderungen in der Röhre. Durch Veränderung der Kopplung lässt sich für eine gegebene Steuerleistung die Einstellung für höchste Ausgangsleistung finden.

7. Gemessene Werte / Measured Results

Siehe untenstehende Tabelle.

See table below.

Pin (W)	Pout (W)	g (db)	Ua (kV)	Ia (mA)	Uk (V)	Ig 1 (mA)	Q An. (W)	η (%)
0	0	-	3,1	80	39	0	248	-
4,8	200	16,1	3,0	520	46	0	1360	12,8
11	400	15,6	2,9	750	47	0	1775	18,3
28	600	13,3	2,8	880	49,5	22	1864	24,3
35	750	13,3	2,8	980	52	30	1994	27,3

8. Röhrendaten / Tube Data

Parameter	Sym bol	Unit	Wert - YD 1334/1336	Wert YD 1335
Heiz- spannung	Uf	V	6 – 6,3	6 – 6,3
Heizstrom	If	A	4,8 -5,8	4,8 -5,8
Kapazität Anode- Gitter	Cag	pF	6,8 – 8	6,8 – 8
Kapazität Gitter-Kath.	Cgk	pF	20 – 30	20 – 30
Kapazität Anode-Kath.	Cak	pF	0,09 – 0,18	0,09 – 0,18
Steilheit	s	mA/ V	ca. 70	ca. 70
Anodensp.	Ua	V	max. 3500	max. 3500
Anodenverl.- Leistung	Qa	W	1800	1900
Gitterstrom	Ig	mA	5	5
Kathoden- strom	Ik	mA	max. 700	max. 800
Gitter- spannung	Ug	V	max. – 200	max. – 200
Vorheizzeit	th	sec.	180	180

9. Diskussion der Ergebnisse

Die veröffentlichten Werte wurden mit einer neuen Röhre erzielt. Die Leistungsmessung erfolgte mit HP432A und vorgeschaltetem 30dB-500W-Abschwächer von Bird, in Reihe mit einem 20dB-Dämpfungsglied von HP. Die Zusammenschaltung der Abschwächer ergab eine Dämpfung von 50,1dB. Die Kabelämpfung von 1m HCF ½" wurde nicht berücksichtigt. So kann die Leistung am Ausgang der PA ca. 5% höher sein. Es wurde zum Test bis zu einer Ausgangsleistung von 900 Watt angesteuert.

Die Frage ist nun, wie hoch darf man die Röhre aussteuern, was sagen die Hersteller-Daten und wie sind diese auszulegen. Es liegen vom Hersteller keine Daten für SSB oder CW Betrieb vor, da diese Betriebsarten für UHF-Röhren nicht vorgesehen sind. Der Röhrenhersteller gibt für einen TV-Verstärker folgende Einstellung an:

Ua 3kV, Ia 650mA, Steuerleistung 7W, Ausgangsleistung 200W, bei einem geforderten Intermodulationsabstand von 54 dB. Für diese hohe Linearität ist der A Betrieb erforderlich mit einem stabilen Arbeitspunkt im linearen Bereich der Kennlinie. Für den Funkamateurbetrieb macht es keinen Sinn, eine Endstufe mit 54 dB Intermodulationsabstand zu bauen, da die Transceiver, von denen die Stufen angesteuert werden, nur Werte von 25 – 35dB aufweisen. Höhere IM-Werte sind für Sprachübertragung nicht erforderlich.

Es bietet sich für SSB der AB-Betrieb mit ca. 100 mA Ruhestrom an und für CW der C-Betrieb. Damit ist eine Ausgangsleistung von 600 W möglich ohne die Lebensdauer bedeutend zu vermindern.

Bei 500 W Ausgangsleistung können die Herstellerwerte eingehalten werden.

Mit dieser Leistung ist die Endstufe im Kontestbetrieb hinreichend mit gutem Ergebnis getestet worden.

10. Spezielle Bauteile

Federkontakt-Streifen (Finger Stock), Form S

Fa. Feuerherdt, Berlin, Tel.:030/7559094

R1 Hochlast-Draht-Widerstand mit verstellbarem

Abgriff, 100Ω 15 W, Fa. Bürklin, Best. Nr. 64E110

R6 Hochlastwiderstand 33Ω 50 W RH50, Fa. Dale

Rp 5x Sicherheits-Metallschicht-Widerstand 1.8

MΩ 1W Fa. Bürklin, Best. Nr.20E932

Lüfter: G2E108-A01-01 von Fa. EBM GmbH & Co.

Distributor: Fa. Breuell & Hilgenfeld, Hamburg

Tel.: 040/538092-20 E-Mail: info@breuell.de

11. Quellennachweis / Literature

Valvo Handbuch, Senderröhren für Nachrichtensender, 1977

12. TNX

Herzlich bedanken möchte ich mich bei Norbert DL9EAJ, der mir die gesamten Zeichenarbeiten abgenommen hat.

Amplifier for the 23cm-band with the YD 1336 Triode

Hubert Tiedmann, DJ3FI

1. Introduction

The chosen tube was originally developed for use in TV repeater stations. Here transistors have replaced tubes due to their unlimited lifetime, but VHF/UHF/SHF-radioamateurs still make their contests and EME-QSOs mostly with tube amplifiers. The limited lifetime (10,000 to 20,000 hours) is sufficient for amateur operation, even if the tube was already used.

It makes sense to use a tube amplifier on 23cms because the tube also has some advantages at these power levels.

The socket for the YD1336 is the same as for the TH 308, TH 329, YD 1333, YD1335 and YD 1337. The anode dissipation of these tubes ranges from 700 W with the TH 308 to 1900 W with the YD

1335, so the length and diameter of the anode radiator and the internal capacitances are different. The input circuit of this amplifier has a broad tuning range, so all the above mentioned types could be matched. However, if a TH308 is used, the anode resonator has to be lengthened by 1 cm.

2. Construction of the Amplifier

The amplifier is constructed as a coaxial resonator and the triode is working in a grounded grid circuit. Two coaxial resonance circuits are built into each other. Brass tubes with a diameter matching to the necessary impedance were used. The contact of the coaxial feeds at the anode and grid is made from finger stock material. The tubes at the cathode and heater ring are slotted to make better contact. The base plate and anode plate also have contacts made of finger stock.

The chimney above the anode radiator is made from a Teflon ring and a slotted aluminium ring. The anode voltage is supplied through this aluminium ring. At the upper side the chimney can be moved in order to prevent any mechanical pressure on the tube.

3. The Output Circuit

The coaxial circuit between the grid (ground) and the anode is $3/4 \lambda$. The tuning is capacitive at the first voltage maximum, located ca. 30 to 50 mm from the lower base plate.

The power output coupling is also done capacitively, on the upper side of the cavity. This realizes the impedance transformation between the load resistance of the tube (ca. 2.5 k Ω) and the impedance of the feedline. The output coupling at the upper end of the cavity gives the advantage that it is the voltage maximum for all frequencies and the resonance in $1/4 \lambda$ mode is also attenuated, resulting in no oscillation possibility around 430 MHz.

To avoid unwanted oscillations it is also important that the contacts from the finger stock are pressed tightly all around the grid ring, so no energy can leak from the output circuit to the input.

The upper side of the cavity is terminated by the terminating plate, the tube (anode) and the brass plate that is insulated with a 0.1mm sheet of Kapton. This results in a capacitor of 600 to 800pF. This plate is centered with a Teflon ring and covered with a 0.2mm Teflon sheet and a plate of brass. This makes the unit closed for any RF.

4. The Input Circuit

The tube is driven between the grid (ground) and the cathode. The coaxial input circuit transforms the input impedance from the cathode to the impedance of the feedline. The power for the heater

filaments also has to be supplied, and the cathode and the grid ring have to be cooled.

The coaxial system (grid and cathode tube) has the electrical length of 1λ . The tuning is done with an insulated and moveable ring of aluminium. When this ring is located in a voltage maximum its effect is more capacitive and makes the circuit electrically longer. When the ring is located in a current maximum, its effect is pushing away the current, lowering the L and the impedance and making the circuit electrically shorter. This results in a resonator with low losses and a broad tuning range, making it possible to match all the above mentioned tubes.

At the open end of the coaxial system a capacitive input coupling is realized with a round disk and an N-connector. The distance from the disk to the cathode tube determines the ratio of the voltage and thereby the transformation from the impedance of the tube to the impedance of the feedline (50 Ω). The required condition is that the main circuit is tuned to resonance to have a voltage maximum at the input point.

Inside the cathode tube there is another tube for the filament contact that is insulated with Teflon washers. The filament voltage is supplied to the cathode tube and the filament tube through a choke with feedthrough capacitors.

The air for cooling the cathode is supplied to the inner tube and leaves the cathode area through holes at the side of the socket. Then the air goes through the grid tube and leaves the unit through holes in the base plate.

5. Regulation of the Cathode Voltage

If the tube is to work at an anode voltage of 3 kV and a bias idle of 80 mA, we need a negative bias voltage of around 38 V at the grid. Since the grid is DC-grounded, the cathode has to be made 38 V more positive. In order to prevent any flash-over in the tube there should be always a small current of 3 to 10 mA. This current is determined by the resistor R1.

If the TX relay is switched on there is a current of around 70mA. This current is determined by R2 and R3. If a higher current is necessary the tap at R2 has to be moved towards the cathode until the Zener voltage of D1 is reached and drives the transistor. 80mA and more can be achieved. The resistor R6 takes a part of the load of the regulator and limits the current through the tube in case of a emitter-collector short circuit after a flash-over.

6. Switching On and Tuning

If the blower, heater supply and high voltage are ready for use, the primary adjustments can be made.

The coupling of the input circuit is adjusted to 1mm distance to the cathode tube. The output

coupling is adjusted to 2mm distance to the grid tube. The R2 divider is set to the middle. Into the anode voltage feedline a resistor of 1 k Ω is inserted. Now the blower and the tube's heater can be switched on.

As all these tubes are older than 10 years it is necessary to heat them for 30 minutes before supplying the anode voltage via a high load resistor. This may not be possible for a long time due to the heat in the resistor, so the process may have to be stopped and repeated again and again. In this way, short-circuits can be determined and voltages and currents can be checked.

If the output of the amplifier is terminated with a dummy load and a power meter, the full anode voltage can then be supplied. For the first matching, 5 to 10 Watts are supplied to the input. Now the tuning ring is moved until the anode current is at a maximum. Then the anode circuit can be tuned to maximum power output. The reflection at the input should be measured with a directional coupler and tuned by changing the input coupling to give a return loss of at least 16 to 20 dB (SWR <1.4). Due to the grounded grid circuit the output and input impedances are in series with the tube. This means that the output coupling and tuning of the anode also has some influence on the input matching. So after optimising the output the input has to be checked and retuned again.

A hint for optimising the output power: if the tuning ring is moved closer to the grid tube (tighter coupling) the working resistance is getting lower. The amplification is going down and the voltage at the anode is going up, resulting in a lower grid current. The opposite happens with a loose coupling: we have higher amplification and have to reduce the driving power, because the grid current gets too high. If the coupling is too loose the working Q gets too high, and due to the small bandwidth of the circuit it becomes sensitive to detuning due to thermal changes from the tube. For a given input power the tuning of the coupling always allows to find the highest output power.

7. Measured Results

See table in the German text.

8. Tube Data (See table)

9. Discussion of the results

The values were achieved with a new tube. The power was measured with a HP432A power meter and a 30dB/500W attenuator from Bird in series with a 20dB attenuator from HP. Both attenuators gave 50.1 dB. The cable loss from 1m HCF 1/2" was not considered, so the power at the output of the PA may be some 5% higher. For test purposes the amplifier was driven up to 900 watts of output power.

Parameter, Symbol		Unit	YD 1334/1336	YD 1335
Filament Voltage	Uf	V	6 – 6,3	6 – 6,3
Filament Current	If	A	4,8 -5,8	4,8 -5,8
Capacity Anode-Grid	Cag	pF	6,8 – 8	6,8 – 8
Capacity Grid-Cath.	Cgk	pF	20 – 30	20 – 30
Capacity Anode-Cath.	Cak	pF	0,09 – 0,18	0,09 – 0,18
Mutual Con- ductance	s	mA/ V	ca. 70	ca. 70
Anode Voltage	Ua	V	max. 3500	max. 3500
Anode Diss. Loss	Qa	W	1800	1900
Grid Current	Ig	mA	5	5
Cathode Current	Ik	mA	max. 700	max. 800
Grid Voltage	Ug	V	max. – 200	max. - 200
Heat. Time	Th	sec.	180	180

The question is what the maximum input power for the tube should be, and how to interpret the manufacturer's data. Of course, there are no data available for SSB or CW, because these tubes were made for other applications. The manufacturer provides this for a TV transmitter: Ua 3kV, Ia 650mA, Input power 7W, Output power 200W, for a necessary intermodulation value (IM) of 54dB. For this high linearity the tube must work in class A mode, but for radio amateurs this is not required because the driving transceivers only have IM of 25 to 35 dB. So for SSB, class AB operation with around 100mA bias current should be a solution. For CW only, class C mode is OK of course. With this an output power of 600 watts is possible without any big decrease of the tube's lifetime.

At 500 watts output all specifications of the manufacturer are fulfilled. With this power the PA was tested under contest conditions with good results.

10. Special parts

Finger Stock, Form S: Fa. Feuerherdt, Berlin, Phone: +49 30 7559094.

R1 High power wirewound resistor (with adjustable tap) 100 Ω 15 W: Fa. Bürklin, Munich, Order Nr. 64E110.

R6 High power resistor 33 Ω 50 W RH50, Fa. Dale. Rp 5x Safety-metalfilm-resistor 1.8 M Ω 1W: Fa. Bürklin, Order Nr. 20E932.

Blower: G2E108-A01-01 Fa. EBM GmbH & Co. Distributor: Fa. Breuell & Hilgenfeld, Hamburg, Tel.: +49 40538092-20 info@breuell.de

Teil	Anz.	Einh.	Benennung	Verstärker	Z. Nr.
1	1		Abdeckung für Luftschacht	Aluminium	3
2	1		Haltering für Teflonschacht	Aluminium	17
3	1		Teflonschacht	PTFE	20
4	1		Luftschacht	Aluminium	12
5	1		Röhrenabschluss	Aluminium	18
6	1		H.V. Platte für Anode	Messing	7
7	1		Abdeckung HV Platte	Messing	5
8	1		Zentrierung HV Platte	PTFE	6
9	1		Abschluss Anodentopf	Messing	8
10	1		Kupferwinkelstück	Kupfer	12
11	1		Kathodenrohr	Messing	14
12	1		Heizungsrohr	Messing	22
13	1		Zentrierung Heizungsrohr	PTFE	24
14	1		Isolierung Kathodenschieber	PTFE	10
15	1		Kathodenschieber	Aluminium	10
16	1		Anodentopf	Messing	13
17	1		Haltering Kathodenrohr	Messing	9
18	1		Zentrierung Kathodenrohr	PTFE	9
19	1		Zentrierung Heizungsrohr	PTFE	24
20	1		Slit für Kathodenschieber	Hartgewebe	11
21	1		Bodenplatte Anodentopf	Messing	1
22	1		Gitterrohr	Messing	25
23	1		Luftschacht Kathode	PTFE	24
24	1		Bodenplatte Gitterrohr	Messing	19
25	1		Gewindebuchse	Messing	2
26	1		Koppelplatte Input	Messing	2
27	1		N - Einbaubuchse		2
28	4		Halbboizen	Messing	2

Stückliste 1

Linear Amplifier 1296 MHz

DJ 3 FI

Ver. Nr. 100

Rev. 100

Rev. 100

Parts List

Teil	Anz.	Einh.	Benennung	Werkstoff	Z. NR.
29	1		Radiallüfter Fa EBM Typ G2E 108-AA-01	—	—
30	1		Zentrierung Kathodenrohr	PTFE	11
31	1		Kupferrohr 18 x 1 55 lang Kathodenbeulung	Kupfer	19
32	1		Kupferrohrwinkel Kathodenbeulung	Kupfer	19
33	1		Kupferrohrwinkel Kathodenbeulung	Kupfer	19
34	1		Kupferrohr 18 x 1 30 lang Kathodenbeulung	Kupfer	19
35	1		Abschirmstempel	Messing	23
36	1		Schlauch 18 x 3 195 lang	PVC	23
37	1		Abschirm C	Messing	23
38	2		Gewindeplatte	Messing	26
39	1		Feststeking	Messing	26
40	1		Buche für Auskopplung	Messing	26
41	1		Auskoppelseinstempel mit 7/16 Buchse		27
42	1		Kappe für Abschirmung	Aluminium	28
43	1		Stelling	V2A	28
44	1		Gewindespindel	V2A	28
45	1		Buchse für Durchführungs C	Messing	28
46	1		Folie für Anodenblock-C oder H V Platte	Teflon	16
47	1		Folie für Anodenblock-C unter H V Platte	Kapton/Teflon	16
48	1		Röhre YD 1335/1336		4
49	1		Isolation	Teflon	
50	1		Mutter aufgelötet	Messing	
51					
52					
53					
54					
55					
56					

Stückliste 2

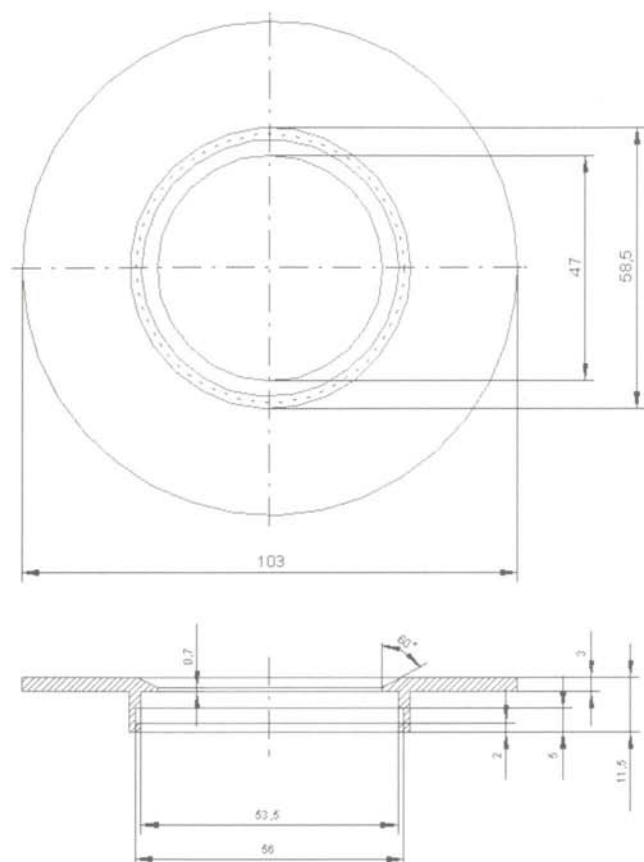
DJ 3 FI

Linear Amplifier 1296 MHz

Rev. 2

Part 8

Part 7

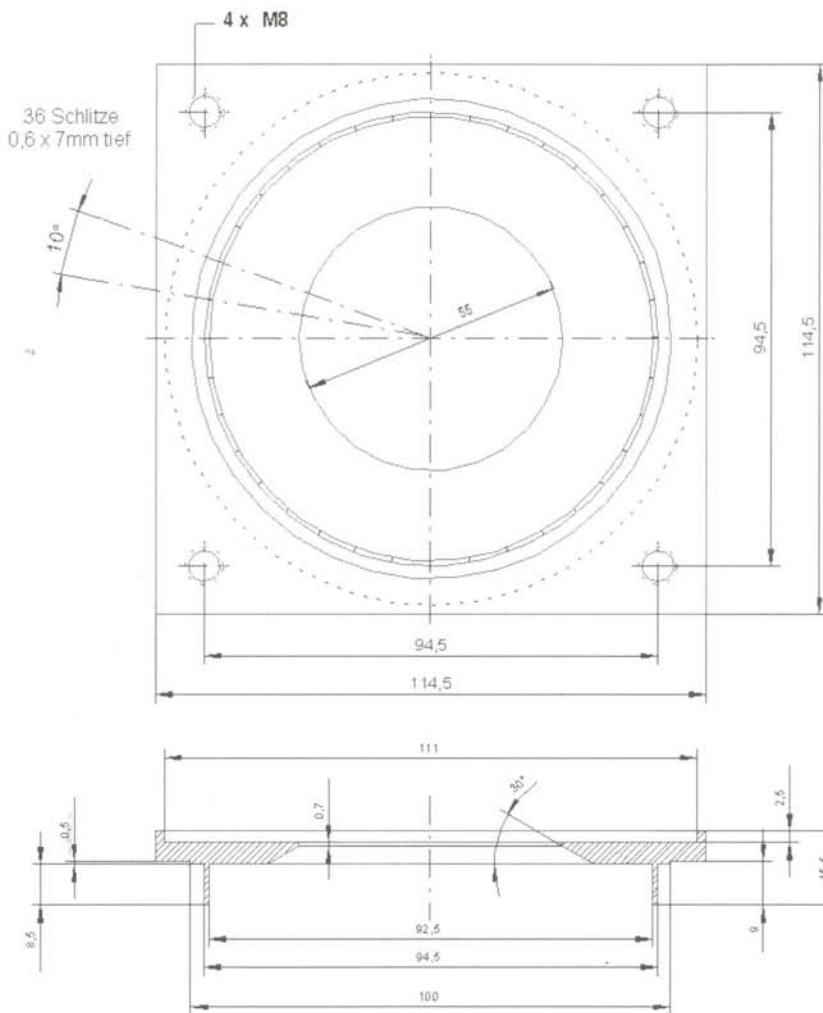


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							Werkstoff	Messing
							H.V. Platte	
							Linear Amplifier 1296 MHz	
							Blatt 1	
							1 Blatt	
Zust.	Änderung	Datum	Name	Urspr.	Ers. für		Ers. durch	

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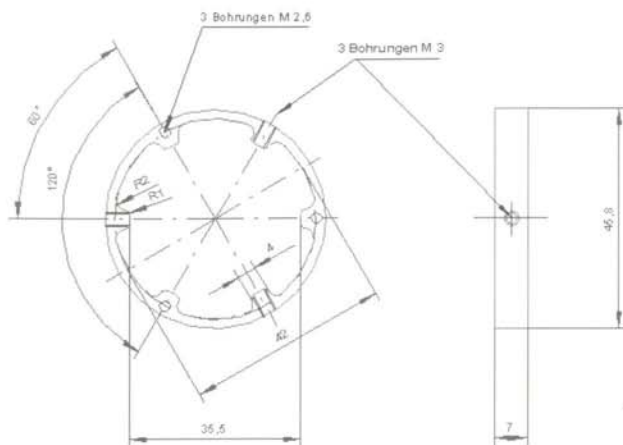
Part 6

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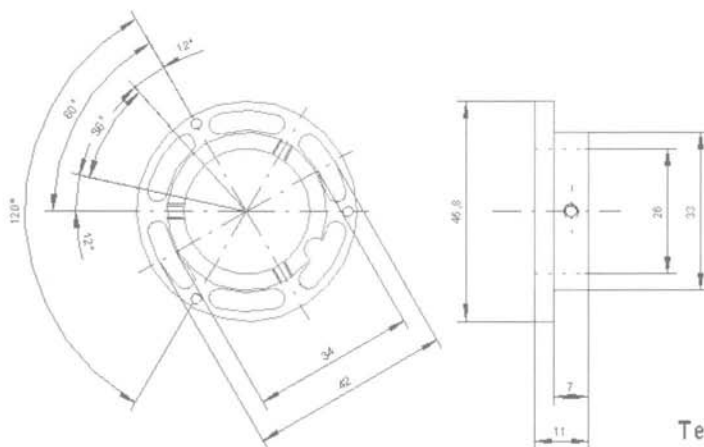


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							Werkstoff	
				Datum	Name	Abschluss Anodentopf		
				Bearb.	Couhorn			
				Gepr.	Tiedmann			
						Linear Amplifier 1296 MHZ		Blatt 1
								1 Blätter
Zust.	Änderung	Datum	Name	Urspr.	Ers. für	Ers. durch		

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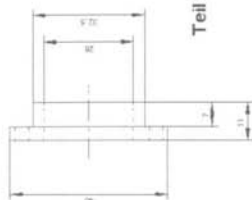
Teil 17



Teil 18

Teil 17 18				Allgemeintoleranzen DIN 7168	9	Maßstab 1:1	Gewicht
						Werkstoff	17 Messing 18 Teflon
				Datum	Name	Halterung Kathodenrohr	
				Bearb.	Couhorn		
				Gepr.	Tiedmann		
						Linear Amplifier 1296 MHZ	
Zust.	Änderung	Datum	Name	Urspr.	Erz. für:	Erz. durch:	Blatt 1 1 Blatt

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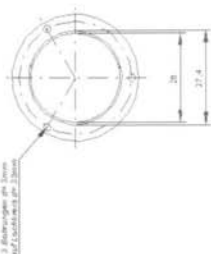
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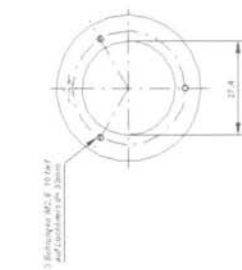
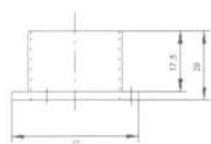
Teil 20

Teil 20, 30		11		Materialnr		Gewicht	
Abgemessen Dimensionen DN 7 168		Name		PTFE			
BohrD 82/800		Datum					
GezD 82/800		Gezeichnet					
		Technisch					
		Teil 30 Zentrierung Kathodenrohr					
		Teil 20 Stift für Kathodenschieber					
DJ 3 FI				Linear Amplifier 1296 MHz		1. Seite 1. Blatt	

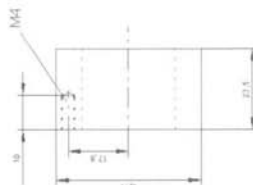
Part 30 and 20



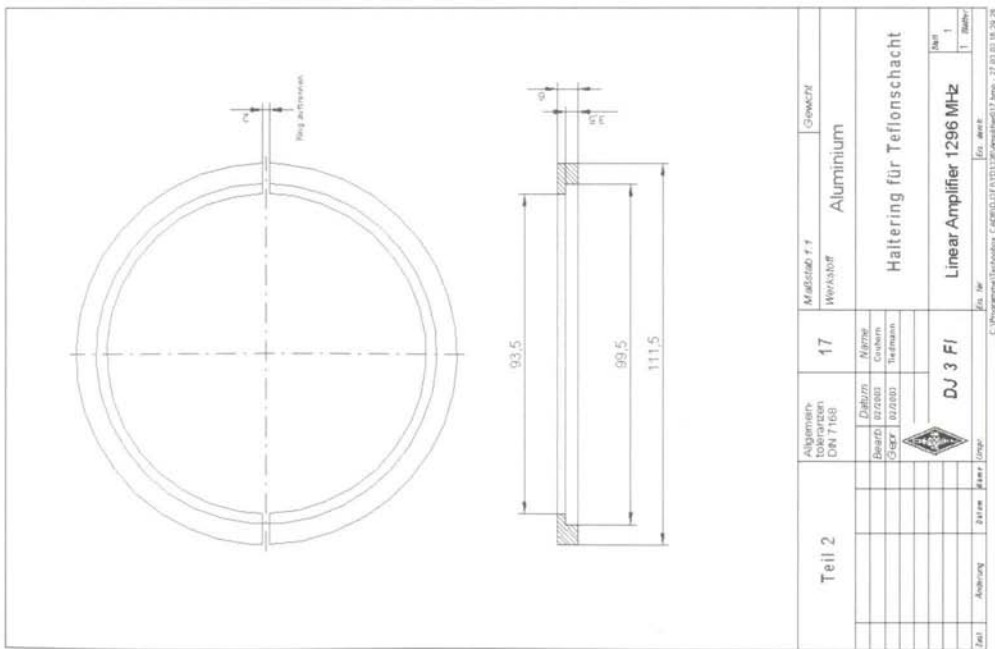
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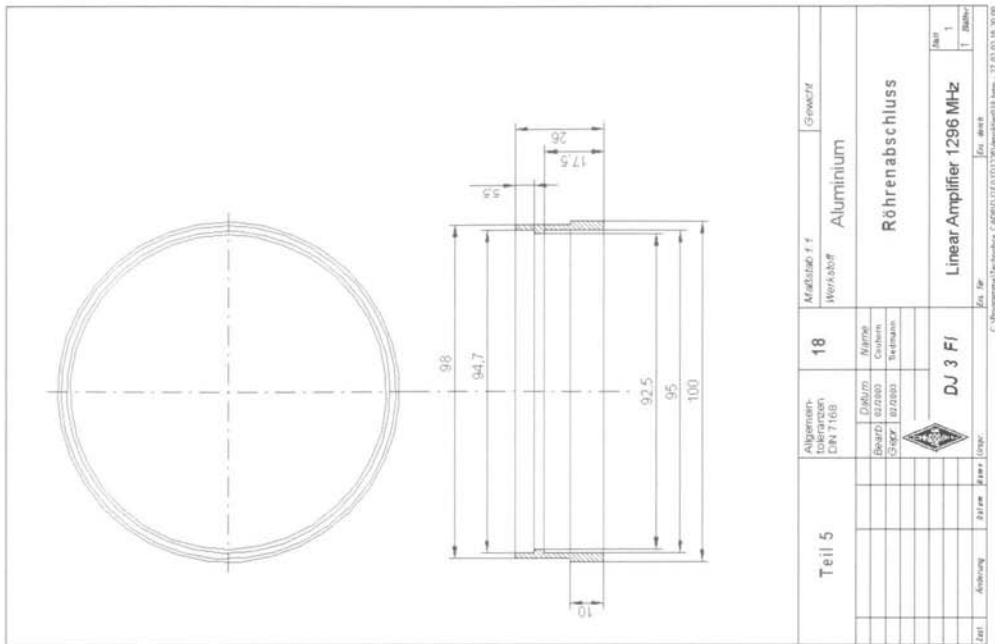
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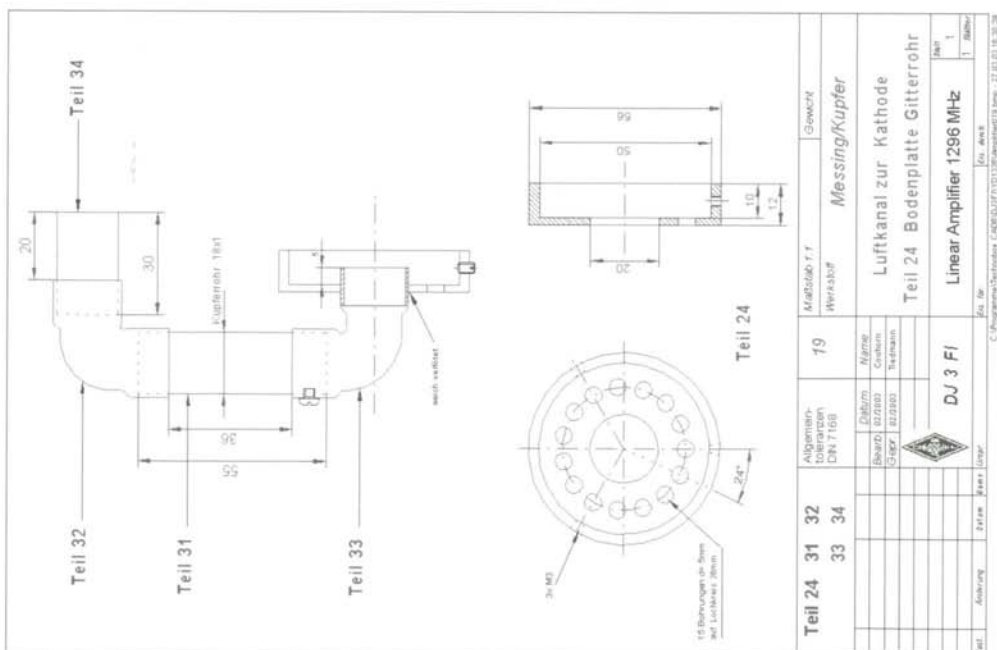
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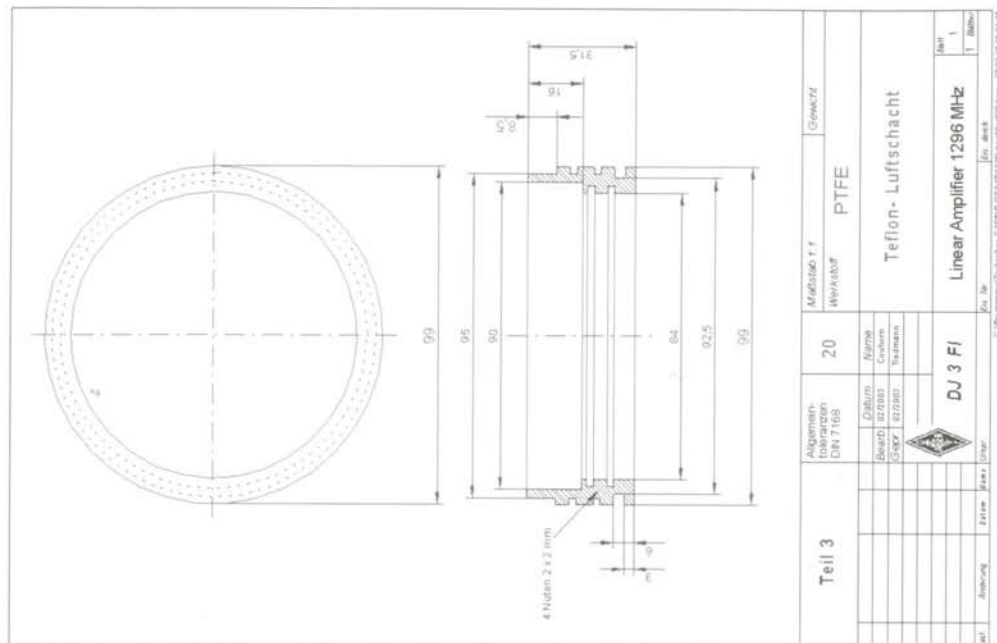
Part 2



Part 5

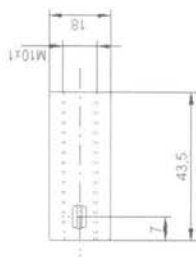


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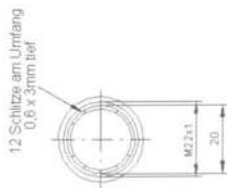


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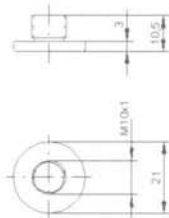
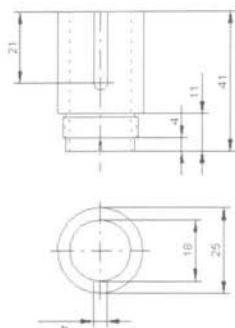
Teil 35



Mutanten aus M4 Zylinderoptische auf einem trockenen



Teil 36



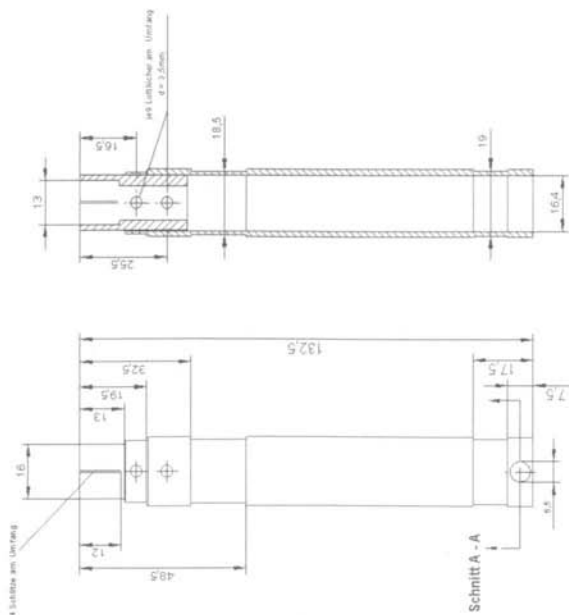
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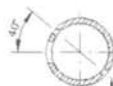
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Schnitt A - A

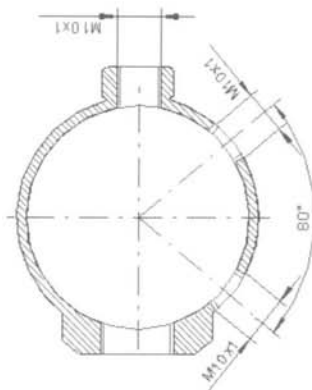
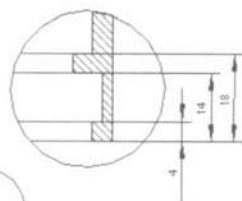
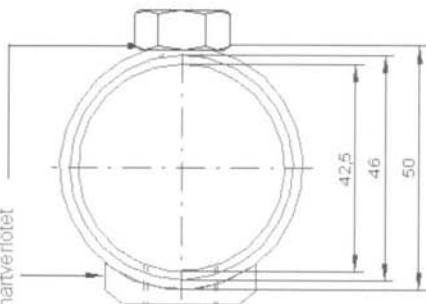
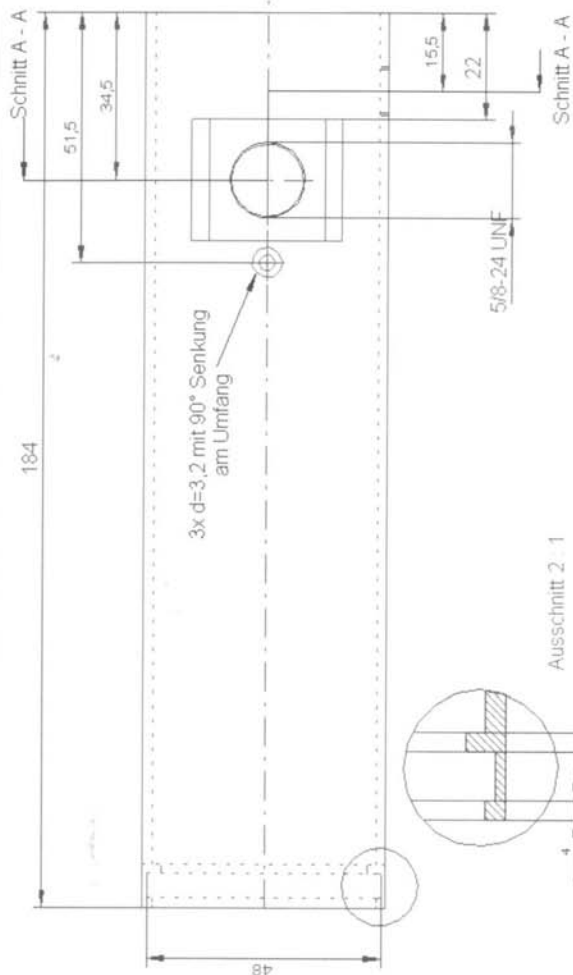


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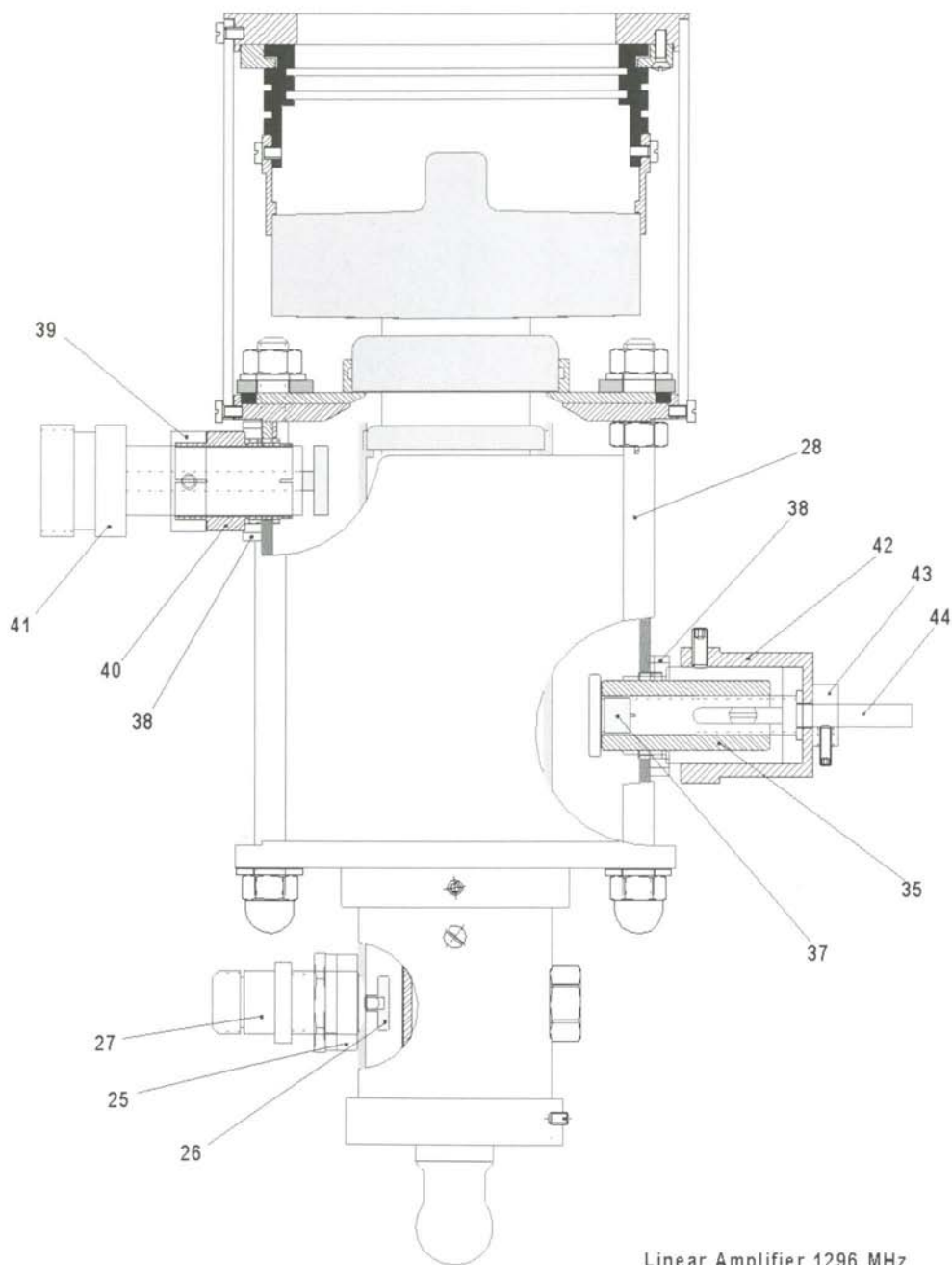
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Part 35, 36, 37



Teil 22		Allgemeintoleranzen DIN 7168	25	Maßstab 1:1 Werkstoff Messing	Gewicht
			Datum	Name	Gitterrohr Linear Amplifier 1296 MHz
		Bearb. Gepr.	02/2003	Couborn Tiedmann	
					
			DJ 3 FI		Blatt 1 1 Blätter

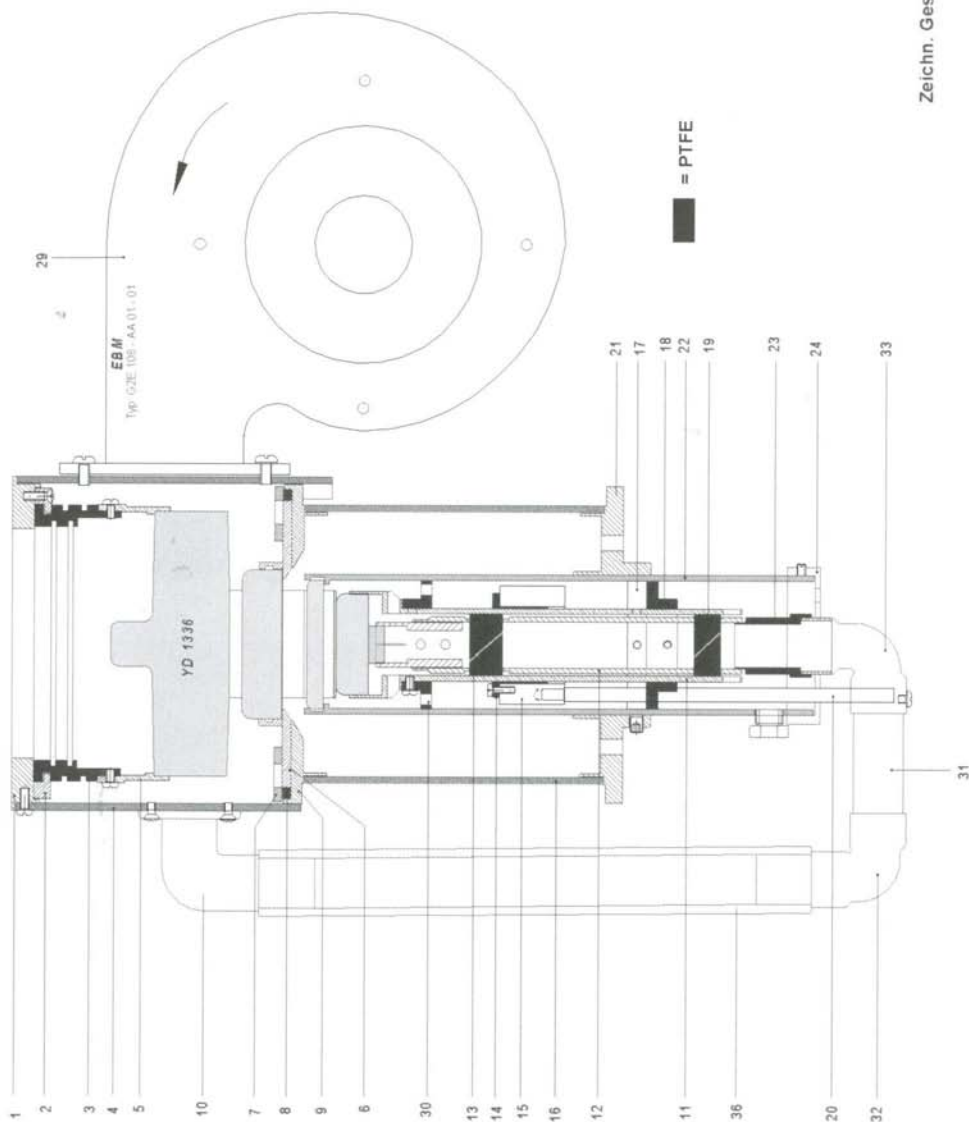
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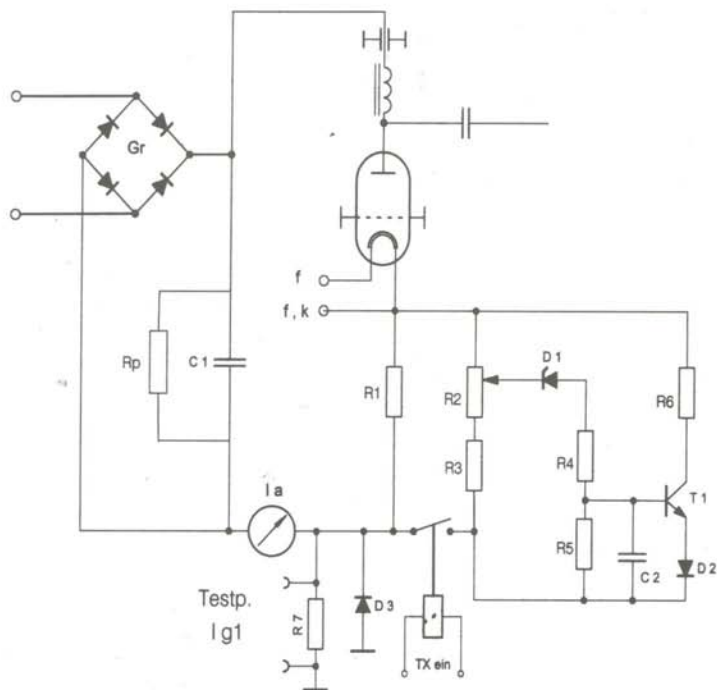
Linear Amplifier 1296 MHz

Zeichn. Ges. Nr.7

PA Komplett, Ansicht auf Koppelstempel



23cm PA DJ3FI, Querschnitt mit Nummerierung der Teile



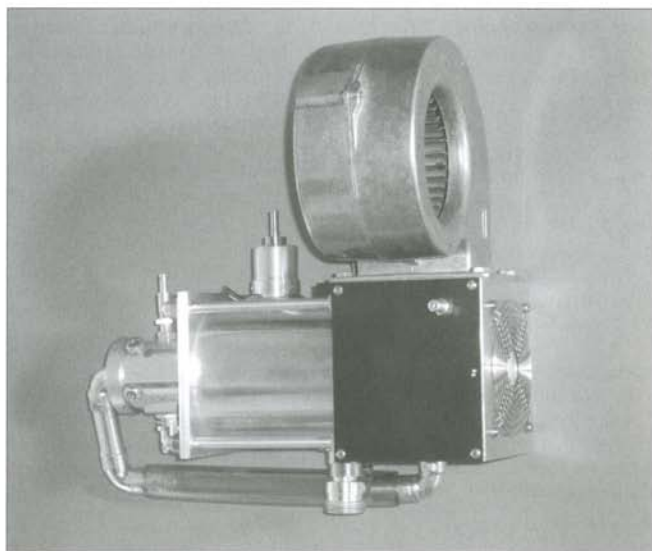
$R1 = 4,7 \text{ k}\Omega$
 $R2 = 100 \Omega$
 $R3 = 470 \Omega$
 $R4 = 47 \Omega$
 $R5 = 1,5 \text{ k}\Omega$
 $R6 = 33 \Omega$
 $R7 = 10 \Omega$
 $Rp = 5 \times 1,8 \text{ M}\Omega$
 $D1 = \text{ZX } 33$
 $D2 = \text{BY } 255$
 $D3 = \text{BY } 255$
 $C1 = 8 - 20 \mu\text{F}$
 $C2 = 10 \text{ nF}$
 $T1 = \text{BUX } 48$

Stromversorgung für Verstärker mit YD1336

dj3fi

Circuit Diagram

Complete
PA



Using MRF284 a Motorola⁽¹⁾ LDMOSFET on our bands

Franck Rousseau, F4CIB⁽²⁾

1. Introduction

As I request more technical articles for our favorite magazine, here is my 2 cents contribution!

I'd read with interest F5EFD⁽³⁾⁽⁴⁾, Maurice Niquel, and F1ANH⁽⁵⁾, Jean-Pierre Lecarpentier, articles about conversion of Motorola devices on 23 and 13cm but wasn't able to get such devices.

These devices were really a good opportunity to replace expensive power bricks but also the well-known 2C39 cavities amplifiers, so I tried to find a good replacement to my M57762 power module.

Being lucky to find a MRF284 device and its demonstration board, I decided to try to make it work on 1296MHz and to share this experience with you...

Transistor characteristics really sound interesting for our ham applications: Broadband (1GHz-2.6GHz), class A & AB specified.

Being QRV only on 23cm no tests have been done on 13cm, but it should work too.

2. Transistors Characteristics

Performances quick overview extracted from product datasheet⁽⁶⁾:

Specified Two-Tone Performance @ 2000 MHz, 26 Volts

- Output Power = 30 Watts PEP
- Power Gain = 9 dB
- Efficiency = 30%
- Intermodulation Distortion = -29 dBc

Typical Single-Tone Performance at 2000 MHz, 26 Volts

- Output Power = 30 Watts CW
- Power Gain = 9.5 dB
- Efficiency = 45%

Capable of Handling 10:1 VSWR, @ 26 Vdc, 2000 MHz, 30 Watts CW Output Power

3. Evaluation schematic

As I got a demonstration board with devices, the schematic is only an adaptation from the original one with a bill of materials update. (Sorry some comments should remain in French as the original paper was written in that language!)

4. PCB layout and wiring

See Figures on next page.

5. Tuning

All measurements have been done on professional calibrated equipment:

- Network Analyzer HP8753: Forward Gain S21 and Input Return Loss S11.
- Power Meter HP438A: RF Output Power.
- DC Power Supply HP6632: DC Sources and Measures.

Single DC supply is required, so first step is to set the bias current with the help of the adjustable resistor. Set the bias to 200 mA and forget it!

Without any capacitors, the amplifier runs as an attenuator, all input power goes back to the source, so it helps to insert an isolator!

Let's first add some stuff in the input area...

Best matching network is done with 3.9 pF closed to DC block cap (but not only DC block, it's part of input network!) and 1.5 pF on the 30 Ohm line.

The 3.9pF capacitor's position is the most critical, 1.5 pF along the 30 Ohm line helps to center the tuned frequency at 1296MHz.

After that step the gain is already around 9-10 dB, though Input Return Loss (IRL) is close to -13 to -14 dB.

In the same way, next thing to set is the output network.

2.2 pF soldered at the 30 Ohm line center and 3.9 pF as close as possible of DC block capacitor (not only once again, keep that in mind...) will push the gain up to 14dB, while IRL remains at -14dB.

6. Measurements results

1dB compressed power, P_{1dB} . (Output power level while gain drop of 1 dB) has been measured at 45.5 dBm (35 W) in accordance with the device specifications.

Under following stress conditions: VCC = 28 V, Input Power = 31.6dBm, the output power level has been measured at 45.2 dBm, while the DC current rose to 2.41 A.

Computing Power Added Efficiency (PAE) gives 47% which exceeds the specifications, but that's not a surprise as the device is able to work up to 2 GHz. The lower the frequency, the more nervous is the transistor! (See Fig. on next page!)

7. Conclusion

As expected, as MRF282 and MRF286, the MRF284 works without any problems on 1296 MHz. Its output power twice as a power brick and its gain should be a good alternative to these such

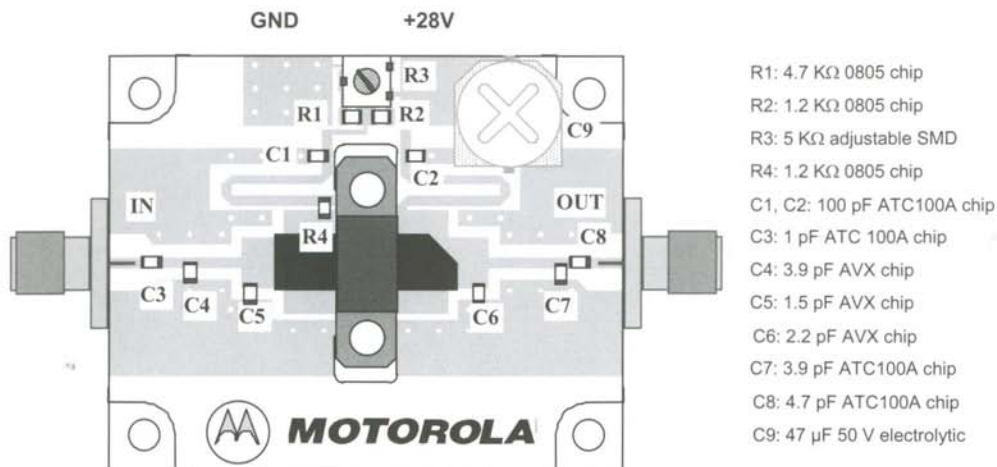


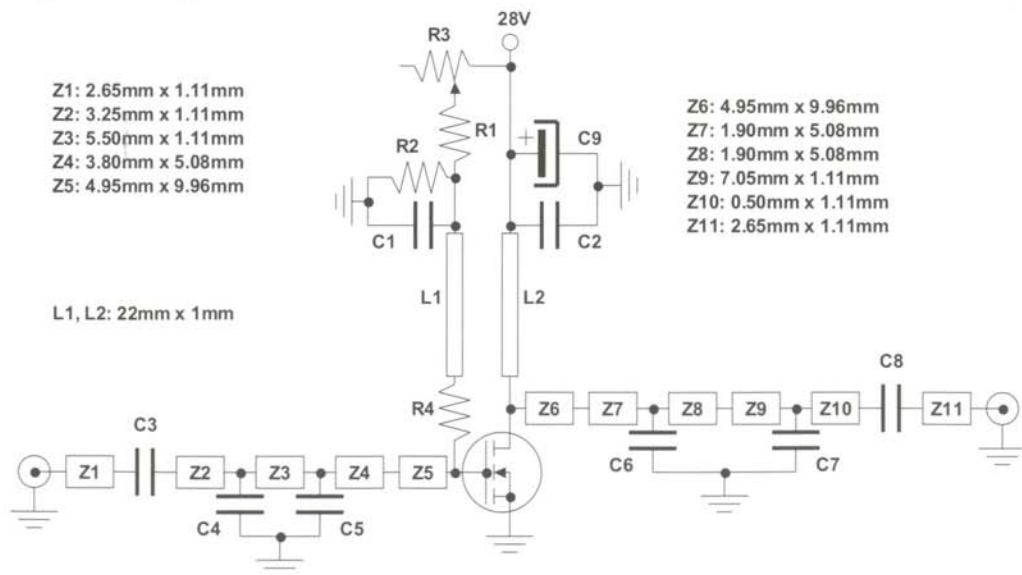
Fig. PCB Layout

expensive power modules. Now, the trick is to find a good and cheap source. These transistors are used in cellular base stations... Should be a clue for some of you!

I think Power Added Efficiency optimization could be pursued, but today a 2.5A DC/DC converter seems to be really easy to build.

Just read the following article....

Fig. Circuit Diagram



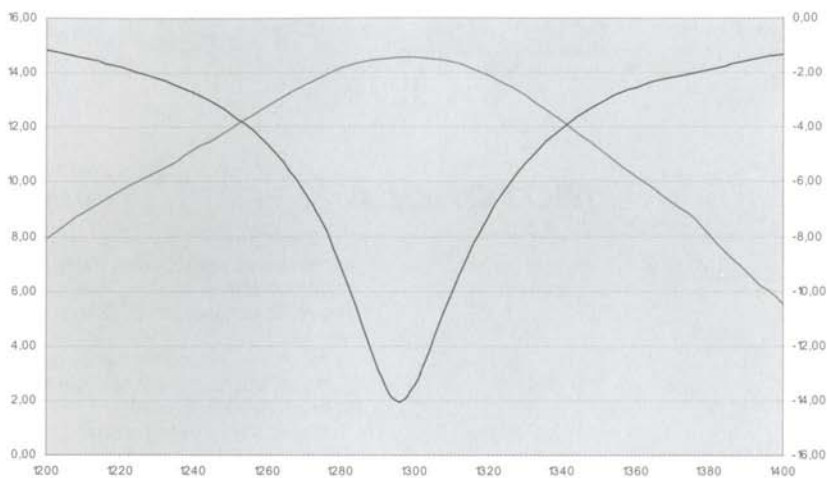
Substrat ROGERS RO4003 H=0.508mm T=35 μ m ϵ_r =3.48

8. References

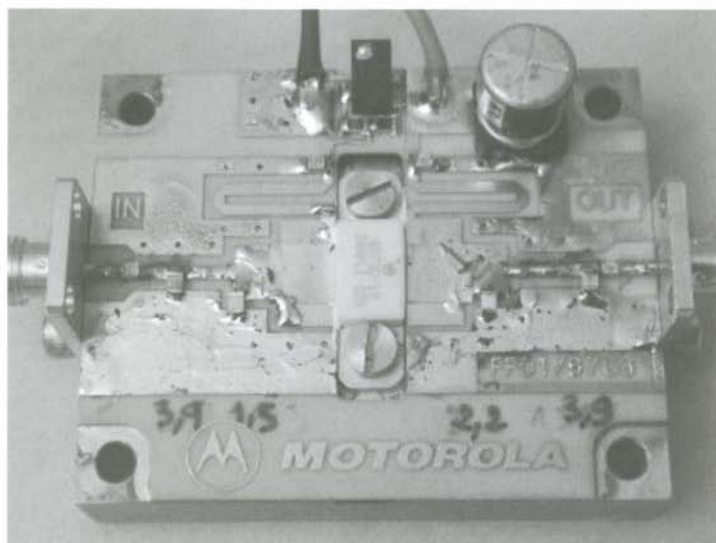
1. <http://www.motorola.com/>
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AppT 31 – 31400 Toulouse – France – email:
f4cib@ref-union.org
3. F5EFD – Hyper n°56 – Février 2001 –
Amplificateur 4W 13 cm à transistor LDMOS
4. F5EFD – Hyper n°65 – Novembre 2001 –
Amplificateur 70 W sur 1296 MHz
5. F1ANH – Hyper n°73 – Juillet-Août 2002 – Ampli
2300 MHz à MRF282; 12 ou 24 VDC
6. [http://e-
www.motorola.com/brdata/PDFDB/docs/MRF284.pdf](http://e-www.motorola.com/brdata/PDFDB/docs/MRF284.pdf)
See also Microwave Parts in this issue !

Gain S21

RL / S11



Frequency



Einsatz des Motorola MRF284 LDMOSFET im Amateurfunk

Franck Rousseau, F4CIB

Einleitung

Mit Interesse habe ich die Artikel von F5EFD⁽³⁾⁽⁴⁾, Maurice Niquel und F1ANH⁽⁵⁾, Jean-Pierre Lecarpentier, über den Einsatz von Motorola-Transistoren auf dem 23cm und 13cm-Band gelesen. Leider war es mir nicht möglich, die dort verwendeten Typen zu bekommen.

Diese Transistoren wären eine gute Möglichkeit, die teuren Power-Module und auch die bekannten 2C39-PAs zu ersetzen.

Deshalb habe ich versucht, einen Ersatz für mein M57762 PA-Modul zu finden.

Ich konnte mit etwas Glück einen MRF284 bekommen, inkl. zugehörigem Demo-Board.

Ich entschied mich, einen Aufbau für 1296 MHz zu realisieren und meine Erfahrungen hier mitzuteilen. Die angegebenen Eigenschaften dieses Transistors klingen für Amateurfunk-Anwendungen sehr interessant: Breitband (1 bis 2.6 GHz), Klasse A und AB. Da ich nur auf 23cm grv bin, habe ich keine Tests auf 13cm durchgeführt. Das ganze sollte aber dort auch arbeiten.

Eigenschaften des Transistors

Hier ein kurzer Überblick der Eigenschaften des Transistors, die dem Datenblatt entnommen sind⁽⁶⁾:

Zweiton-Performance bei 2000 MHz / 26 Volt

- Output Power = 30 Watts PEP
- Gewinn = 9 dB
- Wirkungsgrad = 30%
- Intermodulation Distortion = -29 dBc

Typ. Eintont-Performance bei 2000 MHz / 26 Volt:

- Output Power = 30 Watts CW
- Gewinn = 9.5 dB
- Wirkungsgrad = 45%

VSWR von 10:1 ist möglich, bei 26VDC, 2000 MHz, 30 Watt CW Ausgangsleistung

Entwicklungs-Board / Schaltung

Da ich ein Entwicklungsboard mit Transistor zur Verfügung hatte, stellt die Schaltung nur eine Abwandlung des Originals mit einer neuen Stückliste dar.

PCB Layout und Verdrahtung

Siehe Abbildung vorige Seite.

Abstimmung

Alle Messungen wurden mit folgendem professionellen und geeichten Equipment durchgeführt:

- Network Analyzer HP8753: Forward Gain S21 and Input Return Loss S11.
- Power Meter HP438A: RF Output Power.
- DC Power Supply HP6632: DC Sources and Measures.

Nur eine einfache DC-Spannung wird benötigt. Als erstes wird der Ruhestrom mit dem Einstellwiderstand auf 200mA eingestellt.

Ohne irgendwelche Kondensatoren verhält sich der Verstärker wie ein Abschwächer, sämtliche Eingangsleistung geht zur Quelle zurück. Deshalb hilft es, einen Isolator zu verwenden.

Zuerst wird im Eingangsbereich bestückt.

Die beste Anpassung wird mit 3.9pF nach Masse (DC Block, aber auch Teil des Eingangsnetzwerks!) und 1.5 pF auf der 30 Ohm Stripline erreicht.

Die Lage des 3.9pF Kondensators ist sehr kritisch. Der 1.5 pF Kondensator auf der 30 Ohm Stripline hilft die eingestellte Frequenz bei 1296 MHz zu halten. Danach beträgt die Verstärkung bereits etwa 9 bis 10dB. Der Input Return Loss (IRL) liegt nahe -13-14 dB. Genauso wird als nächstes das Ausgangsnetzwerk bestückt: 2.2pF auf der 30 Ohm-Steifenleitung und 3.9 pF möglichst direkt nach Masse als DC-Block (aber auch Teil der Anpassung!). Nun geht die Verstärkung hoch auf 14 dB. Der IRL bleibt bei -14dB.

Meßergebnisse

Die 1dB kompr. Ausgangsleistung wurde in Übereinstimmung mit den Spezifikationen mit 45.5dB, (35 W) gemessen. Unter Stress-Bedingungen (28V, Input 31.6 dBm) wurde ein Output von 45.2 dBm bei 2.41A DC gemessen.

Die Power Added Efficiency (PAE) ergibt 47%, was die Spezifikation übertrifft. Dies ist aber keine Überraschung, das der Transistor bis 2 GHz arbeiten kann. Je niedriger die QRG, desto „nervöser“ der Transistor! (Siehe Abb.)

Zusammenfassung

Wie erwartet arbeitet der MRF284 – genauso wie MRF282 und MRF286 – ohne Problem auf 1296 MHz. Der Output beträgt das Doppelte von dem der teuren Power-Module. Nun muss versucht werden, Quellen für diese Transistoren aufzutun. Sie werden in Basisstationen im Mobilfunktelefonnetz eingesetzt... Eine Optimierung der Power Added Efficiency könnte weiter verfolgt werden.

Aber ein 2.5A DC/DC-Konverter scheint zur Zeit recht einfach realisierbar.

(Siehe nachfolgender Artikel!)

28 V Switched Power Supply for LDMOS-PA's

Achim Vollhardt, DH2VA / HB9DUN

In recent months, LDMOS PAs for 23 cm became more and more popular. While these new devices have better efficiency (50% vs. 35 %) than the well known Mitsubishi Modules, the supply voltage of 28 V is new to most hams. Although higher voltages result in lower currents and therefore less problems with contact resistances, most ham radio stations are based on 12 V main supply voltage. This is in particular true for portable stations, as they rely on lead acid batteries as main power source. Together with these new PA units, the need for a simple, most efficient power supply appeared.

Due to the ongoing development on switched power supply circuits, many devices and applications are available. Even for newcomers, which are familiar only with standard linear regulators like the 78XX/79XX family, switched power supplies are becoming a real alternative when dealing with high power applications. As a target example, a power supply for the G3WDG 50W PA will be presented.

Component selection

The power requirements of the G3WDG 23cm PA are 28 V @ 4.5A max. (126W). To generate this voltage from a 12 V source, we are going to use the so called boost topology, which is sketched in figure 3.

Basic components of a switching regulator are the inductor, the diode and the switch (which is essentially a MOSFET). When arranged in this boost topology, the input voltage is increased ('boosted') to a higher value. For a more in depth discussion of a boost converter see [1]. When looking on the schematics, we can start identifying the first possible sources of power losses, which will become important later:

- the ohmic resistance of the inductor
- the ohmic resistance of the switch when turned on
- the forward voltage of the diode

To complete the boost topology to a full switching regulator, we need a feedback circuit, which is controlling the regulator's properties according to load demands. While this had to be built with discrete components in former years, the industry is providing us today with several highly integrated switching power supply controllers, which include overcurrent sensing, soft start features and a sense feedback for adjusting the output voltage. I chose the LT1680 from Linear Technologies for this application, as this device is specifically made for a boost topology based on a 12 V input. In addition, this controller is able to handle the high gate capacitances which come with the high power MOSFET devices needed.

Beside the controller IC, the inductor is the central device in such a converter. It must be able to provide the needed inductance for the switching converter operation even when biased with a high current. When the rated current is exceeded, the

core won't be able to handle any more magnetic flux, the inductor saturates. This can result sometimes in catastrophic failure modes and is reason for many people not to build switched regulators. Therefore, one has to be very careful when selecting the inductor. When in doubt, take the larger one! In this case, I chose the DMT-3 inductor from Coilcraft, which happens to be the largest inductor in their catalogue. According to its datasheet, it is able to sustain 35 uH of inductance when biased with 12 A. This will be necessary, as our converter will sink above 10 A from its power supply when operating at 100% efficiency and even more, if we lose some percents, which we certainly will..

The next component to chose is the MOSFET switch. Basic parameters like maximum VDS and IDS should be well above 28 V and 12 A respectively. To minimize ohmic losses during the MOSFET's on-time, look for lowest possible RDSon. This should be in the order of 10 mOhm. In addition, there is another loss mechanism not mentioned before: the dynamic losses of a MOSFET switch. An ideal mechanical switch has either R=0 or R=∞ with an instant transition between both states. In both cases, no power is lost in the switch. A real MOSFET switch however has inherent rise- and falltimes, where the resistance has a finite value and therefore the MOSFET dissipates power. This power can easily be calculated:

$$P = 0,5 \cdot V_{out} \cdot (I_{out} / (1 - DC)) \cdot (T_r + T_f) \cdot f$$

with 1-DC = Vin/Vout (DC: Duty Cycle) and switching frequency f.

It can be seen, that the rise-/falltimes can significantly contribute to converter losses. To minimize these, I chose the FQP85N06 as MOSFET. With a RDSon of 10mOhm and rise-/falltimes around 200ns, this is a very good device for our needs. Of course this is only true, if the gate can be charged and discharged faster than this. Even with the cited LT1680 controller, this can be a problem, as we deal here with times in the order of 200 ns and gate capacitances of 5 nF and more. A simple calculation shows us, that 12 V * 5 nF = 60 nC, 60 nC / 200 ns = 0.3 A. This means, that our gate driving circuits has to be able to drive a current pulse of more than 0.3 A for a risetime of 200 ns and less! To further enhance the overall efficiency, one can now use so-called MOSFET drivers specifically designed to perform such tasks. The used device from Texas Instruments (UCC37322) can drive/sink up to 9 A to a gate.. fair enough. If you don't want to use it, simply connect the controller-IC output to the MOSFET gate directly.

Last but not least, we arrive at the diode. When using now a diode at a forward current of 10 A and a forward voltage drop of 0.7 V, we already let the diode produce 7 W of heat, which is 7% in an 100 W application. To minimize this losses, we can now choose low forward voltage schottky diodes, which are capable of 10 A+. In addition, paralleling diodes will give us some more improvements, as the forward voltage is not constant, but decreases with lower forward current (I assume that every reader knows the basic I/U curve of a diode). I used the MBR20100CT dual schottky diode and used two devices in parallel, which gives me four diodes in total. One also dis-

tributes the heat mechanically a little bit to make cooling easier.

We discussed now the basic losses in a switching regulator. When digging deeper in to the subject, there are some more things to improve, not only in efficiency, but also reliability (remember the catastrophic failure modes?). Those items are partially correlated: a cool circuit (high efficiency) will last longer than a hot circuit. This is particularly true for the used capacitors, and who ever opened a PC power supply will see, that in this case, bigger is better. This holds for the capacitance, the voltage and the size. Larger capacitances will give us less voltage ripple on the output, and increased voltage rating gives us more safety not to cross the operating limits. But together with the mechanical size, this also gives us an important hint to the equivalent series resistance (ESR) of the cap. When operating at the switching frequency of 100 kHz, this series resistance will degrade the capacitors performance and together with the ripple currents, it will heat up the capacitor. Not only we again lose here efficiency, but increased temperatures reduces the capacitors lifetime and leads to blow-up (eye-witnessed..)! I therefore strongly recommend using low ESR electrolytic capacitances with a temperature rating of 105 deg C (search your preferred suppliers catalogue). When in doubt, take the bigger one, as when comparing the sizes of capacitors with same value and voltage rating, the bigger ones have usually lower ESR. This holds only true for modern age devices.. don't try to use these parts sitting in your drawer for 20 years.

To squeeze now the very last percents of the switching regulator, identify the high-current copper traces of your circuit and solder some copper wires (1.5-2.5 mm²) across. This will increase their cross section and decrease the resistance, which is certainly observable at the currents we are dealing with.

Circuit description

After carefully evaluating of all the components, it's time to build up the circuit. The schematics is shown in Fig. 9. Aside from the already described components, there are some more blocking capacitors for the controller IC, the current sense resistor, the slope compensation and the voltage divider for setting the output voltage. The latter one can be calculated easily with

$$V_{out} = R1 / (R1 + R2) * 1.25 V$$

taking $R1 = 2.2 \text{ k}\Omega$ and $R2 = 47 \text{ k}\Omega$, this gives us almost exactly 28 V. The current sense resistors are calculated with the formula given in the datasheet to 0.06 Ohm, which are implemented as three 0.2 Ohm resistors (SMD 2512 case). The slope compensation counteracts an inherent instability of this type of topology. While I won't describe the mechanism here in detail (see datasheet for an explanation), this instability results in decreased efficiency and a high-pitched sound emitted by the circuit, although the more 'experienced' hams won't hear it anymore... While this sound may be known to many of us as 'normal' for switched regulators, surprisingly our design is completely silent when operating properly.

The PCB layout can be seen in Fig. 5 and 6. The used material is standard FR4, 89 x 56 mm, 1.6mm thickness with 35um copper. Note the close proximity of the power devices. While this is not exactly a HF application, short traces are also important here due to the high currents involved.

A common problem when using switching regulators at a ham radio environment is the switching frequency, which is radiated over the supply lines. To make the regulator 'silent', we used a metal box and inserted two little separate chambers at the in- and output, where some lowpass filtering is performed. The needed high-current inductors were taken out of an old PC power supply. The high frequency performance is not so important here as for the main inductor, as the basic component at the input/output lines is the DC current. Feedthroughs into all chambers are done with three 1 nF feed-through caps.

Results

Two units have been built so far by the author and Charles Suckling, G3WDG. To test the behaviour under load, the author used high power resistors. With the enhanced high-current traces and the gate driver, 93 % efficiency could be verified at 4.58 A output current. This is equivalent to a output power of 128 W and an input power of 138 W. The remaining 10 W are lost mainly in the FET and the diodes, but one also can feel the inductor and the current sense resistors getting slightly warm. Nevertheless, 10 W are easily dissipated with a handy heatsink. Testing together with the 50 W 23cm PA kit was done by G3WDG, showing good performance of the combination of the two designs. While measurements showed, that even without the output power filter, no sidebands are generated within the PA, the filter is recommended to keep your own shack 'quiet'. The remaining ripple with the filter is about 20 mV pk-pk (Figure 8).

A 'long-time' test was performed by constantly drawing 128 W power from the regulator for over 7 minutes and temperatures measured on and in the closed box. With an ambient temperature of 20 deg C, the location just outside the MOSFETs heated to about 30 deg C and the inside air temperature to about 40 deg C. The load resistors definitely were warmer..

Conclusion

We presented a high power, high performance switching regulator circuit designed for 12 V use of modern LDMOS power amplifier kits. The overall efficiency is > 90% at > 100W regulated power. As far as possible, standard market components were used. Test were performed with artificial loads as well as with a 23cm PA.

Acknowledgements

I would like to thank first of all Charlie, G3WDG for the nice 23 cm PA design, without which I probably would have never start designing this. He also was beta tester for the circuit. I further would like to thank Coilcraft, Texas Instruments and Fairchild for supplying samples.

Literature

[1] Carl Nelson, LT1070 Design Manual, Application Note 19, www.linear.com

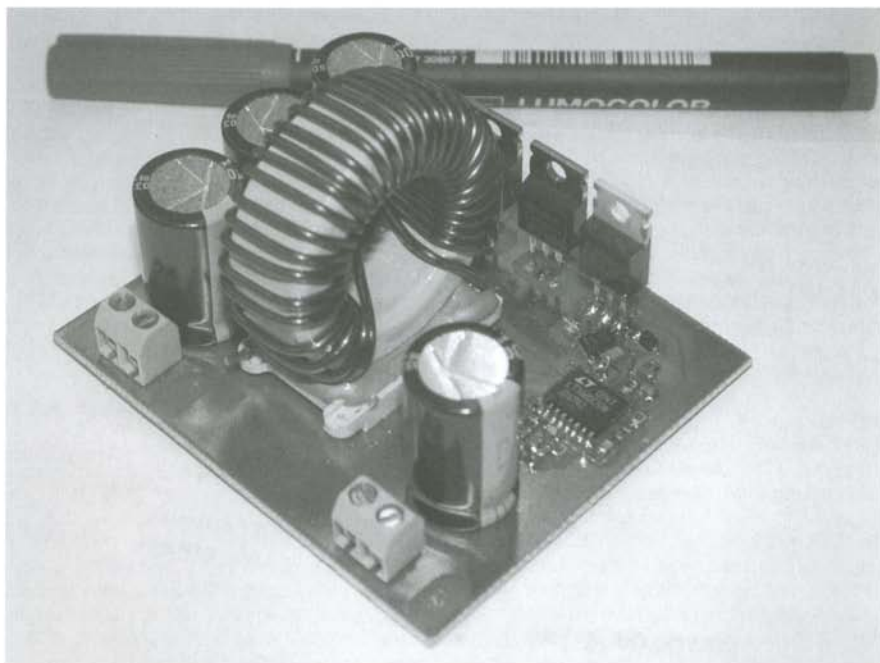
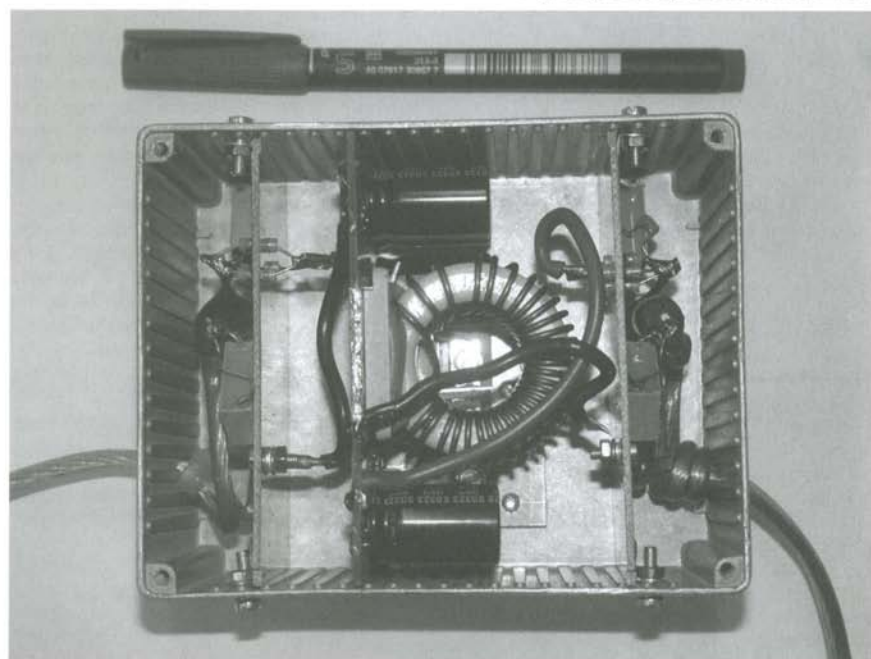


Fig 1. (top): Complete board

Fig 2. (bottom): Board in the box



28 V - Schaltregler - Stromversorgung für LDMOS - PA's

Achim Vollhardt, DH2VA / HB9DUN

In den vergangenen Monaten wurden LDMOS PA's für das 23cm Band populär. Zwar haben diese neuen Bauelemente einen besseren Wirkungsgrad als die allgemein bekannten Mitsubishi Leistungsmodule (50% gegenüber 35%), die Versorgungsspannung von 28 V stellt jedoch Neuland für die meisten OMs dar. Trotz den geringeren Problemen mit Kontaktwiderständen aufgrund der geringeren Ströme bei der höheren Spannung, werden die meisten Amateurfunkstationen mit 12 V betrieben. Dies trifft insbesondere auf die tragbaren Stationen zu, welche auf Bleibatterien als Stromversorgung angewiesen sind. In Verbindung mit den neuen Leistungsverstärkern ergab sich so ein Bedarf an einer einfachen Stromversorgung mit höchstmöglichem Wirkungsgrad.

Aufgrund der anhaltenden Entwicklung auf dem Gebiet der geschalteten Stromversorgungen, sind diverse Bausteine und Schaltungen erhältlich. Selbst für Neueinsteiger, welche sich bisher nur mit Spannungsreglern der 78XX/79XX Baureihe beschäftigt haben, werden Schaltregler zu einer echten Alternative für Hochleistungsanwendungen. Als Beispielanwendung wird eine Stromversorgung für die G3WDG 50W PA vorgestellt.

Bauteileauswahl

Der Betriebsstrom der G3WDG 23cm PA ist 4.5 A bei 28 V Drainspannung (126 W). Um diese Spannung aus einer 12 V Versorgung zu erzeugen, verwenden wir die sogenannte 'Boost'-Topologie, welche in Bild 3 dargestellt ist.

Grundlegende Bestandteile jedes Schaltreglers sind die Induktivität, die Diode und der Schalter (welcher effektiv ein MOSFET ist). Wenn man diese in der Boosttopologie verwendet, wird die Eingangsspannung auf eine höhere Spannung 'geboostet'. Für eine detaillierte Erklärung verweise ich auf [1]. Mit einem Blick auf den Schaltplan können wir die ersten möglichen Quellen für Leistungsverluste erkennen, welche später noch wichtig werden:

- der ohmsche Widerstand der Induktivität
- der ohmsche Widerstand des MOSFET Schalters
- die Vorwärtsschwellspannung der Diode

Um die Boost-Topologie zu einem vollwertigen Schaltregler zu vervollständigen benötigen wir noch eine Rückkopplungsschleife, welche die Eigenschaften des Schalters entsprechend der Last einstellt. Diese musste früher mit diskreten Bauteilen aufgebaut werden. Heutzutage existieren auf dem Markt verschiedene hochintegrierte Schaltregler Kontroll-ICs, welche eine Überstromerkennung, Softstarteigenschaften und eine Feedbackschleife für die Spannungsregelung beinhalten. Ich habe mich für den LT1680 von Linear Technologies entschieden: der Regler wurde speziell für die Boost-Topologie mit 12 V Eingangsspannung entwickelt. Weiterhin kann er direkt das

Gate eines Hochstrom-MOSFETs ansteuern, welche im allgemeinen hohe Gate Kapazitäten haben.

Neben dem Kontroll-ICs ist die Induktivität die zentrale Komponente eines solchen Reglers. Sie muß selbst bei hohen Strömen noch genug Induktivität für den Betrieb des Schaltreglers bereitstellen. Wird der maximale Strom überschritten, kann der Spulenkern keinen zusätzlichen magnetischen Fluss mehr aufnehmen, die Induktivität ist in Sättigung. Dies kann zu einem katastrophalen Versagen der Schaltung führen und ist der Grund für die Abneigung vieler OMs gegenüber Schaltreglern.

Aus diesem Grunde muß man bei der Wahl der Induktivität sehr vorsichtig sein. Im Zweifel lieber eine Nummer größer wählen! Ich habe mich für die DMT-3 Induktivität von Coilcraft entschieden, welche die größte Drossel in ihrem Katalog ist. Laut dem Datenblatt hat sie selbst bei einem Strom von 12 A noch eine Induktivität von 35 µH. Dies wird auch notwendig sein, da unser Schaltregler mehr als 10 A von seiner Stromquelle abverlangen wird, und dies auch nur bei 100% Wirkungsgrad. Da wir eher einige Prozente darunter liegen werden, werden es sicherlich mehr als 10A werden..

Das nächste auszuwählende Bauteil ist der MOSFET-Schalter. Die maximalen Werte für V_{ds} und I_{ds} sollten deutlich höher als 28 V und 12 A liegen. Um die ohmschen Verluste zu minimieren, welche anfallen solange der MOSFET leitend ist, benötigen wir minimales $R_{DS(on)}$. Dies sollte im Bereich von 10mOhm liegen. Weiterhin gibt es einen weiteren, noch nicht diskutierten Verlustfall: die dynamischen Verluste des MOSFET. Ein idealer mechanischer Schalter hat entweder $R=0$ Ohm oder $R=\infty$ Ohm mit einem augenblicklichen Übergang zwischen den beiden Zuständen. In beiden Zuständen ist der Schalter verlustfrei. Ein realer MOSFET hat intrinsische Anstiegs- und Abfallzeiten, während derer er einen endlichen Widerstand darstellt, an dem Leistung verloren geht. Diese Verlustleistung kann einfach berechnet werden:

$$P = 0,5 \cdot V_{out} \cdot (I_{out} / (1-DC)) \cdot (T_r + T_f) \cdot f$$

mit $1-DC = V_{in}/V_{out}$ (DC: Duty Cycle) und Schaltfrequenz f . Man erkennt, dass die Anstiegs- und Abfallzeiten t_r und t_f deutlich zu den Verlusten des Schaltreglers beitragen. Um diese zu minimieren habe ich den FQP85N06 als MOSFET gewählt. Mit einem $R_{DS(on)}$ von 10 mOhm und Flankenzeiten um die 200ns ist dies ein recht guter FET für unsere Anwendung. Diese Zeiten können natürlich nur dann erreicht werden, wenn die Ansteuerschaltung des FET-Gates die Gate-Kapazität auch schnell genug laden und entladen kann. Dies kann mit dem LT1680 trotz seiner recht hohen Treiberströme problematisch werden, da wir eine Gate-Kapazität von ca. 5ns in 200ns laden müssen. 5nF sind bei 12V 60nC, 60nC in 200ns sind 0.3 A, das bedeutet, dass unser Kontroll-IC für 200ns Anstiegszeit einen Strom von 0.3 A liefern muss. Um hier also weitere Prozente herauszuschinden, kann ein sogenannter Gate-Treiber, welcher für genau diesen Zweck entwickelt wurde, verwendet werden. Ich habe den UCC37322 von Texas Instruments eingesetzt, welcher bis zu 9A liefern Puls-

ströme liefern kann.. das sollte reichen. Bei Nichtverwendung wird der Steuerausgang des Kontroll-ICs direkt an das MOSFET Gate angeschlossen.

Kommen wir nun zur Diode. Wenn wir bei einer Diode eine typische Vorwärtsspannung von 0.7V annehmen und einen Betriebsstrom von 10A sind dies bereits 7W Verlustleistung, was bei einer 100W Anwendung 7% Verluste entspricht. Um diese Verluste zu minimieren, verwenden wir Schottky Dioden mit niedriger Vorwärtsspannung, welche für 10A und mehr ausgelegt sind. Zusätzliche Verbesserungen ergeben sich durch Parallelschaltung von Dioden, da die Vorwärtsspannung nicht konstant ist, sondern sich mit niedrigerem Strom verringert (ich gehe davon aus, dass der Leser die prinzipielle I/U Kurve einer Diode kennt). Ich habe die die MBR20100CT Doppel-Schottkydiode verwendet und zwei von ihnen parallelgeschaltet, welches insgesamt vier Dioden ergibt. Als Nebeneffekt wird die Verlustwärme auch mechanisch etwas verteilt, dies vereinfacht die spätere Kühlung.

Wir haben nun die grundlegenden Verlustmechanismen eines Schaltreglers diskutiert. Wenn man sich detaillierter mit dem Thema befaßt, lernt man weitere Verbesserungsmöglichkeiten kennen, die sich nicht nur mit dem Wirkungsgrad, aber auch mit der Zuverlässigkeit beschäftigen (Stichwort: katastrophales Versagen). Dies ist zum Teil gekoppelt: eine kühle Schaltung (hoher Wirkungsgrad) lebt länger als eine heiße. Für die Kondensatoren gilt dies insbesondere und wer jemals ein PC-Netzteil geöffnet hat erkennt, dass hier 'größer ist besser' gilt, und zwar für Kapazität, Spannungsfestigkeit und mechanische Größe. Eine höhere Kapazität ergibt einen geringeren Ripple aus der Ausgangsspannung und eine höhere Spannungsfestigkeit mehr Sicherheitsspielraum um im Betriebsbereich des Bauteils zu bleiben. Zusammen mit der mechanischen Größe bekommen wir jedoch zusätzlich einen Hinweis auf den äquivalenten Serienwiderstand (engl. ESR). Bei der Betriebsfrequenz von 100 kHz verringert der ESR die Leistungsfähigkeit des Kondensators. Zusammen mit den Rippleströmen, welche durch den Kondensator fließen, verursacht der ESR aber auch eine Erwärmung des Bauteils. Dies ist nicht nur ein weiterer Verlust von elektrischer Leistung, sondern auch eine Verringerung der Lebensdauer bis hin zum explosiven Versagen (auch schon erlebt..!) Elektrolyt-Kapazitäten mit niedrigem ESR und einem Temperaturlimit von 105 Grad C sind hier dringend empfohlen (gibts im nächsten besten Elektronikatalog). Im Zweifel den größeren Elko nehmen, da bei gleichen Kapazitäten und Spannungsfestigkeiten das größere Bauteil normalerweise einen niedrigeren ESR hat. Dies stimmt allerdings nur bei modernen Typen, bitte nicht die uralten Komponenten aus der Bastelkiste verwenden.

Um nun die allerletzten Prozente an Wirkungsgrad aus der Schaltung zu quetschen, lötet man am besten Kupferdraht (1.5-2.5mm²) auf die Leiterbahnen, welche die hohen Ströme durchleiten müssen. Damit wird ihr Querschnitt erhöht und der Widerstand verringert.. der Effekt ist durchaus bemerkbar bei den vorliegenden Stromstärken.

Schaltungsbeschreibung

Nach der umfangreichen Diskussion der Komponenten bauen wir nun unsere Schaltung auf. Der Schaltplan ist in Abb. 9 dargestellt. Zusätzlich zu den bereits beschriebenen Bauteilen enthält er noch Abblockkondensatoren für den Kontroll-IC, einen Widerstand zur Stromüberwachung, die Slope Kompensation und den Spannungsteiler für die Ausgangsspannung. Letzterer kann einfach wie folgt berechnet werden:

$$V_{out} = R1 / (R1 + R2) * 1.25V$$

mit $R1=2.2k\Omega$ und $R2=47k\Omega$ ergibt das ziemlich genau 28V. Der Widerstand zur Stromüberwachung wird mit der über das Datenblatt zu 0.06 Ohm bestimmt, die kann mit einer Parallelschaltung von drei 0.2 Ohm Widerständen (SMD 2512) erreicht werden. Die Slope-Kompensation verhindert eine mögliche Instabilität, welche bei der Boost-Topologie auftreten kann. Ohne diesen Prozess im einzelnen zu erläutern (siehe dazu auch LT1680 Datenblatt), sei hier angemerkt, dass die Instabilität sich durch verringerten Wirkungsgrad und einen sehr hohen Ton auszeichnet, welcher von der Schaltung erzeugt wird. Dieser Ton, welcher von den 'erfahrenen' OMs wahrscheinlich nicht mehr wahrgenommen werden kann, wird allgemein mit Schaltreglern assoziiert. Überraschenderweise ist unser Schaltregler bei ordnungsgemäßen Betrieb vollkommen ruhig.

Abb. 5 + 6 zeigen das Leiterplattenbild. Es wird normales FR4 Material, 89 x 56 mm, mit 1.6mm Dicke und 35um Kupfereauflage verwendet. Die Leistungsbauteile sind alle dicht beisammen angeordnet. Dies ist zwar nicht wirklich eine HF Anwendung, durch die hohen Ströme sind jedoch auch hier kurze Wege vorteilhaft.

Ein verbreitetes Problem von Schaltreglern im Amateurfunk ist die Abstrahlung der Schaltfrequenz über die Zuleitungen. Um den Schaltregler 'ruhigzustellen' wird er in ein Metallgehäuse eingebaut. In zwei internen Kammern am Ein- und Ausgang sind Tiefpassfilter eingebaut. Die benötigten Hochstrom-Spulen kommen aus einem alten PC-Netzteil. Die Hochfrequenzeigenschaften sind hier nicht so entscheidend wie für die Schaltinduktivität, da die Hauptkomponente der Ein-/Ausgangsleitungen Gleichstrom ist. Die Durchführungen in die Kammern sind mit je drei 1nF Durchführungskondensatoren realisiert.

Ergebnisse

Zwei Exemplare wurden vom Autor aufgebaut, ein weiteres von Charlie Suckling, G3WDG. Um den Schaltregler unter Vollast zu testen, wurden Hochlastwiderstände verwendet. Mit den verstärkten Hochstromleiterbahnen und dem Gate-Treiber wurde jeweils ein Wirkungsgrad von 93% bei 4.58 A Ausgangsstrom ermittelt. Dies entspricht einer Ausgangsleistung von 128W und einer Eingangsleistung von 138W. Die verbleibenden 10W werden größtenteils im FET und den Dioden verloren, die Induktivität und die Stromüberwachungswiderstände werden ebenfalls leicht warm. Dennoch können die 10W sehr einfach mit einem kleinen Kühlkörper abgeführt werden. Tests mit einer 50W 23cm PA wurden von

G3WDG durchgeführt, welche die Kompatibilität der beiden Designs bestätigten. Obwohl Messungen ohne die beschriebenen Tiefpassfilter zeigten, dass keine Seitenbänder in der PA erzeugt werden, ist das Filter empfohlen, um das Shack 'ruhig' zu halten. Der verbleibende Ripple mit dem Filter ist in Bild 8 zu sehen, er beträgt 20mV pk-pk.

Ein 'Langzeit'-Test wurde mit einer Belastung von 128 W für über 7 Minuten durchgeführt. Hierbei wurden die Temperaturen außen am Gehäuse und innen gemessen. Mit einer Umgebungstemperatur von 20 Grad C wurde die Stelle direkt außerhalb des MOSFET 30 Grad C warm und die Luft im inneren knapp 40 Grad C. Die Lastwiderstände waren noch etwas wärmer.

Zusammenfassung

Es wurde ein Hochleistungs-Schaltregler mit hohem Wirkungsgrad vorgestellt, der es ermöglicht, moderne LDMOS Leistungsverstärker an 12V zu betreiben. Der Gesamtwirkungsgrad beträgt mehr als 90% bei über 100W Ausgangsleistung. Soweit möglich wurden kommerzielle Standardbauteile verwendet. Tests wurden sowohl mit künstlichen Lasten als auch mit einer 23cm PA durchgeführt.

Danksagungen

Zunächst möchte ich Charlie, G3WDG für die wunderbare 23cm PA danken, ohne die ich sicherlich nicht diesen Schaltregler entwickelt hätte. Er war ebenso Beta-Tester für diese Schaltung. Weiterhin danke ich Coilcraft, Texas Instruments und Fairchild für die Bauteilemuster.

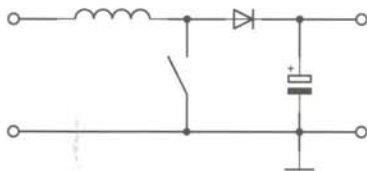
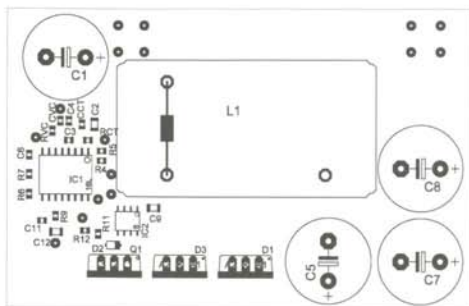


Abb. 3: Prinzipschaltbild
Fig. 3: „Boost“ Circuit



L-Band Oscillator Chain

Daniel Uppström, SM6VFZ

Introduction

When converting up to and down from the bands above 1 GHz one needs some sort of local oscillator (LO). Modern PLL ICs are now available for frequencies up to 2-3 GHz, but the classical approach of using a crystal oscillator at something like 100 MHz followed by multiplier stages is still competitive and features in general less phase-noise than a PLL solution operating directly at the target frequency. By using SMD components only, an oscillator chain with high performance and stability has been implemented in a small board size. It delivers a signal in the L-band region centered at 1.2 GHz, which is achieved by a total multiplication factor of twelve. By the insertion of varactor diodes in series with the crystal it is possible to fine-tune the frequency or lock it to an external frequency normal. Most testing and evaluation have been done with a crystal at 96 MHz and a target frequency of 1152 MHz (which is needed for conversion between 1296 and 144) but the circuit is designed to be versatile for a variety of applications and tests at other frequencies have been done as well, see table 1. Spurious were -40 dBc or better in all listed cases.

Output (xtal) Freq.	Measured Output Power (mW)
1128 (94)	18
1152 (96)	20
1278 (106.5)	12
1291 (~107.6)	11

Table 1

Circuit description

The oscillator circuit is of butler type and follows much of the guidelines mentioned in (1). The overall gain of the two transistors (Q1 and Q2) is designed to be moderate in order to avoid spurious oscillations. The oscillation voltage is also cut at 0.6 V peak-to-peak by the two anti-parallel schottky diodes D1. This ensures that the transistors work in a linear fashion. If the diodes were not present, the amplitude of oscillation would be determined by the saturation point of the transistors and the circuit would behave non-linear and let low-frequency noise mix with the carrier. This noise would then directly add up to the phase-noise of the LO which would worsen the perform-

ance. By lightly loading the transistors the effect of temperature variations on frequency and power level is also reduced.

The crystal operates in series resonance mode at the fifth overtone, which is achieved by tuning C3 to be resonant with L1 at this frequency region. The coil L2 tunes out the holder capacitance. By introducing the varactors D2 and D3 it is possible to pull the frequency upwards somewhat. Since it is often necessary to pull more symmetrically around the specified frequency, a coil, L4, can also be introduced to shift the pulling range to cover downward frequencies as well. But coils always have some loss which will hurt the Q of the resonance circuit and increase phase-noise. For best performance the crystal should probably be ordered some 3 kHz low so that L4 can be excluded. See (1). Resistor R3 removes excessive gain in the oscillator and helps maintain a well defined oscillation and a low phase-noise. It should be noted that in some cases a little more gain is needed for the oscillator to behave well and the value of R3 could be increased or the resistor could even be removed. This seems to be more important the higher the oscillation frequency is.

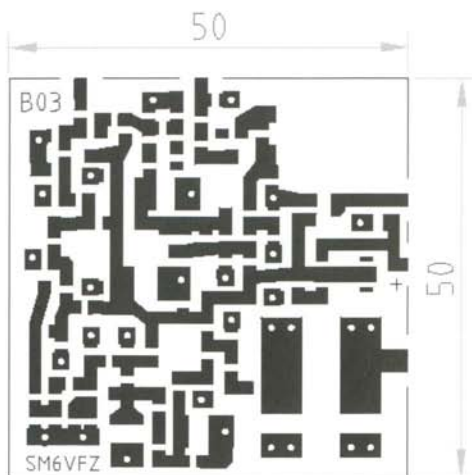
The oscillator is followed by a buffer stage (Q3) that isolates it from the chain and also makes it possible to connect an external circuit in order to monitor the frequency. By feeding the oscillator signal to some PLL/FLL-circuit and the resulting error signal to the Vtune port, the frequency could be locked to some frequency normal. See for instance (2). If no such locking is necessary, a tuning voltage is available from VR1. The tuning voltage is applied through the coil L3 which should be self-resonant near the operating frequency for best performance.

The oscillator signal is amplified in Q4 and fed into Q5 which then behaves nonlinear due to the high signal level. The fundamental frequency is damped in a fixed value filter while the second harmonic is able to pass through to the next stage. The collector circuit of Q6 forms part of a highly selective filter which, when properly tuned, turns the stage into a tripler. To achieve the needed total factor of twelve, the frequency is then doubled in Q7 which is also followed by a selective filter that needs some tuning. The coil L9 should have a value that makes it operate near the self-resonance frequency.

Construction and tuning

The board material is 1.6 mm thick FR4 and it measures 50x50 mm. One side is completely unetched and serves as ground layer. All holes indicated in the layout are simply through connections to ground that needs some millimeters of wire. Since the board is rather thick it radiates at high frequencies and some screening should be mounted around it.

The tuning procedure starts by setting VR1 to its middle position and connecting it to Vtune. With a frequency counter connected to the "osc out" port (or a receiver loosely connected), C3 can be tuned for oscillations in the right frequency region. It should also be checked that the exact frequency could be set by VR1. When the oscillator seems to behave well, the filters in the chain could be tuned. Points of resonance should first be searched for where the trim capacitors are not very far from their middle positions. Notice also that the optimum point of C17 is quite dependent on C20 and vice versa. When the right points are found, an output power in the region of 10-13 dBm should be expected, depending on output frequency. Unwanted spurious should be more than 40dB below the carrier if properly tuned. If one tolerates a little less spurious suppression it should be possible (but tricky!) to tune the circuit without access to a spectrum analyzer. One could then get along with just a power meter and a frequency counter, although spectrum analysis is of course highly recommended.



Layout

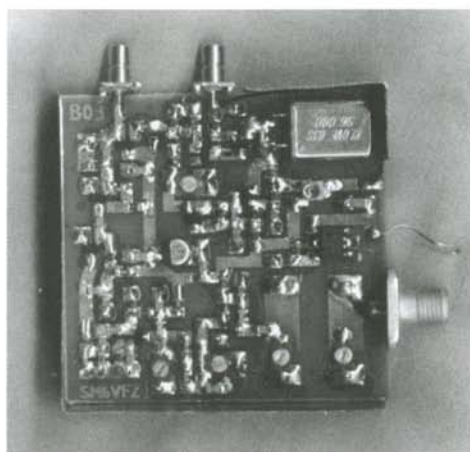


Fig.: L-Band Oscillator Chain

Correspondance with the author could be sent to: upda@eta.chalmers.se

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L-Band Oszillator-Kette

Daniel Upström, SM6VFZ

Einleitung

Wenn man auf bzw. von den Bändern oberhalb 1 GHz umsetzt, benötigt man einen Local Oscillator (LO). Moderne PLL ICs sind zwar inzwischen für Frequenzen bis 2-3 GHz verfügbar, aber die klassische Methode eines Quarzoszillators bei um 100 MHz gefolgt von Verfielfacherstufen ist immer noch konkurrenzfähig und hat normalerweise den Vorteil eines geringeren Phasenrauschens als eine PLL-Lösung, die direkt auf der Zielfrequenz arbeitet.

Durch die ausschließliche Verwendung von SMD-Bauteilen wurde eine Oszillator-Kette mit hoher Performance und Stabilität auf einem kleinen Board realisiert.

Es wird ein Signal im L-Band geliefert, das auf 1.2 GHz zentriert ist und insgesamt durch eine 12-fache Multiplikation erzeugt wird.

Durch die Verwendung von Varactor-Dioden in Serie mit dem Quarz ist es möglich, die Frequenz fein abzustimmen oder an ein externes Frequenznormal zu koppeln.

Die meisten Tests und Untersuchungen wurden mit einem 96-MHz-Quarz und einer Zielfrequenz von 1152 MHz durchgeführt. (Diese wird für die Umsetzung zwischen 1296 und 144 MHz benötigt.) Die Schaltung ist aber so ausgelegt, dass sie für verschiedene Frequenzen variabel ist und es wurden auch Test mit anderen Frequenzen durchgeführt. Siehe hierzu Tabelle 1. Unerwünschte Harmonische werden in allen Fällen -40dB oder besser unterdrückt.

Output (xtal) Freq.	Gemessene Ausgangsleistung (mW)
1128 (94)	18
1152 (96)	20
1278 (106.5)	12
1291 (~107.6)	11

Tabelle 1

Die Schaltung

Es handelt sich um einen Butler-Oszillator, der sich überwiegend nach den Richtlinien von (1) richtet. Die gesamte Verstärkung der zwei Transistoren Q1 und Q2 wurde niedrig gehalten, um unerwünschte Harmonische gering zu halten. Die Oszillator-Spannung wurde ebenfalls auf 0.6 V Spitze-Spitze durch zwei antiparallele Schottky-Dioden (D1) begrenzt. Dadurch wird lineares Arbeiten der Transistoren abgesichert.

Ohne diese Dioden würde die Amplitude der Schwingung durch die Sättigungspunkte der Transistoren begrenzt. Die Schaltung würde sich unlinear verhalten und Rauschen niedriger Frequenz würde sich mit dem Träger mischen. Dieses Rauschen würde sich dann direkt zum Phasenrauschen des LO addieren und die Performance verschlechtern. Durch leichtes Belasten der Transistoren wird der Einfluß von Temperaturschwankungen auf Frequenz und Höhe der Ausgangsleistung verringert.

Der Quarz arbeitet in Serienresonanz auf dem 5. Oberton. Dies wird durch die Abstimmung von C3 auf Resonanz mit L1 in diesem Frequenzbereich erreicht. Die Spule L2 stimmt die Halter-Kapazität ab. Durch das Einfügen der Varactor-Dioden D2 und D3 ist es möglich, die Frequenz etwas nach oben zu ziehen.

Da es oft nötig ist, um eine bestimmte Frequenz symmetrisch herum ziehen zu können, kann eine weitere Spule (L4) eingefügt werden, um den Ziehbereich zu verschieben und auch niedrigere Frequenzen zu erreichen.

Aber Spulen haben immer auch Verluste, die die Güte des Schwingkreises verringern und das Phasenrauschen erhöhen. Für beste Performance sollte deshalb wahrscheinlich der Quarz etwa 3 kHz niedriger bestellt werden, so dass man L4 weglassen kann. (Siehe (1)).

Der Widerstand R3 verhindert zu starke Verstärkung im Oszillator und hilft eine gut definierte Schwingung und niedriges Phasenrauschen zu erhalten. Man sollte hier anmerken, dass in einigen Fällen etwas mehr Verstärkung für den Oszillator benötigt wird, damit dieser sich „gutartig“ verhält und R3 sollte vergrößert werden oder sogar ganz weggelassen werden.

Dies scheint um so wichtiger sein, je höher die Oszillator-Frequenz ist.

An den Oszillator folgt eine Puffer-Stufe (Q3), die die ihn von der Kette isoliert und es außerdem ermöglicht, eine externe Schaltung anzukoppeln, um die Frequenz zu überwachen. Durch Einspeisen des Oszillator-Signals in eine PLL/FLL-Schaltung und des resultierenden Fehlersignals an den VTune-Port, kann die Frequenz an ein Frequenznormal gekoppelt werden.

Siehe z.B. Referenz (2). Falls solch eine Kopplung nicht nötig ist, ist eine Steuerspannung von VR1 verfügbar. Die Steuerspannung wird durch die Spule L3 zugeführt, die für beste Performance ihre Eigenresonanz nahe der Arbeitsfrequenz haben sollte.

Das Oszillator-Signal wird mit Q4 verstärkt und an Q5 gegeben, der dann aufgrund der hohen Pegel nicht linear arbeitet.

Die Grundfrequenz wird durch ein fest dimensioniertes Filter gedämpft, während die 2. Harmonische hindurch bis zur nächsten Stufe gelangt.

Der Kollektor-Kreis von Q6 stellt den Teil eines hochselektiven Filters dar, das, wenn es ordnungsgemäß abgestimmt ist, die Stufe als Verdreifacher arbeiten lässt. Um den benötigten Faktor 12 zu erhalten, wird die Frequenz mit Q7 verdoppelt, ebenfalls gefolgt von einem selektiven Filter, das besonders abgestimmt werden muss.

Die Spule L9 sollte einen Wert haben, der eine Arbeiten nahe der Eigenresonanz ermöglicht.

Aufbau und Abstimmung

Die Platine ist 1.6mm dickes FR4 in der Größe 50 x 50mm. Eine Seite bleibt komplett ungeätzt und dient als Massefläche. Alle Löcher im Layout sind einfache Durchkontaktierungen mit einem mm Draht nach Masse.

Da die Platine relativ dick ist, schwingt sie auf hohen Frequenzen und es sollte eine Abschirmung um sie herum aufgebaut werden.

Die Abstimmung beginnt mit der Einstellung von VR1 auf die Mittelstellung und des Anschlusses an VTune. Mit einem Frequenzzähler, der am „osc out“ – Port angeschlossen wird (oder einem lose angekoppelten RX), kann C3 auf einen Wert für Schwingungen im richtigen Bereich abgestimmt werden. Man sollte auch kontrollieren, dass die exakte Frequenz mit VR1 eingestellt werden kann. Sobald der Oszillator gut zu arbeiten scheint, können die Filter in der Kette abgestimmt werden. Die Resonanzpunkte sollten zuerst dort gesucht werden, wo die Trimmer nahe der Mittelstellung sind.

Es ist zu beachten, dass der optimale Punkt bei C17 ziemlich von C20 abhängt und umgekehrt.

Wenn die richtigen Punkte gefunden wurden, sollte eine Ausgangsleistung von etwa 10-13dBm erwartet werden, je nach Output-Frequenz.

Unerwünschte Harmonische sollten mehr als 40 dB unterhalb des Carriers liegen, wenn sauber abgestimmt wurde. Wenn man mit etwas schlechteren Werten auskommt, sollte es möglich (ist aber nicht ganz einfach!), die ganze Schaltung ohne Spektrum Analyzer abzugleichen. Man geht dann die Stufen nur mit Power-Meter und Frequenzzähler durch. Natürlich ist ein Spektrum-Analyzer der bessere Weg.

Korrespondenz mit dem Autor kann gesendet werden an: upda@eta.chalmers.se

Practical and Cheap Solution of Circular Waveguide for 10 GHz

Ivo Chladek, ZS6AXT, OK2WCG

Tropo operation on microwave frequencies from home usually requires some low loss feedline to reach the antenna on the top of your tower or roof. This is of course so unless you mount your transverter, PA and LNA in a waterproof enclosure somewhere around the dish and live with the access problems when something goes wrong with it.

Before starting on my 10 GHz home tropo setup, good consideration was given to possible options. Great help were articles as listed in the references and comments from few microwave guys. In the end decision for a circular waveguide (WG), "plumber's delight" was made. The following description is certainly nothing new for many, but it may give some new ideas to others. Similar arrangement can be used for 5760 MHz and probably even for 24 GHz if a suitable size of tubing is available.

WG is certainly the feedline with lowest losses in the microwave range of frequencies. For 10 GHz it is WR90 (although WR75 can be used too), but unless you get some surplus it is expensive. In addition, one needs flanges, bends etc. and in general it is quite clumsy and heavy. Loss is about 10 dB / 100 feet, i.e. over 3 dB / 10 m.

The other alternative is circular WG, which is not very often used by commercial users. However, suitable size of copper tubing is available, as used by plumbers for water installations. It is the 21mm I.D. copper tubing, with 0,5mm thick wall, which is just the right size for 10 GHz WG. In addition, straight and angle couplings are available, all of that for reasonable prices. Such copper tubing is also not as heavy as WR90 WG. There is just one disadvantage: polarization may change in the length of such WG, especially when using angle couplings, thus one end of such feedline must have adjustable polarization. However, 10 m length of this WG, assembled with three 45 degree and two straight connections, plus two coax to WG transitions has just 2,2 dB loss in my case. Similar second 10 m length measured about 0,1 dB more. This was measured in home conditions, with very basic equipment. Considering the costs, availability of materials for this WG and ease of assembly, it is really a feedline worth considering.

Coax to WG transition

This is the only bit complicated and critical part of the WG assembly. It consists of the following parts:

- Short piece of the 21mm I.D. tubing, about 15 – 25 cm long,
- Brass plate 16x16 mm, 2 mm thick,
- SMA panel mount connector with four fixing holes and the center lead of 1mm diameter, at least 9,5 mm long and
- Thin round brass plate of 22 mm diameter.

Both ends of the tubing should be square, flat and with clean smooth edges. Drill a 3 mm diameter hole exactly 12,2 mm from one end of the tubing. Prepare the SMA mounting plate by drilling 3 mm hole in the center and four SMA fixing holes tapped for M2,5 screws. Select a short stainless steel M3 screw with nut and fix the SMA mounting plate to the WG tubing. Solder together such, that the solder does not get into the tapped M2,5 holes. Remove the M3 screw and enlarge the hole to 4 mm. Clean the edges of this hole. Cut the TEFLON on the center lead of the SMA to such length, that it is just flush with the inner wall of the WG. The protruding bare center lead should be at least 6,5 mm, but rather 7 mm long, to give chance for optimization. Mount the SMA connector with four M2,5 short screws. Stainless steel type screws are the best here.

For the accurate optimization of the transition one needs a 10 GHz signal source, directional coupler and some spectrum analyzer or similar indication. If you do not have access to such equipment, cut the probe length to exactly 6,5 mm, it should be close enough to optimum. That is of course if the distance of the probe center is exactly 12,2 mm from the WG end, since the optimal probe length depends on this distance.

For the proper adjustment we need some sort of dummy load fitted to the other end of the WG since it's open end has quite a bad SWR, like 2:1. A wooden handle of a broomstick, sharpened over about 7 cm length and inserted into the WG may do the job. With the open end (probe side) the return loss will be about 10 dB. Take a metal plate (I used a PCB off-cut) and move it to this WG end. When just few mm away from the open end, return loss will increase, usually well over 20 dB. With a file, remove a small piece of the probe. Do not cut it with sidecutters, you may shorten it too much! After each probe shortening, maximum of return loss should be with the metal plate closer to the WG end. Shorten the probe few times, until return loss maximum is with the metal plate pressed to the end of WG. It should be at least 25 dB, in my case close to 30 dB. Accuracy of such measure-

ment depends on the quality of directional coupler, it's lower coupling will give more accurate results.

After optimization, solder thin layer of solder onto the edge of the WG and then solder the round plate to cover the WG end. It must be watertight, if used outside. Check the return loss, it should be the previously measured optimum value.

Solder the straight coupling to the open end such, that the WG goes to about the center of it. For the coax to WG transition which will be in the shack, make few longitudinal cuts into the other end of coupling with a thin hacksaw, since this end will be fixed to the incoming WG by means of a hose pipe clamp, so that the polarization can be optimized.

The above described technique of the probe length adjustment was used here also for the optimization of 10 GHz W2IMU horn, which is probably the best choice for dishes with $f/D = 0,5 - 0,6$. In this case one does not need the WG dummy load. Horn should be however aimed into reasonably open, reflections free space, during measurements.

WG assembly and alignment

Surprisingly, assembly of the WG parts is not critical. It is advisable to clean the inside of the tubing before assembly. In the straight couplings, ends of the WG tubing should be square, cleaned and touching each other. In the angle couplings push the tubing ends as far as possible. The easiest way is to assemble most of the WG run on the ground; it is quite difficult to do good soldering of the joints up in the air. Gas torch soldering irons are best suited for this job. All joints should be watertight. Do not use any acidic solder flux. According to W1GHZ not more than 45 degree angle couplings should be used for best results. After assembly, inject signal into the shack side of the WG and measure output at the other end. This should be terminated by a good 50 Ohm load with a detector. By slow turning of the coax to WG transition in the shack adjust for maximum output at the other end and fix the position with the hose pipe clamp.

Some practical applications

There are many practical applications of such circular WG. These are not only for tropo, but also for EME use. Sensitive equipment like transverter can stay in the ham shack, while only PA and LNA will be outside, close to the feedhorn, together with the coaxial relay. Due to the low loss in such WG, drive level for the PA can be low and the LNA does not need to have high gain.

In my case I use two runs of the WG for tropo, since the transverter has separate outputs for TX and RX. Thus only one coaxial relay is necessary. With a single WG run two more coaxial relays may

be required. Saving of two such relays will more than cover the cost of the second WG. When using a TWT for tropo, it can be mounted in the ham shack, some 3 dB loss in the WG may be tolerated, considering how many problems with the HV PSU will be eliminated.

And for further experimentation: Bit larger diameter tubing, about 35 mm can be used for two bands WG, namely 3 and 6 cm. However, dimensions of the coax to WG transitions will have to be recalculated since these depend on the WG wavelength. Necessary details can be found in the W1GHZ article listed in References. Coax to WG transitions, one for each band, could be perhaps at right angles; it may work.

Measuring equipment used

Coax to WG transition: HP signal generator, HP 141 spectrum analyzer and 20 dB directional coupler.

WG loss measurement: 200 mW signal from transverter, diode detector calibrated with the use of 1dB SMA attenuators, 20 dB directional coupler and 50 Ohm SMA microwave dummy load.

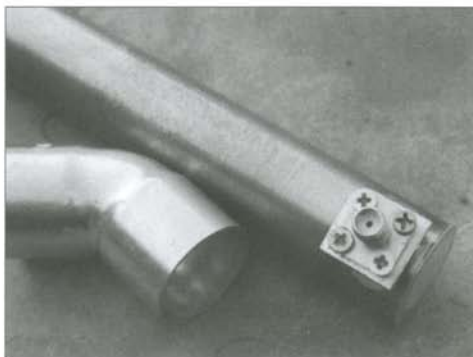
The above measuring equipment is certainly not "state-of-art", but for amateur conditions it was found good enough, after the measured results were confirmed by few contacts.

References

Proceedings of Microwave Update 1999:

-Paul Wade, W1GHZ: Understanding Circular Waveguide - Experimentally.

-Dick Knadle, K2RIW: Circular Waveguide Frequencies, More Accuracy, More Experiment Data.

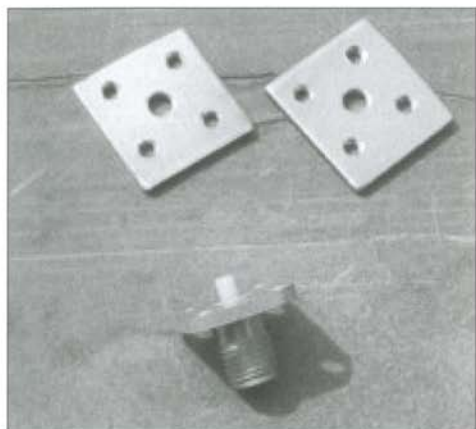


**Complete transition
and 45 degree coupling**

Der Übergang und ein 45 Grad - Verbinder



Brass plate fitted for soldering
Die Messingplatte vor dem Verlöten



SMA and brass plates
SMA-Flanch und Messingplatten

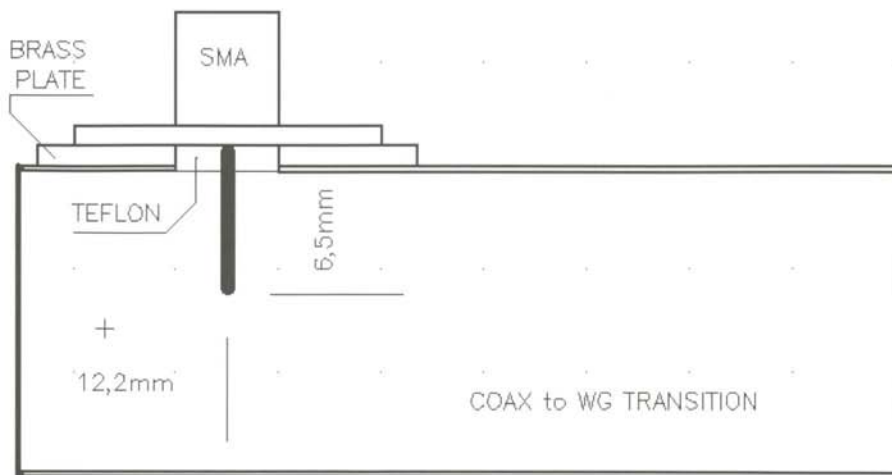


Complete transition
Der fertige Koax-Wellenleiter-Übergang

Ein praktischer und günstiger zirkularer Wellenleiter für 10 GHz

Ivo Chladek, ZS6AXT, OK2WCG

Tropo-Betrieb auf den Mikrowellenbändern von der häuslichen Station aus erfordert ein sehr dämpfungsarmes Kabel zur Antenne auf der Spitze des Mastes bzw. Daches. Natürlich nur, wenn



Tranverter, PA und LNA nicht in einem wasserdichten Gehäuse direkt am Spiegel untergebracht sind. Dann hat man aber ggf. Probleme mit dem Zugang, wenn mal etwas nicht funktioniert.

Bevor ich zuhause meine Tropo-Station für 10 GHz aufgebaut habe, wurden Überlegungen über alle möglichen Optionen angestellt. Sehr hilfreich dafür waren die in den Referenzen aufgelisteten Artikel und auch einige Kommentare von ein paar Mikrowellen-OPs. Letztendlich fiel die Wahl auf einen runden Wellenleiter (Waveguide, WG) der Marke „Klempners Wonne“

Die folgende Beschreibung mag sicherlich für viele nichts Neues sein, kann aber Anregung für neuen Ideen für einige sein.

Eine ähnliche Konstruktion kann auch für 5760 MHz und wahrscheinlich auch für 24 GHz verwendet werden, sofern passende Rohrdurchmesser zur Verfügung stehen.

Der Wellenleiter ist sicher die dämpfungsärmste Möglichkeit Mikrowellen zu übertragen. Für 10 GHz wird WR90 (WR75 kann aber auch verwendet werden) verwendet. Das ist sehr teuer, sofern man nicht Zugang zu günstigem Surplus-Material hat. Außerdem braucht man spezielle Flansche, Krümmungen etc. Das ganze wird recht schwer und ziemlich klobig. Die Dämpfung liegt bei 10db/100feet, d.h. etwas über 3dB / 10m.

Die Alternative dazu ist runder Wellenleiter, der im kommerziellen Bereich recht selten eingesetzt wird. Es gibt jedoch eine passende Größe von Kupferrohr, das von Klempnern für Wasserleitungen verwendet wird. Es handelt sich um ein Rohr mit 21mm Innendurchmesser und einer Wandstärke von 0.5mm. Dies ist eine passende Größe für 10 GHz. Natürlich gibt es passende Winkel und gerade Stücke, die auch verwendet werden können. Und das alles zu akzeptablen Preisen. So ein Kupferrohr ist sogar leichter als WR90. Es gibt nur einen Nachteil: Die Polarisation kann sich beim Laufe eines solchen Wellenleiters ändern, besonders wenn Winkelstücke verwendet werden. Deshalb ist es nötig, an einem Ende des Wellenleiters die Polarisation einstellen zu können. Immerhin haben 10m dieses Wellenleiters, mit drei 45 Grad Winkel-Stücken und 2 geraden Verbindungsstücken plus zwei Koax-WG-Übergängen, nur 2,2 dB Verlust bei mir. Bei einem zweiten 10m langen Aufbau wurden 0.1 dB mehr gemessen. Dies wurde amateurmäßig mit einfachen Messgeräten gemessen.

Wenn man über die Kosten, die Verfügbarkeit des Materials und die Einfachheit des Zusammenbaus dieses Wellenleiters nachdenkt, kann diese Lösung wirklich bedenkenwert sein.

Der Übergang von Koax auf Wellenleiter

Dies ist der einzige etwas komplizierte und kriti-

sche Teil dieses Wellenleiters. Der Übergang besteht aus folgenden Teilen:

- Ein kurzes 15 bis 25 cm langes Stück des 21mm Kupferrohres
- Eine 2mm starke Messingplatte 16 x 16 mm
- Eine SMA-Flanschbuchse mit 4 Löchern zum Festschrauben, Innenleiter 1mm Durchmesser und mind. 9.5mm lang
- Eine dünne runde Messingplatte mit 22mm Durchmesser

Beide Enden des Rohres müssen sauber senkrecht abgesägt sein und die Kante glatt gefeilt werden. Dann wird ein 3 mm Loch exakt 12,2 mm von einem Ende des Rohres gebohrt.

In die Messingplatte für die SMA-Buchse wird in die Mitte ein 3mm Loch gebohrt sowie vier Löcher mit Gewinde für M2,5 Schrauben. Mit einer kurzen M3 Edelstahlschraube und Mutter wird die Platte am Rohr fixiert. Nun wird die Platte angelötet. Dabei aufpassen, dass kein Lot in die vier Löcher für die M2.5 Schrauben läuft. Nun die M3-Schraube entfernen und das Loch auf 4mm vergrößern und entgraten und glatt schleifen.

Das Teflon, das um den Innenleiter der SMA-Buchse liegt wird nun so beschnitten, dass es genau bündig mit der Innenwand des Rohres verläuft und nicht mehr übersteht.

Der nach innen ragende Innenleiter der Buchse sollte mindestens 6,5 mm überstehen, besser sollten es 7mm sein, um Optimieren zu können.

Nun die SMA-Buchse mit den vier kurzen M2.5 Schrauben befestigen. Möglichst Edelstahl-Schrauben verwenden.

Um eine gute Optimierung des Übergangs zu erreichen, benötigt man eine 10 GHz Signalquelle, einen Richtkoppler, eine Spektrum Analyzer oder ähnliche Indikatormöglichkeit. Falls keine Meßmöglichkeit vorhanden ist, sollte der Innenleiter auf exakt 6.5mm gekürzt werden. Damit sollte man dicht genug am Optimum liegen. (siehe Abb.)

Der Wert gilt natürlich nur, wenn der Abstand exakt 12,2mm vom Rohrende beträgt. Die optimale Länge des Innenleiters hängt von seiner Entfernung vom Rohrende ab!

Für eine richtige Justierung muss eine Art Dummy Load an das andere Ende des Wellenleiters angeschlossen werden, da dieser im offenen Zustand ein ziemlich schlechtes SWR von etwa 2:1 hat.

Ein Stück hölzerner Besenstiel, der auf 7cm Länge dem Rohr angepasst wurde und in den Wellenleiter eingeführt wird, kann die Dummy Load Funktion erfüllen. Wenn das Ende der Einkoppelseite des Wellenleiters offen ist, beträgt der Return Loss etwa 10 dB.

Führt man eine Metallplatte (z.B. einen Abschnitt PCB) an dieses offene Ende heran, geht der Return Loss einige mm bevor man das Rohr erreicht, auf gut über 20 dB.

Nun mit einer Feile den Innenleiter ein kleines Stück kürzen. Hierbei keine Zange benutzen, man könnte damit zuviel abschneiden.

Nach jedem Kürzen des Innenleiters rückt das Maximum des Return Loss immer näher an das Rohrende heran. Nach einigen Kürzungen sollte das Maximum dann erreicht sein, wenn die Platte fest ans Rohr gepresst wird. Es sollten wenigstens 25 dB erreicht werden. Bei mir wurden knapp 30dB erreicht. Die Genauigkeit der Messung hängt von der Qualität des Richtkopplers ab, eine niedrigere Auskopplung wird genauere Ergebnisse liefern.

Nach der Optimierung wird eine dünne Schicht Lot auf das Ende des Rohres aufgebracht und die Messingscheibe aufgelötet. Falls dieses Stück im Freien eingesetzt werden soll muss dieser Abschluß wasserdicht sein. Anschließend noch mal den Return Loss überprüfen. Es sollte der zuvor eingestellte optimale Wert sein.

Nun wird eine gerade Rohrkupplung auf das offene Ende so gelötet, dass das Rohr zentriert verläuft.

Am Koax-Wellenleiter-Übergang, der sich im Shack befindet, werden mit einer feinen Metallsäge einige Schnitte längs auf der Rohrseite gegenüber der Einkopplung gemacht. Dieses Ende wird nämlich dort mit einer Rohrschelle gegenüber dem ankommenden Wellenleiter befestigt, so dass die Polarisation eingestellt bzw. optimiert werden kann.

Die oben beschriebene Technik der Justierung der Länge des Innenleiters wurde auch für die Optimierung des 10 GHz Horns nach W2IMU benutzt, das wahrscheinlich das beste Horn für Spiegel mit einem f/D von 0,5 bis 0,8 ist.

In diesem Fall braucht man auch keine Dummy Load im Wellenleiter: Man benutzt das Horn, das während der Messungen aber in möglichst offenen reflektionsfreien Raum zeigen sollte.

Zusammenbau und Justierung des Wellenleiters

Überraschenderweise ist der Zusammenbau der Teile des Wellenleiters unkritisch. Ratsam ist es, das Innere der Rohre vor dem Zusammenbau zu säubern.

In den geraden Kupplungsstücken sollen die Enden des Wellenleiterrohres gerade und sauber einander berühren. Bei den gewinkelten Verbindungsstücken sollen die Enden des Rohres so weit wie möglich hineingedrückt werden. Am einfachsten ist es, möglichst viel des ganzen Wellenleitersystems auf dem Boden liegend zusammen zu löten. Es ist recht schwierig, die Verbindungen hoch in die Luft vernünftig zu verlöten. Am besten lötet man mit einem Gasbrenner. Alle Verbindungen müssen wasserdicht sein. Säurehaltiges Lotzinn sollte nicht verwendet werden.

W1GHZ empfiehlt keine Winkel von mehr als 45 Grad zu verwenden, um beste Ergebnisse zu erhalten.

Nach dem kompletten Zusammenbau wird im Shack ein Signal eingespeist und der Output am anderen Ende gemessen, das mit einer 50 Ohm-Last und einem Detektor abgeschlossen wird.

Nun wird der Koax-Wellenleiter-Übergang im Shack langsam gedreht, bis der Output am anderen Ende maximal ist. Diese Position wird dann im Shack mit einer Schelle fixiert.

Praktische Anwendung

Es gibt viele praktische Anwendungsmöglichkeiten für derartige runde Wellenleiter. Sie können nicht nur für eine Tropo-Anlage, sondern auch bei EME eingesetzt werden. Empfindliche Teile der Anlage wie der Transverter können im Shack bleiben, PA und LNA können nahe der Antenne zusammen mit dem Relais montiert werden. Wegen des geringen Verlustes des Wellenleiters braucht man weniger Treiberleistung und auch der LNA braucht nicht so viel Verstärkung zu haben.

Ich selbst benutze zwei solche Wellenleiter für meine Tropo-Anlage, da mein Transverter getrennte Ausgänge für TX und RX hat. So brauche ich nur ein Koaxrelais.

Mit nur einem Wellenleiter werden zwei weitere Koaxrelais benötigt. Deren Einsparung deckt die Kosten für den zweiten Wellenleiter mehr als ab.

Wenn man für Tropo eine PA mit einer TWT benutzt, kann diese im Shack verbleiben. Die 3 dB Verlust sind hinnehmbar, wenn man die Probleme bedenkt, die das HV-Netzteil bereiten kann.

Zum weiteren Experimentieren bietet sich Rohr mit einem etwas größerem Durchmesser von 35mm an, das für 2-Band-Wellenleiter, nämlich für 3cm und 6cm benutzt werden kann. Jedoch müssen dann die Koax-Wellenleiter-Übergänge neu dimensioniert werden. Details findet man in dem Artikel von W1GHZ, der in den Referenzen genannt ist. Koax-Wellenleiter-Übergänge für jedes einzelne Band, könnten dann ggf. bei richtigen Winkeln realisiert werden.

Verwendete Meßgeräte

Koax-Wellenleiter-Übergänge: HP Signalgenerator, HP141 Spektrumanalysator, 20dB Richtkoppler.

Wellenleiter, Messung der Verluste: 200 mW Signal aus dem Transverter; Dioden-Detektor, geeicht mit 1dB SMA-Dämpfungsgliedern, 20dB Richtkoppler; 50-Ohm-SMA-Mikrowellen-Dummy-load.

Dieses Meß-Equipment ist sicher nicht der Stand der Technik, für Amateurzwecke wurde es aber als ausreichend empfunden, nachdem die gemessenen Ergebnisse durch einige QSOs bestätigt werden konnten.

New Software UKW Tools

Jonathan Naylor, G4KLX

The purpose of this software project is to provide a set of tools which allow for microwave operators and mobile radio implementors to be able to survey locations. It grew out of my need for a Linux implementation of the G4JNT GEOG suite of BASIC programs for DOS. This work well under DOSEMU under Linux but they are limited to just the UK. By its nature (it is based on UK-only NGRs) it is not easy to adapt it to cover outside of the UK.

UKW Tools from version 0.3.0 can make use of the SRTM and GTOPO30 data files that are publicly available. The SRTM (Space Shuttle Radar Topography Mission) data covers most of the world at approximately 90m intervals but has inaccuracies, the GTOPO30 data covers the whole world at approximately 900m intervals and is accurate. UKW Tools can use both sets of data either separately or together in order to provide the best data possible, without paying for it. Despite the imperfections, it is still very useful for creating good terrain models for mapping anything from 2m FM repeater coverage to optical paths for high microwave or optical operation. Or in my case, to help me choose a new QTH which should be good for anything above 136 kHz.

The SRTM data was mentioned on page 73 of DUBUS 4/2003 in the extreme right hand column. The UKW Tools software is available from www.qsl.net/g4klx/software.htm as UKW Tools (just underneath LinWSJT) and it runs under both Linux and Windows. The bad news is that you have to download a lot of data from NASA to run it. Western Europe fills two CDs easily, so a broadband connection is a must to get it.

Here are a few screen captures to see part of the program in operation.



Fig. 1

The first picture is the main screen of UKW Tools and this is the main control centre for the rest of the program. Here you can see a pseudo-3D map of a part of Derbyshire, the county where I live. The area on the right gives information about what is happening on the map. From here the other functions can be started. You can see that there is a location in the centre of the map marked as Alport Heights.

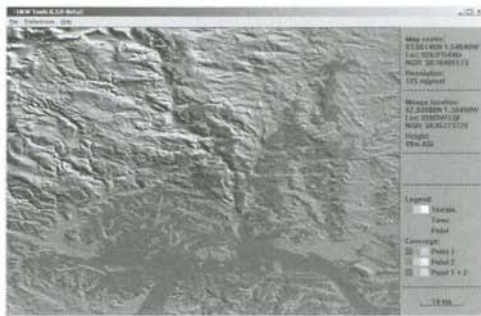


Fig. 2

The second picture is a line of sight coverage plot from the location marked as Alport Heights. This is a major communications site and pictures of it can be seen on my pictures page. The green area (here almost black and mostly in the lower half of the picture) represents the coverage from the top of the mast to any structure which is 8m above ground level. I have also included some local towns on the map. UKW Tools can plot two such sites simultaneously so that overlapping coverage can be seen.

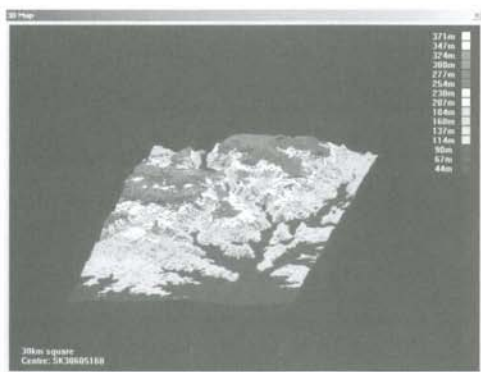


Fig. 3

The third picture is a 3D plot of a 30km square centred on Alport Heights. The palette may be changed as can the viewing angle and other parameters. (The original picture is with colors).

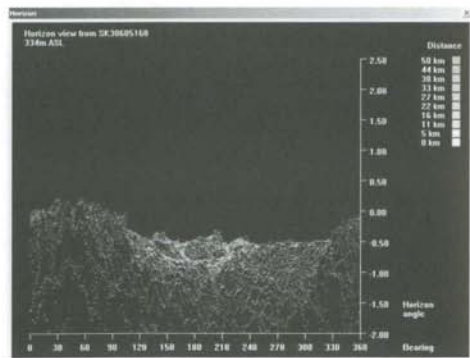


Fig. 4

The fourth picture is the horizon view from Alport Heights up to a distance of 100kms. As you can see, in most directions it has a sub zero elevation angle and it is an excellent location. That is why there are seven communications masts there including the Home Office and more mundane things like the GB3DY 70cms amateur repeater on RB10.

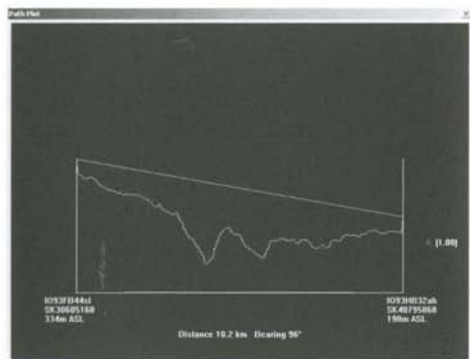


Fig. 5

The fifth picture is of a path plot from Alport Heights to the mast of Zycomm Electronics in Ripley. It is not a great distance but it does show the quality of the SRTM data. At some point I would like to compare this plot with the one generated by a commercial package which we have at work. The first large valley from the left is where Ambergate is located, this is where my first house was, as you can see it was not a good location for anything above 136kHz.

Like LinWSJT, UKW Tools requires the wxWindows library. Unlike LinWSJT it is very portable to Windows etc. I have already had a go at porting UKW Tools 0.3.0 to Windows with success. You will need a copy of MinGW and MSYS as well as the Windows version of wxWindows which is

named wxMSW. The versions that I used were MinGW 3.1.0-1, MSYS 1.0.9 and wxMSW 2.4.2. Apart from some changes in the GUI, it all works fine. In fact all of the screenshots were made under Windows 2000.

The latest version of UKW Tools is 0.3.0 which represents a massive update on 0.2.3 and is probably the first release that looks and feels like something more than the G4JNT GEOG suite ported to Linux. There is a moderately detailed README file within the package that explains how to use it.

The Zip file ending with WinBin is a compiled copy of UKW Tools for Windows 2000 and XP machines. It contains no form of installer at present but I hope to include one soon, however it does allow Windows users to try UKW Tools without having to install a complete development environment. It even runs under Linux using Wine. It is slightly out of date compared to the main release but this will be rectified in due course.

Deutsche Zusammenfassung:

G4KLX hat eine Software "UKW Tools" im Januar 2004 in der Version 0.3.0 veröffentlicht. Sie ermöglicht Landkarten mit Flächen der optischen Sicht von einem QTH aus, Geländeschnitte usw. grafisch darzustellen. Dazu werden Daten, die von der NASA frei erhältlich sind (SRTM, Radarbilder der Erde vom Space Shuttle) verwendet. Diese haben eine Auflösung von etwa 90m, aber einige Ungenauigkeiten.

Sie sind für die ganze Welt verfügbar. Ausserdem können GTOPO30-Daten verwendet werden, die genau sind und aus 900m-Abschnitten bestehen. Das Programm kann beide Datenquellen gleichzeitig auswerten und so die Genauigkeit verbessern. UKW Tools läuft unter Windows und Linux und kann kostenlos im Internet downgeloadet werden unter: www.qsl.net/g4klx/software.htm Für den Download der SRTM Daten benötigt man eine schnelle Internetverbindung bzw. Flatrate, da alleine die Daten für Europa ca. zwei CDs füllen. Das Programm bzw. die Info-Texte dazu sind nur in Englisch verfügbar, weshalb sich eine komplette Übersetzung des Artikels erübrigt. (DL8HCZ)

Microwave Parts

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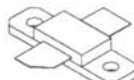
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MAXIMUM RATINGS

Rating	Symbol	Value	Unit
Drain-Source Voltage	V_{DS}	65	Vdc
Gate-Source Voltage	V_{GS}	±20	Vdc
Total Device Dissipation @ $T_C = 25^\circ\text{C}$ Derate above 25°C	P_D	87.5 0.5	Watts W/ $^\circ\text{C}$
Storage Temperature Range	T_{stg}	-65 to +150	$^\circ\text{C}$
Operating Junction Temperature	T_J	200	$^\circ\text{C}$

THERMAL CHARACTERISTICS

Characteristic	Symbol	Max	Unit
Thermal Resistance, Junction to Case	$R_{\theta JC}$	2.0	$^\circ\text{C/W}$

ELECTRICAL CHARACTERISTICS ($T_C = 25^\circ\text{C}$ unless otherwise noted)

Characteristic	Symbol	Min	Typ	Max	Unit
OFF CHARACTERISTICS					
Drain-Source Breakdown Voltage ($V_{GS} = 0$, $I_D = 10 \mu\text{Adc}$)	$V_{(BR)DSS}$	65	—	—	Vdc
Zero Gate Voltage Drain Current ($V_{DS} = 20 \text{ Vdc}$, $V_{GS} = 0$)	I_{DSS}	—	—	1.0	μAdc
Gate-Source Leakage Current ($V_{GS} = 20 \text{ Vdc}$, $V_{DS} = 0$)	I_{GSS}	—	—	10	μAdc

NOTE - CAUTION - MOS devices are susceptible to damage from electrostatic charge. Reasonable precautions in handling and packaging MOS devices should be observed.



TYPICAL CHARACTERISTICS

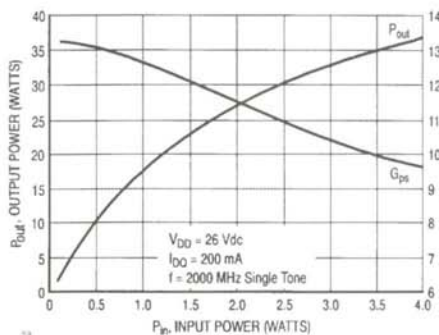


Figure 5. Output Power & Power Gain versus Input Power

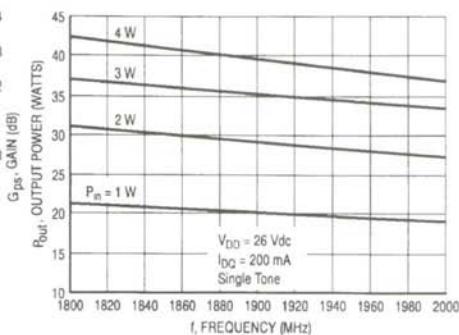


Figure 6. Output Power versus Frequency

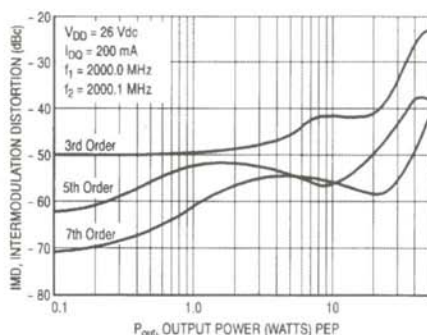


Figure 7. Intermodulation Distortion Products versus Output Power

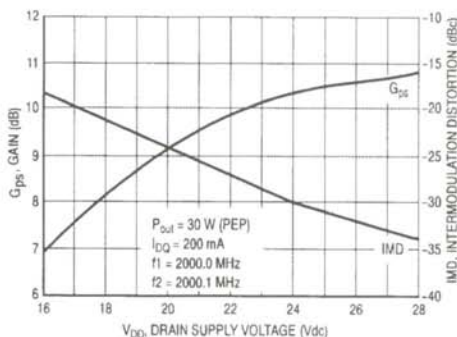


Figure 8. Power Gain and Intermodulation Distortion versus Supply Voltage

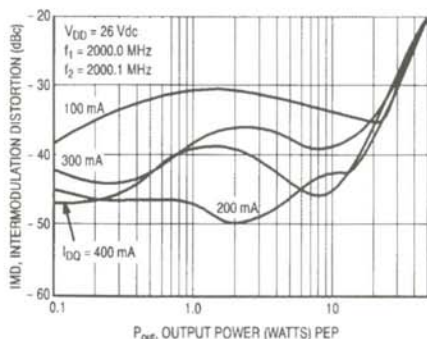


Figure 9. Intermodulation Distortion versus Output Power

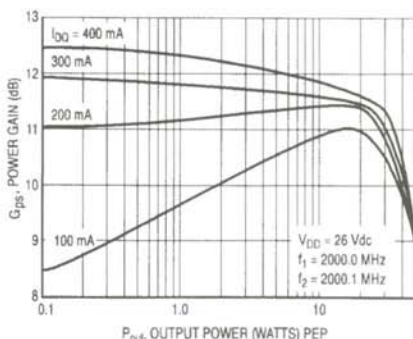


Figure 10. Power Gain versus Output Power

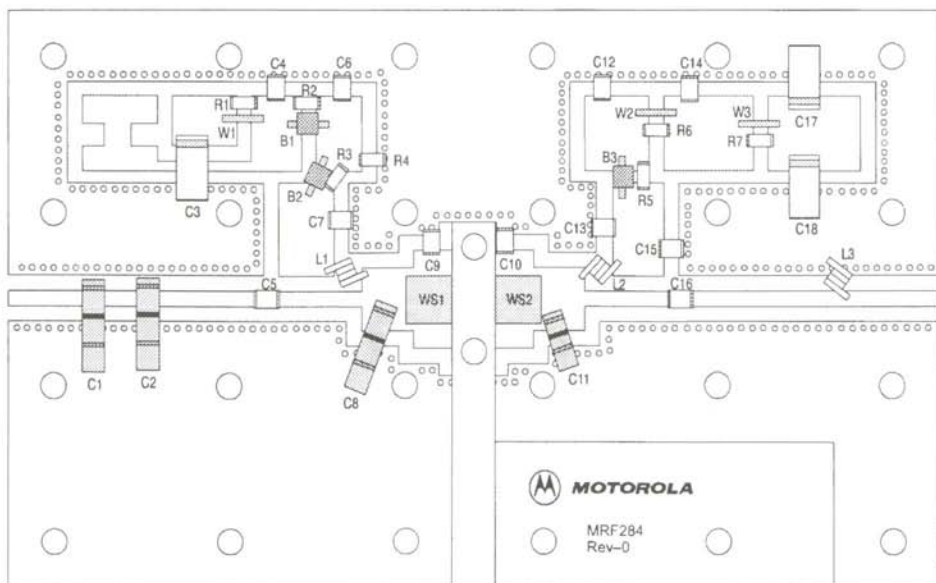


Figure 2. 1.93–2.0 GHz Broadband Test Circuit Component Layout

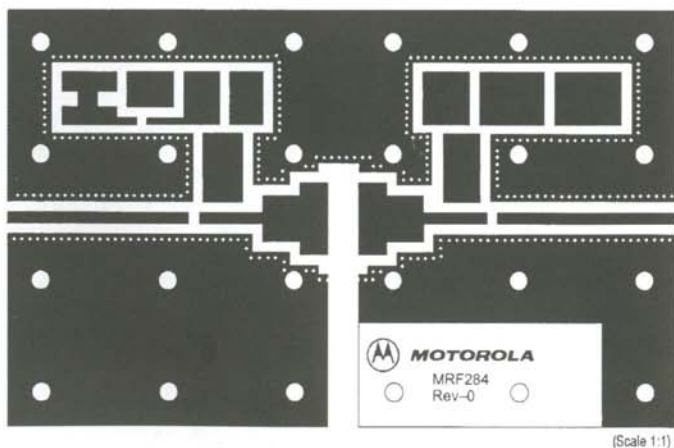


Figure 3. MRF284 Test Circuit Photomaster
(Reduced 18% in printed data book, DL110/D)



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13cm: MT2,3ZQ250WA 2-3 W in, 250 W out, -ask-

9cm: MT3,4ZQ250WA 2-3 W in, 250 W out, -ask-

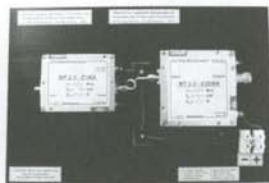
6cm: MT5,7D130WA, 13 W in, 130 W out, -ask-

3cm: MT10ZD50WA, 2,5 W in, 50 W out, -ask-

MT1,3O160W on heatsink
with two fans

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Microwave Europe

Simon Lewis, GM4PLM/DL4PLM

Hello and welcome to another column. Winter has us in its grip again as I write this and it is certainly not a time to be portable! The hills just a few miles to the south of my QTH in Scotland are deep in snow and in the lowlands here in Düsseldorf it is very cold and quite a lot of snow this weekend. But this type of weather certainly reminds us that we should be extremely vigilant and prepared should we go out portable even on much lower ground. If you are one of the adventurers that do brave the weather please ensure you are well prepared. The news regularly carries details of unwary people that get into difficulties due to them being ill prepared. Please don't become a news item! Of course sitting in front of the fire on the farm sends the mind into warmer weather and the thoughts of spring. Make this year different, build for a new band, operate in some of the contest and activity periods, but above all, get out and enjoy the bands.

Please note! Again little news received from Europe! Come on – let know what is happening in EU so we can advise everyone!

Until next time – 73 Simon DL4PLM/GM4PLM

Amateurs Receive Mars Express!

Using what he described as "just a quick throw-together" system, Charlie Suckling, G3WDG, has received a signal in the UK from the European Space Agency's Mars Express spacecraft. Unfortunately the Beagle 2 lander it carried was lost on landing during the Christmas period but Charlie reports he heard the Mars Express signal on X band (8.4 GHz) on December 9 using a 3 metre diameter dish. In a message to James Miller, G3RUH, who provided him with advice on setting up his equipment, Charlie said his system noise factor was about 1 dB, and he used Miller's S-Band 2.25-turn helix scaled to 8.4 GHz as the feed (LHCP). Signals seemed very consistent for about two hours, the signal level was very approximately 0 dB S/N in 2.5 kHz. G3WDG also adds that it

was not too hard to locate the signal, about 10 minutes of searching plus or minus 100kHz and tweaking his azimuth and elevation settings. In mid-November a team of German amateurs were able to copy the Mars Express signal from a far more sophisticated setup in Bochum, Germany, using a 20 meter parabolic antenna. Reception



G3WDG 8 GHz Receiver and Feed

of the Mars Express signal provided a test run for the facility, which will serve as the ground control station for AMSAT-DL's Phase P5-A Mars orbital mission planned for 2007. AMSAT-DL President Peter Gültow, DB2OS, said it was the first time ever that a signal of an interplanetary deep-space probe was received in Germany. It was probably also the first time ever that such a signal was received by Amateur Radio operators he added. You can find a full report on the AMSAT-DL Web site.

WG 24/47GHz switches available surplus!

Peter G3PHO reports "At long last a plentiful supply of waveguide switches for both 24GHz and 47GHz has found its way to the amateur surplus market, the RelComm Technology model RDW SR002 three port waveguide switch. It's in WR28 (which is the same as the British WG22). I have been using a manual WR28 four port switch for several years and have had no problems using it with WR42 (WG20) butt-jointed directly against it. All that's needed is, with a needle file, to carefully elongate two diagonally opposite holes on the WR42 flanges so that screws can be fitted through the flange into the threads on the waveguide switch. The RelComm switch is electrically operated by means of a short pulse of 8.2V DC. In practice this means if you are running from a +12V portable supply you'll need to drop the supply by means of a 1 amp regulator to around 8 or 9 volts. A 3 pin 7808 or 7809 will do the trick. The six pin connector shown here has multicoloured wires coming from it. Changeover switching connections are as follows:

GREY: Ground(0 VOLTS)

RED: +8.2V DC switched

BROWN: +8.2V switched

To change the relay from RX to TX merely apply the 8.2V to either the Grey or the Red lead. To switch back to RX apply the 8.2V to the other lead, using Grey as the common ground. The remaining wires at the connector are from reed relays that can be used to control other modules in your system (eg switch on a PA, etc) or to control indicator LEDs to show when you are in TX or RX mode. You can easily work out these connections using your ohmmeter. By the way, when we first got hold of our relays we applied 9V from a small battery to the Red lead. Sure, the switch went over with a nice thud but it wouldn't go back when we moved the battery to the Brown lead! Reason ... the new 9V battery was flattened! This relay needs a high current pulse for it to latch. The little battery hadn't got the "guts" to work a second time!

These WR28(WG22) switches are available from the Wellington VHF Group, PO Box 12 259, Thorndon, Wellington, New Zealand. Email Leon Toorenberg, ZL2AOC at: leon@surveylab.co.nz



He will give you a quote on price. When we bought a number of these switches a couple of months ago the price was \$50 New Zealand plus postage to the UK. It worked out at around £23 (UK pounds Sterling), including post, for one unit. That's around \$47 US a fantastic bargain! Leon has a UK bank account as well as his usual NZ one. We arranged to pay into the UK account and so avoid any bank charges, etc. Those of you outside the UK may need to enquire at your bank about transfer of funds or merely send Leon US dollars. Check with him first to get an updated quote. In January 2004, Leon told me he had lots of these switches for sale!! Specs from the manufacturer: 26.5-40GHz (they work fine on 24 AND 47GHz and are tested!). Insertion Loss: 0.25dB Maximum Isolation: 50dB minimum Connectors: WR28 (UK WG22)

News Microwave Items

DB6NT has been busy over the past few months and has introduced a number of new items on his sales lists. An 8.4 GHz converter and preamp are now available allowing you at home to have a listen for the Mars probes like G3WDG. The IF is at 1270 MHz. For those more affluent of you, a new 10 GHz solid state PA is now available with an output of 50 Watts this will certainly get you heard on the band! But be prepared to sell the wife! Also a new range of MOSFET amplifiers for 23cms has been released and are well worth looking at. You can find details of all the latest items at the DB6NT website.

Microwave Organisation Changes in UK

Not much news yet on the new arrangements for the UK Microwave Group (UKuG) and the details on how they will support the UK microwave community. These details are still being discussed with the RSGB and the only news is that the society have agreed that the UKuG can provide the support of the 2004 contest calendar and the programme for this has now been agreed. You can find more information on the webpage of the Microwave Newsletter editor, Peter Day G3PHO. The group will also take charge in running of the Microwave News - previously an RSGB publication. This will now pass to the UKuG who will run and support the publication of this very popular newsletter.

The contest calendar for 2004 (until July) for the UKuG is as follows. The rules follow the contest calendar. All dates shown are Sundays, except where specifically indicated. Detailed rules can be found on the web site of Peter Day G3PHO.

O = Open, R= Restricted section, WB = Wideband

Date	UTC	Contest
March 28	0900-2100	Low Band uWave Contest (1.3G/2.3G/3.4GHz) O
April 4	0900-2100	1. 24 GHz Cumulative O
	0900-2100	1. 47 GHz Cumulative O
May 1 (Sat)	1400-2200	RSGB 10 Ghz Trophy O
May 9	0900-2100	2. 24 GHz Cumulative O
May 9	0900-2100	2. 47 GHz Cumulative O, WB, R
May 23	0900-2100	1. 5.7 GHz Cululative O, WB, R
May 23	1400-2200	1. 10 GHz Cumulative O, WB, R
June 6	0900-2100	3. 24 GHz Cumulative O
June 6	0900-2100	3. 47 GHz Cumulative O
June 20	0900-2100	2. 10 GHz Cumulative O, WB, R
June 20	0900-2100	2. 5.7 GHz Cumulative O, WB, R

Interesting AC Scatter Pictures

Recently talking to Nick GM4OGI about AC scatter from GM to JO31 he informed me of some interesting pictures found on the Internet. I guess the reproduction of this picture in printed form will not do it justice but this picture highlights how AC scatter can be such a useful mode on uwave frequencies! This image shows a myriad of contrails from AC over Southern G and N France.



Portugal: ATV on 10 GHz

On 19/8/03 at 16:30utc CT1FP, CT1AMG and CT1BKF made an ATV QSO from IN51NV to IM59HI at a distance of 285.5 Km on 10455MHz with P4. Weather was dry and temperature was 26C. Tx 0.4W / RX 0.65 dB NF with 90cm Dishes both ends. They believe this sets the CT record for ATV distance on 10GHz, and also the Iberic distance record on ATV 10GHz. (tnx to CT1DMK for QSP)



CT1FP working 10 GHz ATV

Links

Peter Day G3PHO - <http://www.g3pho.org.uk/>

DB6NT - <http://www.kuhne-electronic.de/>

Mars Reception / AMSAT-DL - http://www.amsat-dl.org/p5a/reception_g3wdg.htm

UK Microwave Group -
<http://www.microwavers.org>

Microwave Japan

Seiji Fukushima, Ph.D, JH6RTO

On the 13th of January 6-m EME QSOs became legal in Japan. The following four topics are picked up, which happened recently in Japan.

First Six-metre EME from Japan JH6RTO, Seiji

Because the Japanese telecom authority regarded EME communications as the amateur satellite services, the 6-m EME had not been permitted in Japan for many years. Now it is legal in a small segment of 50.0-50.1 MHz. As long as I know, four Japanese amateurs succeeded in 6-m EME QSOs on the first day- January 13th this year, JST. It became effective at 1500Z on the 12th of January. The first-day QSOs reported are:

01/13 JH2COZ worked W7GJ on JT65B
using a single 8-element Yagi

01/13 JE1BMJ worked WA4NJP on JT65?

01/13 JM1SZY worked W7GJ on JT65B

01/13 JA1DXA worked W7GJ on JT65B

JA1DXA is a new call sign of JH7OHF, Katsu, who participated in the dxpeditions of K5K, T32R, and VK9ML. In addition to the first-day QSOs, I received the following logs:

JM1SZY, Akira wkd using 2x10el Yagis and a 1kW amplifier:

01/13 1000Z SM7BAE 6m JT65B

02/08 1038Z WA4NJP 6m JT65B

JH2COZ, Yasuhiro wkd on 2m JT65B:

02/08 KE7NR, K7MAC, P6GWB.

Figures 1 and 2 are JH2COZ, Yasuhiro's 8-el Yagi for 6-m EME and his shack, respectively.

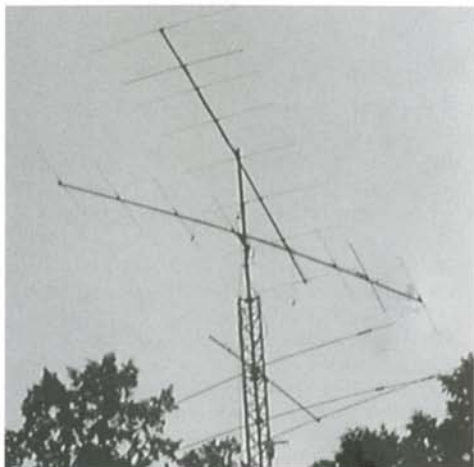


Fig 1.: JH2COZ's Yagi for 6m EME



Fig. 2: Shack from JH2COZ, Yasuhiro

Geminids and Quadrantids MS Summary, JH6RTO, Seiji

Here is a log collection of the major winter meteor showers.

J11FLB, Den (PM95ws) wkd:

2003/12/14 1415 JF3IPR 6m JT6M 27 27 493km
and hrd: JE2XBY on 6m, JF1RYU on 6m JT6M,
and JR8MCR QN02ut 798km on 2m FSK441

J14POR (PM65nk) wkd on 2m FSK441:

2004/01/03 JA7BJP QM07gn 705km
2004/01/03 JM1LQW PM96qa 569km
2004/01/03 J11FLB PM95ws 613km

JA7BJP, Satoru (QM07gn) wkd on 2m FSK441:

2003/12/13 JA9BOH PM86fa 26 26 403km
2003/12/14 1205 JR8MCR QN02ut 26 27 540km

JR8MCR (QN02ut) wkd on 2m FSK441:

2004/01/03 J14POR PM65nk 1104km
2004/01/03 JA9BOH PM86fa 879km

JA9BOH, Kimio (PM86fa) wkd on 2m FSK441:

2003/12/13 2042 JA7BJP QM07gn 26 26 403km
2004/01/03 0145 JR8MCR QN02ut 27 27 879km

JH9TJT (PM86hk) hrd on 2m FSK441:

2003/12/13 2104 JA7BJP QM07gn 370km
2004/01/03 2319 JR8MCR QN02ut 831km

Each mode of the WSJT modes has been being popular in Japan as described in these two topics of EME and MS communications. Many, many thanks to K1JT, Joe.

249 GHz Domestic Distance Record Updates, JA1KVN, Bunpei and colleagues

JA1KVN, Bunpei and JA1ELV, Nori sent me a couple of reports that they updated the Japanese 249GHz dx records.

From Nori, "The first trial was performed on the 22nd of November, 2003 in Ohtone-machi, Saitama Prefecture. Conditions: wx cloudy, temperature 12 degrees C, and, humidity 35%. Time in JST (UTC+9).

14:37 JA1KVN/1-JA1ELV/1 55 47 FM

15:00 JA1KVN/1-JA1ELV/1 539 559 CW

15:25 JA1KVN/1-JA1ELV/1 52 44 SSB

The latitude and longitude of JA1KVN and JA1ELV is N36-09-40-01, E139-40-15-02 and N36-09-09-08, E139-41-28-07, respectively. The distance was measured as 2.06km. My equipment is a homebrew 1-microwatt transverter and a 35-cm dish, as shown in Fig. 3."



Fig. 3: JA1ELV, Nori's transverter

Bunpei reports "The second update was successfully done on the 13th of December, 2003 in Ohira-machi, Tochigi Prefecture. Conditions: wx sunny, temperature 15 C, and humidity 25%.

JA1KVN/1-JA1GLD/1 (N36-20-10, E139-40-55)
59 57 FM 2.79km

JA1KVN/1-JG1MFR/1 (N36-20-10, E139-40-55)
52 52 FM 2.79km

JA1KVN/1-JG1MFR/1 (N36-19-28, E139-39-07)
58 54 FM 4.7km

During our trial, JA1KVN's QTH is always the same, N36-21-23-08, E139-42-01-05 at Mt. Ohira.

Figure 4 is a transmitter used at our experiment. My colleagues used the similar ones. The final diodes are anti-parallel 6th-order harmonic mixers MA4E2039 available from M/A-COM and the antennas are 45-cm TDK with homebrew feeders. The output power is estimated to be -30dBm. In conclusion, the Japanese dx record on 249GHz is 4.7km now." Congratulations! (Editor)



Fig. 4: JA1KVN, Bunpei's tvtr and antenna

Simple, one-board 5.7GHz Transverter - JA0DFR, Hisakazu

This is a brief report of my simple 5.7GHz transverter. A lot of people could reproduce the same one because of its extreme simplicity. You could understand quite easily that this circuitry uses only 5 MMICs as key components. Let's begin with the transmission part at the right of Fig. 5. The final amplifier is just two cascaded NEC MMIC NE3210S01s that output 10 or 12 dBm. So this transverter is really QRPP if you use this as it is.

Looking down from the TX stage, you will see the two cascaded Hittite MMIC HMC320MS8Gs as a RX amplifier. The IC is a narrow-band one while you can use it without any adjustment on 5.7 GHz. Finally, the mixer is on the left hand. It is a Mini-Circuits MBA-591L with LO/RF of 4950-5900MHz and IF of 0-1000MHz. When a LO power is 4 dBm, its conversion loss is 9 dBm. Due to the wide coverage of IF and LO/RF, 144-1296-MHz bands are acceptable as the IF transceiver.

Some characteristics are summarized: voltage 13.5 V, currents 50 (RX) and 110 (TX) mA, output power 10dBm, case size 40x68x10 mm only! The entire transverter is shown Fig. 6. The LO is not included so an adequate one should be prepared separately. The LO power recommendation is 5-10 dBm for IF of 144 and 432 MHz but you need 10 dBm for 1296 MHz because 1296 MHz exceeds the specification of the mixer IC.

References: www.minicircuits.com and www.hittite.com

Editor's address: S. Fukushima, 1182-2506 Hase, Atsugi 243-0036 Japan. Email jh6rto@m.ieice.org
URL <http://www.geocities.com/ah0r/>

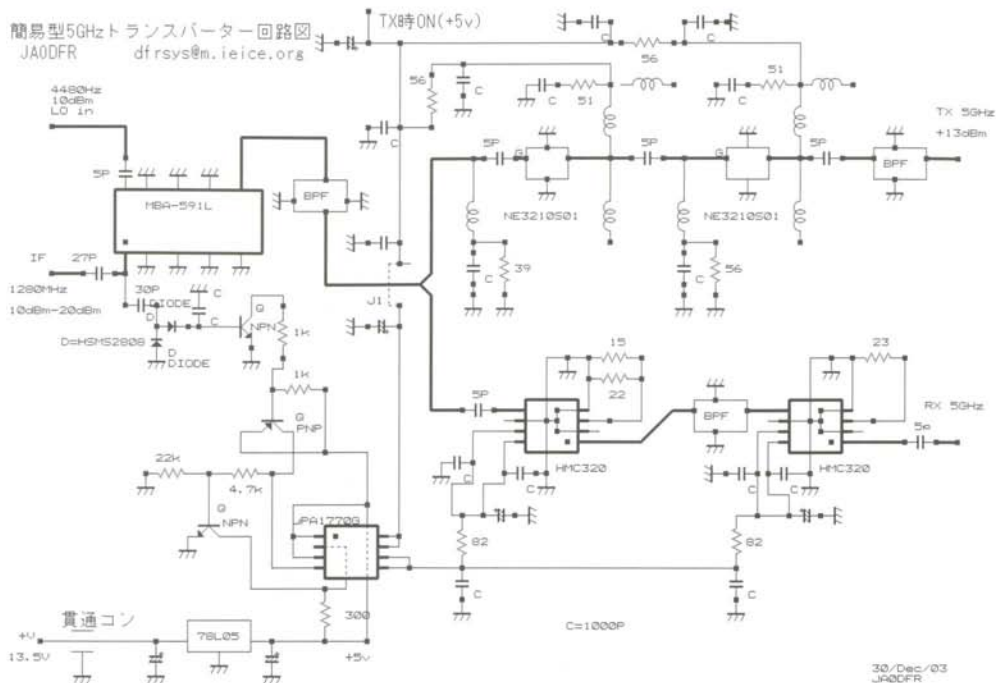
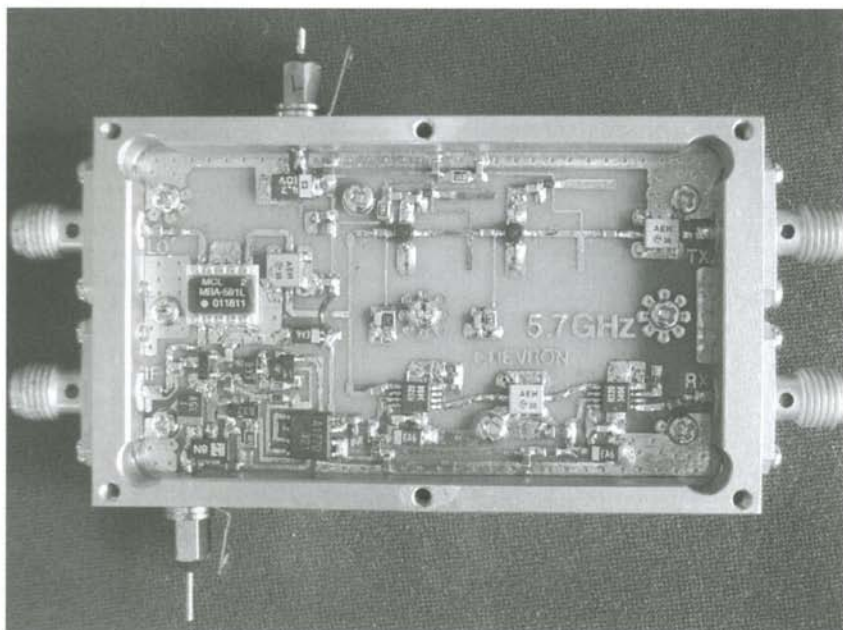


Fig. 6: Photo of the transverter



Microwave USA

Kent Britain, WA5VJB

1st 241 GHz VUCC

and new 241 GHz world record

On their 4th attempt, WA1ZMS/4 worked W2SZ/4 17 Feb on 241 GHz. This QSO earned Brian his 5th grid for the ARRL VUCC award and a new world 241 GHz record of 79.6 km.

A technique commonly used by the VLF crowd on 76 KHz was used for this QSO, QRSS CW. With a direct synthesizer fed by a 5 MHz Ultra Low Noise OCXO from Wenzel, they were able to keep the 241 GHz CW signals in 1 Hz filters for DSP processing. While some of the QSO could be copied by ear, most of the signals were observed on the computer. Dots were about 1 second long, Dashes 3 seconds, with each station sending the exchange several times taking well over an hour to complete the QSO.

The equipment was 30 cm dishes (53 dBi) and -1 dBm from a tripled 80.666 GHz Oscillator.

The weather varied from -8C, Humidity 53%, Dewpoint -16C, Calculated loss .68 dB/km at the WA1ZMS's end, with -8C, Humidity 22%, Dewpoint -26C and Calculated loss of .27 dB/km at the W2SZ end of the QSO. Absolutely amazing work.



WA1ZMS's 30cm dish for 241 GHz



Pair of dishes



WA1ZMS's Cards
for the 1st 241 GHz VUCC



241 GHz station

As expected (see last issue's column p. 56) the 241 GHz world record was extended already before to 61.8km on Dec. 3rd 2003 at 0148z by a QSO between W2SZ/4 (WA1ZMS op) FM07fm and W4WWQ/4 EM97xe. The weather at the time of the QSO was -6.1C Dew Point: -17C Relative Humidity: 40% Station pressure: 876mb These WX conditions result in a total atmospheric loss of 0.541dB per km.

Microwave Update 2004

I would like to invite everyone to Microwave Update 2004 Conference October 14th to the 16th. I will be co-hosting the conference with W5LUA at the Harvey Hotel DFW. The Harvey is just a few km from the North Entrance of the Dallas-Fort Worth International Airport. Lots of good speakers and our world famous "Surplus" tour. For more information please check: www.microwaveupdate.org

FAI News

Reports, Info, Theory
Editor: Dom Dehays, F6DRO
f6dro@aol.com

FAI SEASON 2004

2003 was not very good for FAI. In my opinion, it was not because of conditions, but because of low activity. Very few openings were reported, and more surprising, almost nothing from the usual SCP around JN36, which usually provides the vast majority of the FAI QSOs.

Furthermore I received very few reports from the operators who worked stations via FAI. Thank you anyway to those who did.

What can we do to improve the situation?

Main problem is to get warned that FAI conditions are OK, and to get infos in real time of the scatterer position. This can easily be done using the tools we get already: Cluster or/and ON4KST web site.

Another problem is that for a scatterer to be located, we must try it.

I suggest that the well equipped and familiar with FAI stations run daily skeds, in case a sked works: please spot the info on the cluster network.

These skeds should be run, using the places known to be the most often working: like JN36/JO10, ect...

We could publish the known skeds in this column, so that everyone is aware and can try to listen on them. So please send me your informations, as soon as possible, in case you decide to try to run skeds.

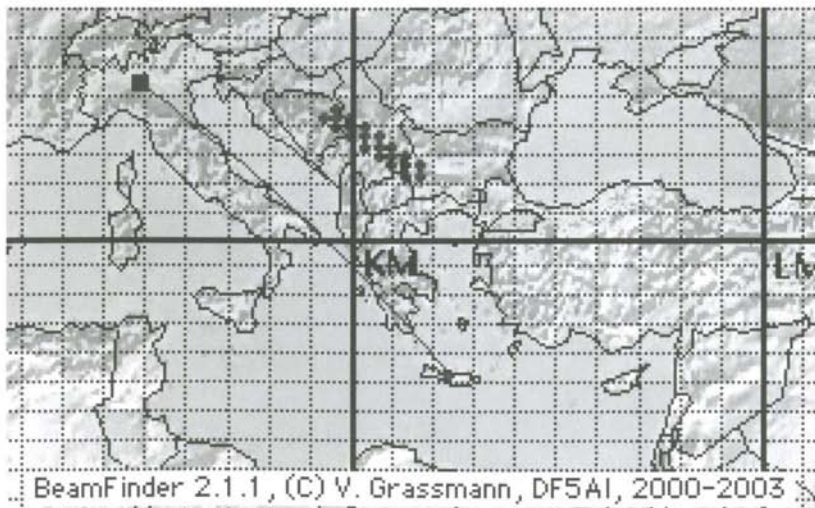
Let's hope this, will help to increase activity. Don't forget to send me your reports for any FAI qso during the next season (May to September).

FAI on the Edge

Northern Italy to Crete

Volker Grassmann, DF5AI

In this column, FAI opportunities are shown which might encourage radio amateurs to conduct unusual dx propagation tests.



FAI scatter volumes (dot marks) corresponding to the path I - SV9.

Please contact the author at feedback@df5ai.net or, alternatively, visit the www.df5ai.net webpage if a detailed analysis of antenna headings, target areas and scatter volumes is required in serious propagation tests.

2-m-EME News

Bernd Mischlewski, DF2ZC

e-mail: BerndDF2ZC@t-online.de

Stan, **OK1MS (JO70)** is more than happy with his DXCC. In November 2003 he received the award he had been working towards for so many many years. Congratulations! Stan's award is no. 14. During last year he worked quite a number of new stations (all on cw):

N9LR, SP2OFW, K0GU, W4SW, N1KI, PA0OOM, IK1UWL, JF4TGO/8, WW8M, DF2CQ, W0EKZ, W4ZRZ, OH1FF, DF0BV, SV3KV, KL6M, PA2DWH, SM3XZF, RN6BN, 5T6M (the first QSO on 2m OK/5T), DF7KF, IW5CNS, SK6EI, DK3EE, W4AD, PE1LWT and F8CXO #840. In ARRL EME contest OK1MS scored 78x36 (all CW). Stan is looking forward to skeds via e-mail: ok1ms@seznam.cz



OK1MS (JO70SL): 8 x 15 ele. DJ9BV (mod.)

Jac, **PA3DZL (JO21)** worked in January (all JT65B worked with his 4x17 el array):

06 Jan
21:18 EA3DXU JN11CM
22:08 KB8RQ EM79SV
07 Jan
22:08 OE3FVU JN78VE # 296
22:42 S52LM JN65TX
08 Jan
00:10 KB8RQ EM79SV
00:34 WA8CLT EN80LA
12 Jan

07:59 K7XQ CM97QI # 297

26 Jan
10:30 KG6DX QK23KL # 298, DXCC
13:38 UA9SL LO71NS
29 Jan
22:59 K7MAC DN13TU
23:33 DK8ZJ JO54AG
30 Jan
00:02 W8WN EM77BQ
15:38 ES6RQ KO28WA
23:41 KB8RQ EM79SV (wkd with only 100 W out)

Joop, **PA0JMV (JO21)** sent an impressive list of stations worked with his 2 yagi setup: "Hi friends, stations with a * behind the call are new initials. The degradation figures are from ZT3 by N1BUG (path loss and bg-noise)

03 Jan
NQ20* JT65b (2x17/1k) -4.8dB
10 Jan
N7WNX* JT65b (2x9/1k) -1.8dB
11 Jan
K7MDL* JT65b (2x10/150w) -1.9dB
SP7DCS* CW (4x21/QRO) -1.9dB
VK7MO* JT65b (1x10/400w) -1.9dB
24 Jan
VK2KRR* JT65b (4x13/400w) -2.4dB
25 Jan
F6FHP JT44 (4x6w/qro) -2.6dB
IK7EZJ JT65b (4x7/qro) -2.6dB
SV1BTR* CW (16x6/qro) -2.6dB

26 Jan
OH1JCS JT65b (2x8/150w) -3.2dB (14 minutes qso!)

30 Jan ZS6DDG* JT65b (1x17el/200w)

(Derek is using 35m coax, no preamp. After 5 x NC skeds it was now pretty easy, even at degr. of -4.8dB). So it seems to me that JT65b really does its job when both stations are performing technically correct (freq drift < 10Hz and low phase noise). AND: there is NO need to be active ONLY when the predicted DEGR figures are below -1dB. VERY small stations can be worked as well when the forecast doesn't look excellent. Working successful EME is being ACTIVE ANY TIME, paying attention to every 0.1dB you can improve, and hope for luck :-). More results and pictures of some worked small stations on my homepage: www.planet.nl/~pa0jmv

Adrian, **YO4FRJ (KN34)** managed to find few more days of 2m and 70cm cw EME operation in January 2004. On 2m, first station found and worked in 2004, W5UN (RO/O), was on random at January 4th at 23:00: big signal from Dave, like a "local" station. On January 9th, at 21:00 Adrian worked with SM7WSJ (O/RO) on sked. Hakan's signal reports were more than +10dB over noise with his new antenna system. Next morning, on January 10th, YO4FRJ worked on random with W3EME (O/RO) and WB9UWA (O/RO) from 03:30 to 03:59. Mike's signals were +20dB and Jim's signals "only" +16dB over noise. Same morning conclude the sked with WA4MVI (RO/O) at 04:30 in good conditions. Later in the evening, during the next moon pass, worked I2FAK (RO/O) on random very strong like usual at 20:26 and failed the sked with SP7DCS. However, Adrian found mixed conditions during that

day. On January 11th, he noticed very good conditions, despite some ionosphere perturbations reports higher in latitude in Western Europe. Average reported echoes were at +15dB over noise most of that time. He worked K9MRI (RO/O) at 04:51 on random very fast. Joe's signal was very strong with more than +15dB over noise. Then, Adrian finished on sked with SP7DCS (RO/O) at 05:00 very easy. Chris did a great job with +15dB over noise. At 06:18 DK8ZJ (RO/O) was worked on random with +10dB over noise signal, followed by PE1LWT at 06:28 with "only" 2 x Xpol antenna system and Linrad. Jürgen's signals were affected by some QSB, but with a strong +12dB over noise average. At 06:41 worked on random N5BLZ With (559/O) close to Adrian's local Moonset. Ed's signals were reported very strong (more than 20dB over noise) affected by some QSB, despite the local noise at YO4FRJ's QTH in that direction. The same day, after the next moonrise, Adrian worked on random very fast PA0V (RO/O) at 21:08. Nanko's signals were very strong (peak at +20dB over noise but also affected by strong QSB). Soon after that, failed to the sked with SM3BYA. NIL on both sides. Skeds are welcome to: yo4frj@xnet.ro

RK3FG wkd with JT65 in Jan 2004: K1JT and K7MAC.

UY2QQ is a new EME station from KN77 with 4x 12 el LY and GS35.

Less EME Activity? by WB9UWA

There have been comments on Moon-net about less EME activity. There are many reasons for a decrease in 2 EME activity. Here is my top 10 list.

- 10) Lower world ham population.
- 9) Worse world economy.
- 8) Restrictive antenna installation rules.
- 7) Fewer hams becoming VHFers.
- 6) Fewer VHFers exploring EME.
- 5) Current EMEers exploring more bands.
- 4) Current EMEers getting too old to do antenna maintenance.
- 3) CW forever EMEers getting bent out of shape over new digital modes.
- 2) Many EME "first" have already been done.
- 1) Poor EME conditions.

In this article, I will explore the last reason; "Poor EME conditions". There is much that could be said about the role of the Sun and EME conditions. I won't say them here because I am not an expert. There are several factors that affect conditions. Most of them are predictable.

- 1) Path loss (Perogee or Apogee; There is a 2db difference.).
- 2) Sky temperature.
- 3) Sun noise (New moon or overactive sun).
- 4) Local noise levels (line noise, carriers and other broadband noise).
- 5) Effects from Solar flares. (I'll leave that one general on purpose).

I first became QRV on 2M EME in late 1999. I started with a very small array, so I was not well in tune with the band. By late 2000, I had a bigger, but moderate sized array. Most recently I have been noticing less performance from my array even after switching to Xpol. It seemed like in the "good old days" (being 2001), perogee would nearly always be at minimum sky temperature. Now it is not the case.

I decided to do a limited study on sky noise vs Perogee, vs time. Using Z-track as my resource, I took a snapshot of only the day of Perogee, over a range of the years 1996 to 2011. Please note that often only one day away from perogee, there is a substantial change in sky temperature, but the trends are still there. Also this study does not account for new moon, whether the active side of the sun faces Earth, or what the sky temperature is at Apogee or some other time of the month. This study does not factor in declination either. Perhaps someone would take it upon themselves to add these other factors and reach some interesting conclusions.

So to be clear, I am looking at the sky temperature, on one day each month (Perogee). Instead of giving you the raw data here, I made some observations about each year.

Max = Maximum sky temperature in db at perogee in the year.

AVG = Average sky temperature in db (add all temps and divide by 12)

Min = Minimum sky temperature in db at perogee in the year.

> 2 = The number of months containing perogees that have a sky temperature of 2db or less.

The year 2001 was an exceptional year. We will not have another one quite like it until 2010. Do not sell the station yet. 2006 will be nearly as good as 2001 was. If we can get through 2004, 2005 should be a better year. You can see by the data that there are cycles, but they are not very regular.

You can also make a case in that in the current year, there will be many low degradation days around 2db.

This is true, but they will be at or near Apogee. The Apogee, Perogee issue relates to actual signal path loss. These are measurable hard numbers that show on your S-meter. The sky noise is something that can become less meaningful if your local noise level is high. You can see why there is so much confusion about when to be QRV. here are many weekends in each month that have moderate overall degradation, but no single weekend tends to stand out as the obvious choice. It is in these times especially that we need to all do our best to observe times of Activity Weekend and Activity Nite.

Jim Shaffer, WB9UWA

Year	96	97	98	99	00	01	02	03	04	05	06	07	08	09	10	11
Max	10	3.8	3	4.3	4.3	2.7	2.5	4.3	10.6	10	2.8	3	4.4	4.5	2.1	3.2
Avg	5.5	2.1	2	3	3	1	1.6	3	7	4	1.8	2.1	3.1	2.7	1	1.8
Min	2.4	1.2	1.2	2.1	.6	.2	.6	2.1	3.2	1.6	1.3	1.4	2.4	.7	.3	6
> 2	0	6	5	0	2	11	6	0	0	2	8	7	0	4	11	6

EME News

Editor: Bernd Wilde, DL7APV

Intro

The 11th International EME Conference will be held on August 6, 7 and 8, 2004, at The College of New Jersey.

Papers on VHF-UHF and Microwave techniques will be presented, and they will be of interest to a large number of radio amateurs.

Registration for the Conference is now open. Papers are also being accepted for the Proceedings.

Please visit: www.qsl.net/eme2004

or email at: info@eme2004.com 73, Marc N2UO

DUBUS/REF EME Contest, 2nd part March 27/28

Digital

432.044 appears to be calling freq.

Watch/monitor 432.030 - 432.050 MHz digital modes freq for JT44 and 65.

After the ARRL contest many microwave EMEers see the must of a new contest or change of a well known contest.

At one or two weekends operating six bands is a big problem. Changing feed at bad weather or be on the right band at wrong time are the common problems as well as the "three Band division of 13cm" 2304/2320/2400 MHz.

Some proposals:

ZS6AXT

So let's get the ideas for this contest from as many of you, as possible. In my humble opinion there should be no more than two bands per weekend. Repeated contacts should be allowed at least after say 24 hours, to increase activity with low numbers of stations. QRO and QRP categories should be introduced. I hope that we get somebody who will organize such contest. I am prepared to do the logs evaluation, if necessary. So - sharpen your pen and give your ideas! 73, Ivo

F2TU

For ARRL contest weekend, I have to change 10 times of sources for the 23, 13, 6, 3 cms... The European contest is already a better solution. For making well, one would need 3 weekends:

1- 144/1296

2- 432/2300

3 - 5760/ 10 GHz with 1 day for each.

Who will manage to make modify this contest?

73 Philippe

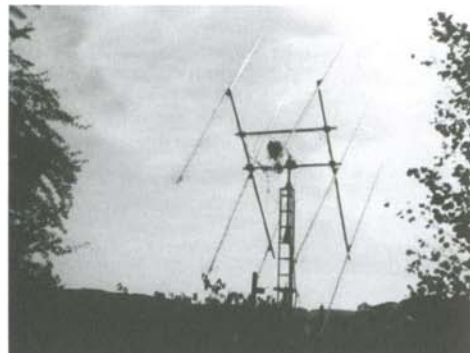
WA6PY

I agree with the need of separate Microwave EME contest. In my opinion contest should start from

2.3 GHz and UP. 2.3 GHz band is difficult and probably this is the reason why activity on this band is low. 1296 today is like 144 MHz 20 years ago and I would not include this band. We had discussions regarding separate Microwave EME Contest for many years now, but we should just agree date and rules and go. VY 73 Paul

OK1CA

I agree with the separate Microwave EME Contest or the microwave weekend, the part of the ARRL EME Contest. I propose on day 2,3 GHz, the second day 5,6 GHz and the both days 10 GHz. Franta



DK8VS 6x9wl at 432 MHz

432 MHz

EA3DXU

Activity very low in Feb 7/8. I was fortunate to complete with 1 new #, and the first ever 2Y / 1Y EME QSO on 432 Mhz band working OK1DFC on JT65B. Feb 8 OK1DFC O/O JT65B sked on his Moonset # 160. 2x38 el M2 + GS23B; 73 de Josep

HB9JAW

Here a small report and comment to 2003 ARRL EME Contest from first Weekend:

By the start off the contest elevation readout had a water problem, no way of fixing it at 2 in the morning outside. The next pass 70cm HV supply and Amp broke, 2m preamp blown up and 23cm much down in performance. A absolute disaster. Only worked JW/SM2BYA #, this manually without readout...hi

Second weekend:

On 432mc worked following new station:

SM5IOT #, OK2BDO #, GW3XYW #, N2IQ #, KU4F #, KE2T/4, DK8VS #

On 1296mc I worked to following new stations:

KA0Y #, W7BBM #, K2UYH #, W7SZ #, OE5EYM #, DL4MUP # JH5LUZ #, JA8IAD #. Just before the end of the contest UR5LX called, but he faded away. To weak. Score: 43 Qso x 27 multi

HB9BBD helped me improving my RX. Before the second weekend, he tuned my IMU horn to the best possible. Now signals and echos are much stronger then before but there is still a lot to learn and improve on 1296mc for me. All the best and see you soon. 73 de Michel HB9JAW

KE2N

Now that the dust has settled, here are my results on 432.

I concur with other observations that the first weekend was fair and the second was poor. Murphy intervened Saturday of the first weekend with a vaporized coax connector out on the tower. I could hear the solar flare on the Sunday of the first weekend - it raised my noise floor!

First leg : DL9KR, HB9Q, F6KHM, N2IQ, OH2PO, K5GW, JW/SM2BYA, N9AB, WA4NJP, K1FO, K4EME, UA3PTW, KU4F, SM2CEW, OZ4MM, K2UYH, JA6AHB, second leg: HB9JAW.

It seems like I only hear/worked the big guns.

Ken (16x28 el, coax fed, 0.33 dB nf, 0.8 dB feedline loss 1500 watts output)

PA2CHR

EME Contest, Nov. 15/16:

During the 2nd. part of the ARRL contest I made: 6 QSO's with 6 multipliers on 70 cm. Conditions where like the first part: up and down, lots of polarisation chance. New stations where: KU4F and N9AB on 70 cm. Total score is 70 cm: 16 QSO's with 14 multipliers. I was also very happy to work FY5DG on Nov. 19 for DXCC # 43 on 70 cm. Rig 70 cm: FT847, MGF1302 pre-amp, nf: 0.32 dB, YL1050, 4 x 35 el.BVO/11wl. Hm, gain: 23.0 dBd. Good DX to all and best 73's, Chris, PA2CHR

PA3DZL

Wkd 2003: HB9Q, F6KHM # 126, JW/SM2BYA # 127 new DXCC, DL9KR, HB9Q.

2004: EA3DXU O/O JT65b # 128, KL6M O/O CW # 129. The QSO with Josep, EA3DXU was my first 2Y to 2Y QSO, UFB signals !! Also my first QSO on 70cm in JT65b. JT65b is more difficult on 70cm then on 2m due to doppler and freq.drift 73's de Jac, Ant.: 2 x 21 el. 6.6L Yagis

OK1DFC

I made QSO with DL9KR 559/559 UFB signal from Jan and with OZ4MM 539/549 also UFB signal from Stig. Sorry I am only during Moonset and Moonrise QRV but it is possible to do skeds. Partly QSO JT65B with EA3DXU OOO and received RO, but no final RRR, moon below horizon. Hope to hear some of you, too. Regards, Zdenek. 1.5kW / single 38 el M2 / 0.34dB N/F, CW/JT65B

QRP EME on 432?

Yes, if you do have a station running 4 yagis and 50W, 2 yagis and 100W, or 1 yagi and 200W you

are capable working HB9Q on CW random!

HB9Q (JN47cg) is QRV since April 2000 on 432 and 1296 EME. They are also QRV on 144 EME since 1983.

On 432 MHz they have worked so far 235 initials, 46 DXCC, 34 maidenhead, 197 grid square, 23 states of USA and WAC July 2002.

On 432 they are using the following station: 15.28m dish, f/d: 0.53, cross-dipole RX: switchable H/V, TX: circular 7/8" hard-line, GaAs FET 0.40 dB NF tuned by HB9BBD, 8874 by HB9SV, 1000W, SSB Electronics LT 70 MKII into IC-756 (modified) NF Filter: DSP 599zx, Additional visual DSP by AF9Y (FFT DSP)

In summer 2003 they worked on fine-tuning their 432 station getting very good results. During the 2003 ARRL EME Contest they have worked 119 stations in 40 multipliers. All QSO's were on random CW. Several stations were QRP!

This are a few examples of QRP stations worked recently:

Dec. 2003				
Call	Ant.	P/w	Mode	Comm.
JJ3JHP	4x21el	50	CW	
DL7UDA	4x21el	70	CW	PA was QRT!
SK6EI	2x28el	100	CW	first EME QSO
A71AW	3m dish	100	CW	first EME QSO
Jan. 2004				
VK3BRZ	4x15el	120	CW	first EME QSO
I1PIK	4x31el	120	CW	first EME QSO
Feb. 2004				
PE1ITR	2x28el	180	CW	
VK7MO	1x17el	200	JT44	first EME QSO
VK3HZ	1x21el	200	JT44	1. EME QSO
DG1KJG	4x27el	200	CW	
ES6RQ		200	CW	
GW4DGU	2x4m	220	CW	
DL7UAE	2x9el	300	JT65B	
DJ8MS	4x13el	300	CW	
PA5DD	1x28el	800	CW	
JA9BOH	2x2.8m	1000	CW	
OK1DFC	1x38el	1000	CW	

Satellite stations and good tropo and contest stations have a fair chance to work them. They are also QRV JT44/JT65B. Interested? Contact them at dan@hb9q.ch. For more info visit www.hb9q.ch

Expeditons

The 3B9C exp. will be qrv from March 16th about 4 weeks with 70cm EME 2 yagi + 100w.

432 & up (net) news

DL4KG

I just want to inform the 432 MHz EME community that a very heavy wind gale tilted my tower and destroyed my open feeder system. Fortunately the antennas survived. I was able to repair the tower

but temperatures and snow did not allow to repair the feeder system. I will be QRT until further notice. 73 de Gerald www.qsl.net/dl4kg

DL8OBU

DL8OBU is qrt on 70cms EME.

KL7FH

Father and son team KL7FH and KL9A are QRV on 432 FT847 100W 16x12el w/polarity rotation ready for first QSO. Anyone interested in trying? -Chris KL9A and Frank KL7FH, Grid BP51BB kl7fh@gci.net

LU7DZ

is fully operational on 432 now with four good yagis, 1500 watts and hearing sun noise. He is actively looking for skeds now. His E-mail lu7dz@yahoo.com.ar

N5ITO

I have new 736-R coming soon and I have a 100 watt amp for 432. Antenna's are Cushcraft DX-420 Colinear arrays. 2X2 with AZ/EL 80 elements total. Old testings shows the DX-420 has around 12.5 dBd so I figure I'm running around 16-17Dbb. Running 7/8 hardline to the antenna. 73 Dave

PE1LWT

Will start EME with 2 BVO70-8.5 antenna using as a xpol yagi next spring. 73's Jurgen

SV1BTR

In the meantime with the help of my good friend Marko LZ2US my new 70cm array is up for some time now, with 8*39jxx. 13wl. No power yet on that band but hopefully soon. Soon I will be posted abroad (Malta or Philippines) for 1-2 years so will be coming back to operate only in contests and in a few weekends with high declination, 73 Jimmy

DK3WG writes: I am the QSL Mgr for more than 25 Russian/Ukrainian EME-stn. I answer QSL requests direct and also via bureau.

Recently I checked the situation of 2002. I received QSL cards for only 30% of the initials! If QSL- card supply drops further down I will stop the service.

The most of the Russians told me, that they do not receive (and mostly do not send) cards via bureau! We have a lot of post-robbers, I am posting to Russia/Ukraine only "Registered-letters" and check it. Even also in this mode letters from time to time are lost and I must ask for new cards.

I am missing answer for my letters with QSL cards from 24 initials:

9A2AE, DK1PZ, DK4JN, EA3PL, F5JJI, F8SQ, HA5BDJ, I5MPK, K5GW, K8ISK, KD0GS, N2HLT, N6OVP, W0RRY/5, W8MQW, WA5ETV,

OE5EYM, OK1DIG, OM1TL, OK3DQ-now OM3DQ, VE1BVL/1-now VE1ZJ, ZL2AQE.

Best 73, see you via the moon. Jurg, DK3WG

Active Italian EME stations

432 MHz	Loc.	Ant.	Pwr/W	qrz
I0UGB	JN62BO	3mt.dish	800	N
IK0BZY	JN61GW	1x33BV	600W	Y
IK0EQJ	JN61FO	4x26el	1K	Y
IK1HWG	JN44DS	4x21FT mod	700	n
IK1MTZ	JN35UB	8x26el	3CX800	Y
I22FOB	JN45MQ	4x19FT	300	Y
IN3AGI	JN56	8x32el.	TH338	Y
IN3KLQ	JN56RG	2x21FT	600	Y
IK4DCO	JN64GB	4x26el.,	1,5K	N
IK4DCX	JN64GB	4x26el.,	1,5K	N
I5CTE	JN53OS	4x26el	650	Y
I5MPK	JN53JU	32x27el.,	500	Y
IK5WJD	JN53PS	4x25elBV	8877	Y
I25EME	JN52NS	8x25el.,	1K	Y
I6PNN	JN72BL	4x26el.,	700	N
1296 MHz				
I0UGB	JN62BO	5,08 mt.	TH308	Y
I21BPN		4 mt.dish	300	Y
IK2MMB	JN45PQ	3,5mt. D	TH308	Y
IK2RTI	JN55AD	4,8mt.D	230	Y
IK3COJ	JN65BL	3mt.dish	250	Y
I6PNN	JN72BL	4 mt.dish	400	N
2304				
IK2RTI	JN55AD	4,8mt.dish	130	Y
I6PNN	JN72BL	4 mt.dish	140	Y
5,7GHz				
IK2RTI	JN55AD	4,8mt.dish	25	Y
I6PNN	JN72BL	4 mt.dish	28	Y
10GHz				
IK2RTI	JN55AD	4,8mt.dish	20	y
IW4BTJ	JN54PN	4mt.dish	45	Y
Team:	I4CHY	I4BER	I4JED	
I4TTZ	JN54KI	7 mt.dish	55	Y
Team	I4TMA	I4BER	I4ZAU	
I4QIG	IW4APQ	IW4BYP	I4ZBEH	
IQ4DF	JN54VJ	7 mt.dish	70	y
I5PPE & IK5WJD	JN53NS	3 mt. dish	20	y

Info de I1ANP, Mario



UT3LL & his new dish

23cm

UT3LL

I am very glad that in second part ARRL-contest 2003 I have made my first EME QSOs on 1296 MHz. See the photos of his 1296 MHz new dish and 432+2400MHz antennas. 73 Valery

OE9XXI

On November 15/16 during the ARRL EME Contest I achieved a final score of **79x31** on 1296 and 1x1 on 2300. Initial contacts on 1296 were SK0UX (549/549), VA7MM (549/559) 2kHz above my echo, UR5LX (O/O) for DXCC 51 and IZ1BPN (449/559). Many times I have had heavy QRM from a neighbour's TV SAT LNC radiating IF signals, I sent several QRZs to UT3LL but I realized his callsign too late and he disappeared. I sent also several QRZs to GJ3Q? (marginal signal). Concerning initial counting I found following multiple callsigns from my logs and QSL cards of the last few years:

DD0SB = DL2SBY

DJ5MN = DL5MAE / DJ3MY / DK5MV / DF4MAA / (?)DO6ME (Oct 14, 01)

DL0ZAB = DL4DTU / DL1HRH / DL1HTT / DL3HPT / DL3HXS

DL6YDH = (?)DH00AH (Oct 21, 00)

F5VHX = F/G8MBI

F6ETI = F8COZ

IY4FGM = IY4ARI

Other callsigns appearing only once are (?)DC3SAW (Apr 20, 02) and (?)DM3UH (Apr 20, 02), possibly that are also guest OPs of any stations. The (?) indicates that it is supposed, all others are confirmed.

Since about a year an advanced cancer with painful skeleton metastases hampers many of my activities. From medical therapeutic view, recovery looks not promising, but don't be worry - I will not give up! Despite this problem, I had a lot of nice QSOs on Feb 7 during the SSB Contest on 1296 included two way SSB contacts with VA7MM (43/56) and VE7BBG (43/56) for a score of 37 (18+1). On Feb 8/9 unfortunately I missed the moved sked with EA9/DL3OCH but worked him on Feb 13 as ZB2/DL3OCH (O/O) on JT44 for an initial and new DXCC. 73's Peter OE9PMJ

HB9BBB

Wkd in ARRL contest:

18.10.03

F1ANH, JH5LUZ, IK2MMB, HA5SHF, OE9XXI, LA8LF, OM6AA #195, OZ6OL, G4CCH, OZ4MM, JA6CZD, OE5EYM, JA6AHB, DF3RU, F6CGJ, OH2DG, G3LTF, SM3AKW, HB9Q, OK1CAQ, IK3COJ, DL4MUP, F5VHX, N2UO, DJ5MN, DL3HRT, W9IIX, DK0ZAB, W2DRZ, K5JL,

W7BBM, F2TU, W7SZ, WA6PY, K5GW, I0UGB, K2UYH, PA0SSB# 196, ON4QQ# 197, K7XQ, DL1YMK, WA4NJP, JA8ERE, HB9JAW, ON5RR, JH1EFA.

19.10.03, OE9ERC, HB9SV, JH1KRC# 198, JR4ZZS# 199, ZS6AXT, F6KHM, G4DZU, G3LQR, PA3CSG, DL3HXS, W2UHI, W7GBI.

15.11.03

JA8IA, UT3LL#200, DL8OBU, IK2RTI, K9BCT, VE6TA, K0YW, SM2CEW, VE7BBG#201, WA4OFS, SK0UX#202, VK5MC, UR5LX#203, OE5JFL,

16.11.03

9H1ES, JF3HUE, DJ9YW

Final score: **75x31**. 73, Dominique

VA7MM will be QRV on 1296 MHz. Station is comprised of 3m dish, 200 W into circularly polarized feed, 0.4dB receive NF. GPS frequency reference will be deployed. 73, Mark, VE7CMK
VA7MM Moon Bounce Team, Coquitlam, British Columbia, CN89OG, e-mail: va7mm@rac.ca
OPs, M. Mattila, VE7CMK, ve7cmk@hotmail.com
Toby Haynes, VE7CNF.: haynes@npsnet.com

UR5LX is active 1296 EME. He wkd in Jan 2004:

9 Jan G4CCH, 10 Jan G4CCH, OE9XI, 11 Jan HB9SV#, GW3XYW# **first QSO GW-UT**, OZ6OL#, JA6AHB# **first JA-UT QSO**, 12 Jan IK2MMB # 19 **first QSO I-UT**. Loc.: KO70wk, 3,2 m dish, 150 w FHX35, 73l Sergej. <http://www.ur5lx.vhf-dx.net>
Results from the **R/U-stn** Jan 2004:

DL0SHF/DK7LJ

Is now QRV on 23cm. We got it! If you work hard enough, you get it. The 750+W PA is right behind the feedpoint of the 9m dish. Echoes are smashing. Absolutely noise free in FM. Per made the first QSOs. 73 de Christoph

DL3OCH's EA9 and ZB2 DXpeditions

I had a lot of fun during the dxpeditions, although it was a very long drive of 2300 km from Weingarten to Gibraltar. Everything seemed to work fine until I tried to make some EME QSOs on 8 Feb from Ceuta (EA9/DL3OCH, IM75HV). I set up the antenna and tried to QSO DJ9YW. The "Guardia Civil" (local police) showed up and wanted to see my permission. They took my passport and my German license and checked everything. I started to operate, but they became upset because I had to start my car engine to have enough power. They thought I was pulling their leg when I pointed my antenna to the moon. It was very hard to explain. They did not speak English or German and my Spanish is just enough to order something in a restaurant. We did not have to wait long for more police to come to see what was going on. After about 30 minutes there were 9 policemen. Fortunately I was able to complete with Heinrich (FIRST DL-EA9). Then the boss arrived and I had to turn off everything and take down the antenna. Some of these guys wanted us to stop and some of

them thought that it was ok. They nearly had a fistfight. The wrong group won, and I had to go to the police department. After a while, somebody who seemed to know something about radio came. He looked at my equipment and asked me about the frequencies that I used. They all came to the conclusion that I could continue. That was good, but the moon was gone. I tried to rearrange the skeds and I was able to work OE9ERC and K2UYH the next day (Feb 9th). The WX turned very bad and it was not possible to make any other skeds. I had to wait 2 extra days because the ferry was not running as a result of the storm. We eventually could make it to Gibraltar. There, somebody from the intelligence services showed up, but it was easy to explain what I was doing and he left. At times I had problems with bad QRM on the sked frequency, but was able to move some of the skeds around and completed (ZB2/DL3OCH, IM76HC) 3 QSOs with DJ9YW, OE9XXI and OE9ERC, which were FIRST ZB-DL and ZB-OE. I was also briefly QRV on EME from Spain (EA7, IM86) where I QSO'd DJ9YW, OE9XXI and OE9ERC. I made 5300 QSOs on HF and now have to fill out lots of QSL cards. More trips are planned for the future. Heinrich is developing a circular yagi that should give me 3 dB more gain. I hope to use this new antenna next summer. Special thanks goes to DJ9YW and OE9ERC for their support of the dxpeditions and to Andreas who assisted with the driving and setup. See Bodo's webpage at www.qsl.net/dl3och/zb2.htm for additional pictures and details.

Equipment: IC-706 + Transverter (by DJ9YW), PWR: 83 Watt at dipole, ant: 59 element, 19.7dB



**ZB2/DL3OCH's 1296 EME
operating position in Gibraltar (IM76)**

2320MHz

OH6NVQ

I am intersted to try a sked. Did some echotesting, I can hear my echoes quite well now! Only qrv 2320 MHz. 150w @feed and 3m dish. 73's Tomas
oh6nvq@sral.fi

EME-TOPLIST

Call	Sqr	DX	US-	WAC	Initials
Small	CC	States	WAS	#	
144 MHz					
I2FAK	540	64	122	50 1988	1993 1241
HB9Q	476	66	112	50 1986	1994 954
PA2CHR	338	55	79	50 1993	2002 526
CT1DMK	154	0	52	0 --	-- 160
DM2BHG	140	26	40	28 --	-- 221
YO2AMU	132	18	44	20 2001	-- 167
OK1DIG	125	28	37	25 --	-- 155
RW3PF	99	23	26	13 --	-- 103
DD0VF	92	23	33	19 --	-- 101
OK1MS	?	58	103	50 1982	1992 812
432 MHz					
DL9KR	391	53	83	50 1987	1978 768
DJ6MB	310	43	63	44 1984	-- 368
OK1KIR	254	45	63	44 1982	-- 359
DL7APV	174	32	38	42 1984	-- 228
OE9ERC	148	30	36	28 1993	-- 226
HB9Q	143	31	37	25 2002	-- 194
UA3PTW	103	26	30	22 2002	-- 117
S52CW	92	27	29	22 --	-- 110
UT3LL	78	26	21	20 --	-- 84
OK1CA	71	24	26	22 1996	-- 121
CT1DMK	61	0	21	0 --	-- 64
PA2CHR	54	20	22	13 --	-- 65
OK1DFC	39	17	19	11 --	-- 42
OK2BDQ	29	14	16	7 --	-- 32
1296 MHz					
OE9ERC	168	32	40	30 1993	-- 268
OK1KIR	148	32	39	28 1989	-- 190
DJ9YW	132	29	40	23 1998	-- 154
HB9Q	102	25	29	20 2001	-- 129
CT1DMK	96	0	24	0 --	-- 102
OK1DFC	78	23	32	15 1999	-- 98
OK1CA	76	26	26	19 --	-- 108
OK1UWA	12	10	11	5 --	-- 22
UT3LL	8	6	6	1 --	-- 8
2300 MHz					
OE9ERC	54	16	25	9 --	-- 58
OK1KIR	39	13	19	10 --	-- 41
OK1CA	19	11	15	2 --	-- 20
OK1UWA	11	7	10	1 --	-- 12
3400 MHz					
OK1CA	1	1	1	1 --	-- 1
5760 MHz					
OE9ERC	23	12	16	4 --	-- 25
OK1KIR	16	10	12	3 --	-- 17
CT1DMK	15	0	10	0 --	-- 16
10368 MHz					
OK1UWA	24	10	14	3 --	-- 25
CT1DMK	21	0	11	0 --	-- 22
OK1KIR	19	9	13	2 --	-- 21

OE9ERC	11	8	7	3	--	--	11
I5PPE	9	9	14	3	--	--	26
OK1CA	7	4	5	1	--	--	8



UT3LL 432 MHz & 24 GHz antennas

Software/Internet

1296 MHz 50W amps see :

<http://www.g3wdg.free-online.co.uk/>

It's available as a kit from him.

Nice maps all over the world look at

<http://www.wm7d.net/azproj.shtml>

I did put some schematics + pictures of the Siemens HV convertor I have in use on my site. It's part of the FM1800/11200 amplifier (11.2GHz) which uses a RW 1125 also. Take a look at <http://pa0c.mine.nu>. Click the 10 GHz button and follow the Siemens link on that page. 73 Henk PA0C

Designer SV is available for download from Ansoft. It is the free student version of Design. This software is very capable and can be used to optimize your favorite RF preamp on your favorite band. Circuits can be input to a schematic in much the same way as in PCEExpress. It is designed to run on Windows 2000 or Windows XP. Here is the web site to download it from: <http://www.ansoft.com/downloads.cfm>

Alain, ON4KST has released a new chat channel, especially for EME JT65 (and other modes). Have a look at www.on4kst.info you will find that the chat channel not only offers generic chatting facilities, but also a wealth of other information, and all is user-configurable. 73 - Franz OE3FVU

SETI

Heppenheim, Germany.: First scientific meeting in Europe, in March of 2004.

SETI League members should watch the organization's website at www.setileague.org/seticon for further details on the upcoming meeting, as they become available.

The SETI League and the 11th International Ham

Radio Moonbounce Conference are pleased to announce that they have scheduled Prof. Joseph Taylor, K1JT as featured speaker at their joint SETICon04 Awards Banquet at The College of New Jersey on 7 August 2004.

MARS Express at DL0SHF

It seems to be the first time that amateurs are able to track spacecraft at another planets orbit.

The reception of MARS EXPRESS is very successful at DK7LJ/DL0SHF. Signal level is quite good at 8,4 GHz. Here are some words on the equipment used:

7,5m precision dish with a tracking accuracy of 0.001 degree. 10 GHz LNA at feed horn I followed by a filter for 8,4 GHz and again a 10 GHz amp. Another filter and another amp followed by 60m hardline. Then a 10 GHz designed mixer fed with an HP generator sets the signal to ca 1,5 GHz where it goes into an ICOM receiver, which is normally used in the 23cm EME station.

What do Per DK7LJ and friends get ? The signal is too weak to decode any data. But the center carrier or whatever it is can be received. The ESA Mars Express peaks ca 15 to 20 dB in 2,4 kHz bandwidth with a good amount of doppler shift caused by the Earth's rotation.

I have updated the DL0SHF page, and there is a link to a (german) site where you find screenshots of the spectra, of Per DK7LJ, the antennas and more. Also I have updated the information about DL0SHF.

<http://www.dl0shf.de.vu> (nags you with pop-ups)

<http://www.qsl.net/dl0shf>

<http://www.df9cy.de/dl0shf> (thats were it really is)

73 de Christoph <http://www.df9cy.de>

MARS Signals?

Detecting signals from the Mars Global Surveyor
Doug McArthur (VK3UM)

During the period August 26-29 2003, NASA's Jet Propulsion Laboratory (JPL) in cooperation with the Stanford Research Institute (SRI) will be conducting a test of the Mars Relay transmitter aboard the Mars Global Surveyor (MGS) spacecraft currently in orbit around Mars. The test involves the spacecraft transmitting a 1 watt 437.1 MHz CW carrier for reception by the 46 meter dish at SRI. Amateurs with 70cm EME class stations using DSP techniques may be able to detect this signal also. Well here was a challenge. This is my account of how I went about it and the results achieved.

I should initially explain that I have been actively involved with EME (Earth Moon Earth) for over 20 years and have, after considerable effort and with the help of close friends, erected a 10 metre fully steerable (computer controlled) parabolic reflector. At 437.1 MHz it has a gain of 30.5dBi. The dish surface is fully specified to 7.5 GHz and usable to 10.3GHz. (Kennedy reflector, all aluminium welded construction, installed on a home brew mount)

The first thing I needed to establish was the feasibility of actually detecting the MGS signal. Given the 1 watt transmitter output to a unity gain antennae at a distance of 55,770,000 kms it is fair to say, that even with what you may consider is a large antennae, the signal would not exactly "blow your head off"! In fact the free space loss at that distance is 240.08 dB. Mind you this is some 20 db less than the EME path on 432 MHz but the MGS was not running the

power or antennae gain of my fellow EMERs!

Using my EMECalc software (obtainable from www.qsl.net/vk3bez/vk3bez.htm) it could be seen that given my receiver capabilities, I could only expect to receive MGS at a S/N equivalent to **-1.2dB in a 1 Hz** bandwidth. This was based upon my receive system temperature of 97°K. (degrees Kelvin) In fact this figure was not correct as I failed to account for another 3dB loss due to the space craft radiating right hand circular polarisation. I am only capable of receiving linear horizontal or vertical polarisation. At the time I believed I should be able to detect the signal and pressed ahead to check and confirm my systems capability.

The receive antennae feed system (and used for transmit) consists of two sets of folded dipoles in an H configuration, remote switchable for both horizontal or vertical polarisation. The spacing has been designed to align the radiation pattern at the 10dB point corresponding to the rim of the dish. This provides a reasonable compromise between antennae gain and antennae temperature. The pre amplifier or low noise amplifier (LNA) at the feed consists of a home brew cavity with a FH35 HEMFET. Y factor and noise figure measurements indicate it has a noise figure of better than 0.35dB and with the relay switching and feed losses provides a receiver noise figure of about 0.7 dB or 51.5°K. Given a Sky temperature of 15°K (hopefully Mars would be in the quiet part of the sky) and ground noise (dish spill over) the actual system noise temperature is about 1.25dB or 97.2°K for elevation angles above 4°. The preamplifier at the feed is followed up at the shack end with an identical LNA which then passes through a 3 cascaded precision stepped attenuators. (10, 1.0 and 0.1 dB steps). These attenuators form the key element to my measuring capability. At this point I have two down converters 432-28 MHz and 432-15 MHz. The former I use to feed my DJ9BV automatic noise figure meter whilst the latter feeds a W & G Peglemesser (SPM6 precision tuneable selective voltmeter). This output also feeds sound card and another accurate HP AF voltmeter. The output of the down converters also feed an Icom 751 (receive only) whilst at 432 MHz (switched attenuator output) I also feed the receive section of Icom IC910H. Other test equipment allows for signal generation and frequency measurement when required.

I should explain what Y factor measurements are and how they are used. Y factor measurements (in this case) are simply the difference between two noise source level measurements. One of these sources should be known or the level derived from another source. The Y factor, measured generally in dB, is simply the difference between the two sources. If you know one then you can derive the other. I prefer to use the substitution measurement and you will note that my precision attenuator is placed in the 432 MHz signal path. The reference indicator is the 2.5 dB (full scale) analogue level meter which part of the precision selective voltmeter.

As an example, to measure Sun noise, I would first direct the array at the quiet area of the sky (Aquarius) and note the level meter reading. I would then direct the antennae at the Sun and add attenuation until the previous level is again obtained on the meter. The Y factor is the amount of added attenuation inserted. This value can then be entered in my software and knowing the Solar Flux Index (SFI) at the time you can derive your Receive System noise temperature. The Sun is not always a good choice of a noise source as its level can vary significantly. Those fortunate to have a large array then choosing a known noise source like Cygnus A is a better and a more accurate noise source. Sagittarius is not a good choice neither as it is not a point source which may impact adversely on your measurement. You can see that by using the substitution

method you do not require a heap of expensive test equipment. There is so much you can achieve by the simple substitution method and I get some enjoyment in deriving the SFI from my measurements and then checking the results off the web. It is rare to be more than 10 SFI units out and if there is a discrepancy you have a problem Houston!

If all of the above seems just too complicated and the figures difficult to understand then down load the software and you will find that by substituting your figures all the hard work and calculations are done for you. More over the software provides for "reverse engineering" calculations to derive other unknowns. It is better than the average game programs as it has a large number of buttons and variables you can play with even if you don't know what you are doing!!

Back to checking out the system. Given that my normal operation is at 432 and the MGS is transmitting on 437.1 MHz it was essential to know just how much degradation may occur at the higher frequency. By using the Sun Noise measurement procedure, as described above, the system gain was found to be 10 dB less (due to fixed tuning in inherent with the down converter). The Y factor however was within 0.2 dB and totally acceptable. The difference I normally see between the horizontal and vertical dipoles (impedance variation at LNA input) was also much the same (0.25 dB) so I was fairly confident that at the higher frequency would not be detrimental to the receive performance. I also performed a quiet sky to Cygnus A check at the time and confirmed that the System Noise Temperature was normal as expected. (Y factor measured at 3.5dB)

The only way I believed I would be able to detect the Spacecraft was by use of **DSP software**. **Spectran**, written by I2PHD and IK2CZL is simply brilliant. I have used it for several years observing Aircraft Doppler Shift and Meteor Activity on both 144 and 432 MHz. The simplest description I can give here is that it takes the receiver audio input to a sound card and displays the signal in an amplitude time domain format on the screen as a white trace or line. Its main feature is to digitally reduce the bandwidth down to 0.032 Hz (selectable) and thus realise a most significant signal to noise improvement. (refer www.qsl.net/padan/spectran.html)

When the bandwidth is narrowed the "signal window" is narrowed as well. In this case the viewable signal window at 0.032 Hz is only 10 Hz and at 0.125Hz it is 80Hz wide. That is fine if you know the exact frequency, and have the time to allow for the overhead processing required to resolve and then display the signal. With a 1GHz processor, it takes about 10 seconds to display a signal in a 0.125 Hz bandwidth. In other words, when you vary your tuning the displayed signal is delayed by 10 seconds. Remember of course that you have to be within 80 Hz of the signal frequency to even see the signal display!

The problem we face with the MGS is that the frequency is not constant as the Doppler Shift between you and the spacecraft is varying rapidly for most of the viewable time. It is rotating around Mars and Mars and our selves are also moving at differing relative speeds. To calculate the Doppler shift requires some pretty smart mathematics to derive just where you should be tuned. As was described earlier, the space craft antennae (in use) is pointing towards the Martian surface and will be blocked by the body of the spacecraft except for a 5-15 minute period as it enters (ingress in spacecraft tracking parlance) or exits (egress) "occultation" or blockage by the planet.

During the approach to us the Doppler shift is positive (occultation egress...received frequency is higher than the signal frequency) and as it heads away from us it is negative. (occultation ingress. re-

ceived frequency is lower than the signal frequency) In round figures it in the order of 4 KHz but varies from orbit to orbit! MGS will be entering and exiting occultation (blockage by Mars with respect to Earth) with each 118-minute orbit. Again remember that the UHF antenna (with about 0 dB of gain, EIRP-30 dBm) will only be viewable from Earth for a few minutes (5 to 15 minutes) just before ingress and just after egress during each orbit.

As you can see, given the combination of varying Doppler Shift, the narrow bandwidth with a very narrow signal display window, it is vital that you know just where to tune at any given time. Fortunately I was supplied with one-minute predictions accurate to the Hz. I was some what taken aback when I first received them in an e-mail. After a cursory glance hit then hit the print button only to find there was 81 pages! I learnt to edit what I wanted first after that!

The next thing I had to do was to relocate my main computer down to the EME shack some 200 metres distance. This was necessary to run Spectran as the maximum speed I could muster was essential for fast decoding. My dish tracking and house keeping is done on a small IBM Notebook where speed is not important.

When I tuned on the computer and tuned to the appropriate frequency, low and behold Murphy has already predicted my actions and was producing the weakest of weak signal precisely where I did not want it! After removing the covers from the computer the source was traced to one of the Crystal Oscillators on the video card. Grabbing a bare alligator clip (no lead) I attached it to the Crystal case and the birdie moved outside the desired frequency range! Be aware that when using Spectran it is essential to keep audio leads well shielded with no current loops etc as its sensitivity is extreme and it loves to display 50Hz traces!

Now to check out the antennae tracking. Not a problem expected here, just click on Mars and away we go. This is my own tracking software that I (and many others) have been using for years. I was only using it the previous weekend looking at Jupiter's IoB pulses and "working a few off the moon". Clicked on Mars and away hummed the motors. Hang on, Mars is not setting it is rising!!! A day later I finally tracked the tracking error to the 2000 Epoch Ephemeris data having a 12-hour right ascension error! All the other Planets were spot on of course only Mars was wrong! Murphy again!

Now we were tracking Mars and JPL was about to turn on MGS at 1200 UTC on August 25th 2003. Incidentally, given its distance from Earth, the transit time for the signal to reach Earth is 3:05 minutes. Thus when commanding the "bird" it takes 6:10 minutes for you to know that your actions worked!

My first attempt to detect the MGS signal (CW only) was from 1221 UTC, (some 20 minutes after switch on) and at 1226 I "painted" a very weak signal at 437.103.500 (in a 1Hz BW) for 25 seconds. At this time I had not receiving the accurate Doppler calculations for my location and was relying upon the predicted "4 kHz shift". What immediately appeared suspicious was the observed signal was not exhibiting the characteristic Doppler Shift as expected. MGS at that time would have exhibited about a 1.5Hz per second downward shift. The signal I was "painting" showed little shift what so ever. Over the next few days I saw this very same signal on many occasions and finally deduced it was as a result of aircraft reflection from a source unknown.

During the first night I listened to 3 orbits from 5 to 15 minutes just before ingress and after egress. Unfortunately no signals were detected.

The following day, 26th September 2003, I received the Doppler predictions for my location. This allowed me a greater degree of confidence to know exactly where and when to tune. As a consequence it allowed me to narrow the bandwidth to the maximum possible. Alas still nothing, although at times one was becoming "line happy" thinking you could see a coherent line in the snow. This is very similar to "hiss happy" that many of us are familiar but in a visual format!

It was not until 28th September (after monitoring many orbits) that at 1328 I noticed (0.5Hz BW) a broken trace exactly where I expected to find the signal. The trace indicated that the signal's amplitude was varying at about a half second rate. I switched polarisation from Horizontal to Vertical and it seemed to "paint in" the missing signal for that time period. In other words it appeared that the signal might have been the result of the spacecraft spinning. I watched it for about 90 seconds and during that time the Doppler shift appeared to follow the identical signature as predicted. At 1356 I again saw a similar signal for another 15 seconds or so in the right Doppler "ball park area". I saw nothing after that time until I ceased tracking at 1410.

I reported my findings and asked about the possibility of the signal variation and received the following response. *The spacecraft is 3 axis stabilized, so signal variations cannot be explained by spin modulation. Also, the antenna on the spacecraft is circularly polarized ... which to me indicates that what I was seeing may not in their opinion have been the signal in question? (all I was really saying is that the amplitude variation could not be explained by spacecraft spin ... what you saw certainly could have been the spacecraft ... particularly since the Doppler signature seemed to match well.)* All the same I do feel fairly confident that it was the MGS given the exact Doppler signature it exhibited. Why I had not seen the signal in earlier orbits? I believe was due to the background sky noise. It was lower on the 28th and would have remained that way for the next few days. Unfortunately I was unable to track MGS on the following night prior to JPL turning it off at 1956 on the 29th due to 80 kph+. I could dearly like to have confirmed categorically the received signal was MGS. Murphy again!

It was a fun exercise commensurate with our Amateur radio pursuits and if nothing else showed just what can/cannot be achieved. A bit like proving Ohms Law I suppose? The result is known before hand .. but achieving it can be challenge.

Over view of the Mars Global Surveyor **Refer**
<http://mars.jpl.nasa.gov/mgs/>

Proposal: EU EME Contest 2005

Due to the obviously high demand we propose a change for the European EME Contest for 2005. We propose 3 weekends:

1st weekend: 144 MHz and 1296 MHz

2nd weekend: 432 MHz and 2.3 GHz

3rd weekend: 5.7 (+ 3.4?) GHz and 10 GHz.

For the first and second weekend we suggest 48 hours for the both bands. For the third weekend the same could be done OR may be 24 hours 5.7 GHz on Saturday and 24 hours for 10 GHz on Sunday.

Any comments would be most welcome. Also if there is need for a 3.4 GHz section. In any case the contest will be (at least) on 3 weekends next year.

73, Joe, DL8HCZ.

info@dubus.de

DXCC	CALL	Loc	ANT.	PWR/KW	Rem.	QSL mangr
ES	ES6RQ	KO28wa	12x14el. BV	1.5	Qrt	
	ES8X	KO18xc	4x	1	Exp. 2003 Exp 1991	
EU	No stn					
RA1	RW1AW	KP50eb	12x15	1.5	Qrt	DK3WG
	UA1ZCL	KP78tx	5,7m D	1.5	Qrt	(qsp)
	RA3DCR/1	KO55jr	4x29	0.3	Qrv?	Qsp
	RA3LE	KO64ar	8x10wl	1.5		qsp
	RA3YCR	KO73dh	8x26	1	Silent key	
	RW3PF	KO93cd	4x24	1	50-170°az el<30°	DK3WG
	RW3PX	KO84tf	4x21	1		qsp
	RW3RW	LO02rw	4x21	1	Not active	
	RZ3BA/1	KO56bc	1x38	1	Not winter time	RW3BP
	UA3PTW	KO93bs	20x15 BV	1.2	Qrv	DK3WG
	UA3TCF	LO26iu		0.5	qrt	DK3WG
	RA4AOR	LN29la	12x29 BV	0.1	Son from UA4ALU	DK3WG
	RV4AQ	LN28fg	4x18 BV	0.5	Not active	
	UA4ALU	LN29la	12x29 BV	0.1		DK3WG
	UA4API	LO20qc	4x19	1	Not active	DK3WG
	UA4NM	LO48up	4x25	0.8	qrt	DK3WG
	RA6AAB	KN94wr	4x30	1	Silent key	
	UA6LGH	KN97lf	16x25 KR	1		qsp
RA9	UA9FAD	LO88da	8x16	1.5		DK3WG
	UA9XEA	LP63un	4x	1	qrt	(qsp)
	UA9YLU	MO92eo	4x 7.5m		Will	DK3WG
UK	No stn					
UR	UR5LX	KO70wk		1.5	Qrt	DK3WG
	UT1PA	KO21fc	4x10wl	1	El. >45°	qsp
	UT3LL	KO80ac	4x27	0.4		qsp
	UT5DL	KN18ep	8x25	1	Not active	(qsp)
	UT5EC	KN78mm	8x27	1	Random only	qsp
	UT5UAS	KO50gk	6m-D	0.2	qrt	
	UT7VF	KN68mq	4x17 KR	1	Silent key	
	UY5HF	KN66hp				

Info on 70cms EME stations and exped. from ES, UA and UR

Tropo Reports

2m, 70cm and up

Editor: Wolfgang Schneider, DJ8ES
DJ8ES@gmx.de

Dear DUBUS-Readers,

Om Ivo Chladek (ZS6AXT) has sent the following very interesting report about the situation on VHF-, UHF and especially microwaves in South Africa:

VHF-, UHF- and SHF-activity in ZS is virtually non-existent at present. ZS tropo records in the bands 23, 13 and 6 cm are 23 years old and in the bands 10, 24 and 47 GHz are 10 years old. On all bands the distances achieved are very short as compared with Europe or USA. Nobody is QRV on the bands 13 cm and above, so that my first QSOs in the 13 and 6 cm bands were done on EME!

In 2002 myself, ZS6AXT and John, ZS6JON started to improve that. I have built up QRP portable transverters for all bands from 13 to 3 cm. We started in the 6 cm band, where the record was just 26 km (!) established in 1978. The 6 cm portable transverter had 400 mW and antenna was for the first two tests just an old horn with few dB gain. On my side I used the EME setup, 5m dish, with circular polarization, which due to surface inaccuracies works as a 3m one and 20 W transmitter power.

On 7. December 2002 we tried it in initial tests over 30 and then 50 km. In both tests signals were 59 on SSB. Next test was on 23 January 03 and signals were 519/529 only over the distance of 92 km. ZS6JON/p was operating only from the side of a motorway to the north, no high site. I built a 20 dB horn and on 27. February 2003 we made it over the distance of 151 km with 51/55 SSB reports. Tests over larger distance were negative. Only later, when checking my preamp on sunnoise I discovered that it was shot! Still, we did it.

After these tests we decided to improve the equipment and antenna further. The portable transverter is upgraded now, with a new basic unit which has cup filters on both, TX and RX side and output power just under 1 W. The previous transverter had just PCB printed filters. We got hold of a 60 cm dish and will use circular polarization for both ends. We have earmarked good spot bit over 300 km far and we hope to make it, to come closer to the distances achieved in northern hemisphere. Next band which will be tested soon is 13 cm, where we have 4 W portable transverter power and 1 m dish. On my side, it will be again my EME setup, centre of my dish is 6 m above the ground and it just about clears the surrounding houses. Present ZS record (1979) here is 198 km. All equipment for the above bands is designed and

built by myself. After that we will go for 3 cm, where the present record is standing at 268 km, established in 1993.

Hopefully more ZS/ZR stations will join by then, so that we will be able to try rainscatter and other modes over greater distances. For contact please E-Mail: ZS6AXT@GLOBAL.CO.ZA
vy 73, Ivo Chladek (ZS6AXT, KG33VV)

2m

LZ3GM, KN32rl, wkd (>800km):

25/11/03 11:00 RA4AOR	LN29LA	1535km
25/11/03 11:08 UT2IO	KN88XA	1025km
25/11/03 11:11 UA6MA	KN97VE	1102km
25/11/03 11:12 RV6LJK	KN97VE	1102km
25/11/03 11:13 RN6MT	KN97LN	1064km
25/11/03 11:22 RV6LJK	KN97VE	1102km
25/11/03 18:50 UA9XEA/6	KN97KG	1044km
25/11/03 19:31 RW3WR	KO71IM	1144km
25/11/03 19:58 UT2HN	KN79OI	970km
25/11/03 20:03 UT2QA	KN86KS	884km
25/11/03 20:09 UY5ON	KN89CW	1066km
25/11/03 20:54 RA3WDK	KO81BR	1218km
25/11/03 21:21 UR5LX	KO70WK	1093km
25/11/03 22:54 UR5LLW	KN89TQ	1119km
26/11/03 09:03 RV3ZR	KO80CL	1112km
26/11/03 09:15 RA3QTT	LO01GQ	1420km
26/11/03 09:28 RN3QR	LO00BK	1310km
26/11/03 09:42 RW3PF	KO93CD	1428km
26/11/03 10:22 UR5LX	KO70WK	1093km
26/11/03 10:25 RW3WR	KO71IM	1144km
26/11/03 10:29 UT2IO	KN88XA	1025km
26/11/03 10:38 UY5ON	KN89CW	1066km
26/11/03 10:51 UT8AL	KO61WP	1128km
26/11/03 15:43 UR4MBN	KN98ON	1136km
26/11/03 15:45 UR5MGW	KN98QO	1149km
26/11/03 16:00 UT5ER	KN78ER	878km
26/11/03 18:36 UY2QQ	KN77OT	848km
26/11/03 19:18 UX7LT	KN89TR	1122km
26/11/03 19:34 UR5EOA	KN78LC	883km
26/11/03 20:38 UA3RBO	LO03NG	1569km
26/11/03 23:11 UA4AQL	LO20QB	1608km
26/11/03 23:53 RA4AOR	LN29LA	1535km
27/11/03 16:11 RN6BN	KN95LC	967km
27/11/03 19:36 RA3QTT	LO01GQ	1420km

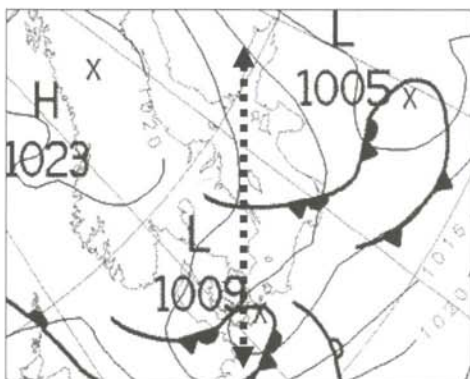
Rig: 3w BVO ant, MGF1302, 100 W.
tnx for info Ivan

PA5DD, JO22FE, wkd Jan - Feb '04:

06/01/04 18:52	OZ6TY	JO55XE	590km
06/01/04 18:51	OZ1BNN	JO55PM	577km
06/01/04 18:50	DL3TW	JO44SK	415km
06/01/04 18:48	SK7MW	JO65MJ	660km
06/01/04 18:45	OZ1ALS	JO45VA	468km
06/01/04 18:13	OZ9KY	JO45VX	545km
06/01/04 18:09	OZ1IEP	JO55XU	633km
06/01/04 18:07	OZ2LD	JO54TU	552km
06/01/04 18:06	OZ2PBS	JO55XJ	602km
03/02/04 19:27	OZ1ALS	JO45VA	468km

03/02/04 19:23	OZ7CQ	JO55GL	538km
03/02/04 19:22	OZ1ANA	JO55DK	523km
03/02/04 18:56	OZ6TY	JO55XE	590km
03/02/04 18:55	OZ1BEF	JO46OE	542km
03/02/04 18:53	OZ1DLD	JO45RL	490km
03/02/04 18:48	SK7MW	JO65MJ	660km
03/02/04 18:46	DL0RSH	JO43SV	382km
03/02/04 18:44	OZ1HLB	JO55RT	606km
03/02/04 18:44	OZ1IEP	JO55XU	633km
03/02/04 18:43	DL3TW	JO44SK	415km
03/02/04 18:43	OZ9KY	JO45VX	545km
11/02/04 17:48	SP1FPG	JO73GN	681km
11/02/04 17:47	LA4YGA	JO48AE	710km
11/02/04 17:45	GW3HWR	IO71XN	603km
11/02/04 17:44	OH2KW	KP20IJ	1533km
11/02/04 17:41	OH1NOR	KP00XL	1409km
11/02/04 17:39	SP3MGM	JO73QE	731km
11/02/04 17:36	OH5LK	KP30ON	1664km
11/02/04 17:34	ES6RQ	KO28WA	1493km

tnx for info Uffe



Tropo on Feb. 11th from PA to OH/ES on 2m:
1600km on the front side of a high pressure system are quite rare

RW3PF, KO93CD, wkd 26/11/03 (>700km):

08:33	UY5HF	KN66HP	559/579	825km
09:40	LZ3GM	KN32RL	579/579	1428km
12:52	RN6BN	KN95LC	59/59	895km
14:27	UR5GIM	KN76CL	59/59	794km
14:30	LZ2QA	KN43EK	579/559	1298km
15:50	UU1DX	KN74BW	599/599	959km
16:00	YO4GJH	KN35XG	56/59	1146km
16:35	LZ2PI	KN23XU	579/579	1367km
16:47	RA6DA	KN96HA	569/559	792km
16:51	UU2JD	KN75AC	599/579	943km
16:53	UT5JCW	KN64SN	57/58	1012km
19:42	LZ5GM	KN32QL	599/599	1431km

tnx for info Peter

DL8EBW, JO31NF, wkd (>700km):

06/11/2003 19:24	SK7HW	JO76JU	800km
06/11/2003 19:48	SM7XFD	JO77JO	864km
06/11/2003 20:02	SM6LIF	JO67RT	837km

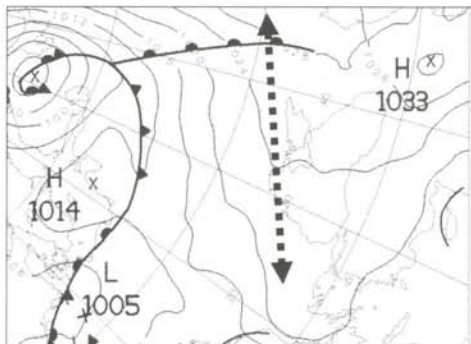
06/11/2003 21:10	SM4XIO	JP70NI	1129km
06/11/2003 21:16	SK6HD	JO68SD	872km
06/11/2003 21:33	SK6HD	JO68SD	872km
06/11/2003 21:34	SM6VSZ	JO67EB	733km

Rig : ICOM IC-275-H, 11 el DL6WU, GU74b
tnx for info Guido

LZ1KG, KN31CS, wkd on 26/11/03 (>600km):

07:28	UT5JCW	KN64SN	672 km
11:12	UY5HF	KN66HP	745 km
12:10	UT2QA	KN86KS	1015 km
12:38	UR5LX	KO70WK	1217 km
20:28	RV3ZR	KO80CL	1236 km
20:31	RA4AOR	LN29LA	1668 km
20:33	RN6MT	KN97LN	1196 km

Rig: IC275E, 2x 16el F9FT, 12m agl
tnx fr info Mitko



Tropo on Nov 25/26th from LZ to UA3/4
with QRBS on 2m up to 1700km

DL8BDU, JO43AA, wkd (>700km):

13/08/03 04:02	F8BON	IN96CV	979km
13/08/03 04:03	F6KSV	IN97RJ	875km
13/08/03 04:06	F5GHP	IN96LE	1003km
13/08/03 04:10	EA1NQ	IN73EM	1460km
13/08/03 04:16	F6FHP	IN94TR	1108km
13/08/03 05:10	F6APE	IN97QI	882km
13/08/03 05:13	F1YJ	JN17ET	705km
13/08/03 05:25	F5APM	IN97HG	927km
13/08/03 05:26	F6CJA	JN07UU	729km
13/08/03 05:41	F1DUZ	IN97NJ	891km
13/08/03 05:43	F6DZA	IN87GS	1003km
13/08/03 05:44	F5EGD	JN07IL	808km
13/08/03 05:47	F6FYA	JN07II	819km
13/08/03 05:51	F1SAC	IN96HA	1033km
13/08/03 05:55	F8DBF	IN78RI	1028km
13/08/03 05:56	F4ACE	IN96KC	1014km
13/08/03 05:58	F6FMR	IN99TE	727km
13/08/03 05:59	F5MKM	JN07XU	719km
13/08/03 06:01	F3VX	IN97SK	867km
13/08/03 06:05	F5HGO	JN05AI	1032km
13/08/03 06:09	F5FT	IN87GS	1003km
13/08/03 06:13	F5MGD	JN07AW	800km
13/08/03 06:13	G3OYF	IO81RL	742km

13/08/03 06:20	GW5NF	IO81KQ	774km
13/08/03 06:22	GW8JLY	IO81JM	785km
13/08/03 08:31	F6FAI	IN88HM	947km
13/08/03 08:55	F5SVO	IN97IG	923km
13/08/03 10:04	MU0FAL	IN89RL	836km
01/11/03 20:49	G4ASR	IO81MX	752km
01/11/03 23:03	F8KCF/P	JN36BP	722km
08/12/03 16:10	MM3ERP	IO87RJ	828km

Rig: TR-751E, PreAmp 0,65dB, PA 450W
tnx for info Klaus

DH9NFM, JO50RF, wkd (>700km):

05/09/03 15:01	HA8V	KN06HT	774km
06/09/03 15:13	9A2RY	JN84OF	795km
06/09/03 15:14	IK0DDP/6	JN63KC	799km
06/09/03 19:55	T90O	JN94IR	818km
06/09/03 22:19	9A1O	JN95IT	729km
07/09/03 10:49	9A3V	JN74OC	737km

tnx for info Chris

OZ4VW, JO45VT, wkd Nov '03 (>700km):

06/11 22:58	F2YT	57/59JO10GK	765km
07/11 18:16	LY2HM	53/52KO15CX	779km
07/11 18:19	YL3DW	57/59KO26CW	900km
07/11 19:08	LY2IC	55/53KO14WW	899km
08/11 19:34	SP7TEE	55/55JO91SS	786km

tnx for info Arne

70cm

PA3DZL, JO21HM, wkd (>600km):

19/03/03 21:36	GM6VXB	IO97AQ	805km
19/03/03 22:40	SM6ESG	JO67CC	791km
08/07/03 20:26	OZ9KY	JO45VX	600km
10/08/03 20:03	LA4C	JO29OG	862km
10/08/03 20:15	LB8SE	JP20MN	1005km
10/08/03 20:29	LA6MV	JO59EJ	947km
10/08/03 21:05	LA6LCA	JO59FE	928km
15/09/03 19:15	YL3AG	KO26AW	1391km
06/11/03 19:12	SK4BX	JO79OF	1084km
06/11/03 19:17	SM7HGY	JO86DR	953km
06/11/03 23:31	SM3LBN	JP80IO	1254km

Rig: 400W, 2x21 el. 6.6wl, PreAmp 0.3dB
tnx for info Jac

RW3PF, KO93CD, wkd 26/11/03 (>500km):

08:35	UY5HF	KN66HP	579/579	825km
11:29	UT2IO	KN88XA	579/579	570km
11:36	UA9XEA/6	KN97KG	559/559	655km

tnx for info Peter

DH9NFM, JO50RF, wkd (>500km):

04/10/03 20:03	G8P	JO01QD	717km
04/10/03 20:22	ON4KBE	JO20BI	521km
04/10/03 21:13	IK1PJU/1	JN44GK	680km
04/10/03 22:47	F6HLC	JN26HS	633km
04/10/03 23:01	G5LK/P	JO01OC	728km
05/10/03 05:06	PA6NL	JO21BX	547km

tnx for info Chris

PA2CHR, JO22XA, wkd (>1000km):

10/11/03 19:52	ES2WX	KO29JN	1439km
10/11/03 20:14	ES2RJ	KO29JN	1439km
10/11/03 21:39	SM1NJC	JO97DO	1005km

Rig: FT847, PreAmp MGF1302 with NF: 0.32 dB,
YL1050, ant 4 x 35el BVO/11wl
tnx for info Chris

DL8OBU, JO42XI, wkd (>900km) in CW:

10/11/03 18:30	OH6HFX	KP14	1515 km
10/11/03 18:40	ES2RJ	KO29	1220 km
10/11/03 19:19	OH4LA	KP20	1275 km
10/11/03 20:30	SM2DXH	KP03	1412 km
10/11/03 20:40	OH6QR	KP22	1413 km
09/12/03 19:20	SM4DXO	JP70	985 km
09/12/03 19:26	SK4AO	JP70	982 km
09/12/03 19:50	SM3BEI	JP81	1083 km

tnx fr info Jürgen

PA5DD, JO22FE, wkd:

10/02/04 20:11	G8OHM	IO92AJ	454km
10/02/04 20:08	G8VHI	IO92FM	426km
10/02/04 18:30	OZ1DLD/P	JO45SK	490km
10/02/04 18:23	SK7MW	JO65KN	661km
10/02/04 18:18	OZ1FF	JO45BO	451km
10/02/04 18:14	OZ9KY	JO45VX	545km
10/02/04 18:12	OZ6HY	JO45WA	472km
10/02/04 18:11	OZ1IEP	JO55XU	633km

tnx for info Uffe

23cm + up

DH9NFM, JO50RF, wkd (>400km):

04/10/03 15:35	ON7WR	JO20EP	503km
04/10/03 16:55	IK4ADE	JN54OE	672km
04/10/03 17:07	IK3COJ	JN65BN	521km
04/10/03 17:09	IW4BET/4	JN54OE	672km
04/10/03 17:52	PA0WMX	JO21XI	406km
04/10/03 18:00	PA5DD	JO22IC	514km
04/10/03 20:07	G8P	JO01QD	717km
05/10/03 06:34	I4CVC	JN54WH	658km
05/10/03 07:10	I4LCK/4	JN54PD	676km
05/10/03 07:38	IW3RMR	JN66PA	487km
05/10/03 09:05	PI4ZLD	JO11WM	551km
05/10/03 09:08	PA0S	JO21FW	524km
05/10/03 11:06	PA3AWJ	JO21GW	519km
05/10/03 11:56	ON4CDU	JO20KW	471km
05/10/03 12:04	PA6NL	JO21BX	547km
05/10/03 12:28	S57C	JN65XM	555km

tnx for info Chris

OZ4VW, JO45VT, wkd Nov '03 (>500km):

06/11 20:13	DK1VC	57/59JO31RG	528km
06/11 22:31	SM4DHN	59/59JP60VD	538km
07/11 15:48	SM3LBN	57/55JP80IO	672km

2320 MHz:

06/11 22:34	SM4DHN	59/59JP60VD	538km
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tnx for info Arne

DL8OBU, JO42XI, wkd (>1000km) in CW:
 10/11/03/03 18:45 ES2RJ KO29 1220 km
 tnx fr info Jürgen

OZ1FF, JO45BO, wkd 23cm on 17. Feb '03:
 SK7MW (JO65KN), G3XFY (JO02OB), SM7ECN
 (JO65NQ), SK0UX (JO99BM), DL1SUN
 (JO53PN), SM6VTZ (JO58XJ), LA2Z (JO59EJ),
 DJ6JJ (JO31LG), SM7GEP (JO77IP), PA0S
 (JO21FW), PA0EHG (JO22HB)
 tnx fr info Kjeld

PA3DZL, JO21HM, wkd (>300km):
1296 MHz:

13/07/03 21:26	G4JTJ	IO92SD	356km
13/07/03 22:09	G4BRK	IO91DP	438km
10/08/03 20:31	LA6MV	JO59EJ	947km
10/08/03 21:06	LA6LCA	JO59FE	928km
10/08/03 21:15	SM6DJH	JO58	878km
16/09/03 18:49	OZ9KY	JO45VX	600km
16/09/03 18:53	OZ2LD	JO54TU	595km
16/09/03 19:04	OZ1FF	JO45BO	509km
06/11/03 18:25	SM6DJH	JO58PI	870km
06/11/03 18:28	SM6AFV	JO57GQ	784km
06/11/03 18:32	SK7MW	JO65MJ	703km
06/11/03 19:03	OZ6OL	JO65DJ	665km
06/11/03 19:07	SM6CEN	JO57XO	824km
06/11/03 19:25	SM6OEJ	JO67DA	787km
06/11/03 19:29	OZ2LD	JO54TU	595km
06/11/03 20:41	F5HRY	JN18EQ	353km
06/11/03 21:36	F6HMQ	JN18GP	352km
06/11/03 23:34	SM3LBN	JP80IO	1254km
07/11/03 00:03	OZ1CTZ	JO46OE	599km

Rig: 150W, 23el F9FT yagi, PreAmp 1.0dB

2320 MHz:

13/07/03 22:14	G4BRK	IO91DP	438km
13/07/03 22:17	G3XDY	JO02OB	242km
13/07/03 22:22	G4FEV	IO92RG	364km
10/08/03 21:08	LA6LCA	JO59FE	928km
11/08/03 20:07	LA6LCA	JO59FE	928km
06/11/03 18:43	SK7MW	JO65MJ	703km
06/11/03 20:48	F5HRY	JN18EQ	353km

Rig: 50W, 95cm dish, PreAmp 1.0dB
 tnx fr info Jac

PA5DD, JO22FE, wkd Jan – Feb '04:
23cm:

20/01/04 18:59	OZ6OL	JO65DJ	620km
20/01/04 18:57	DL3TW	JO44SK	415km
20/01/04 18:55	SK7MW	JO65MJ	660km
20/01/04 18:46	SM7ECN	JO65NQ	681km
20/01/04 18:29	OZ1BGZ	JO65AP	623km
20/01/04 18:25	OZ2LD	JO54TU	552km
20/01/04 18:22	OZ1FF	JO45BO	451km
20/01/04 18:17	OZ9KY	JO45VX	545km
17/02/04 20:02	OZ2OE	JO45VV	538km
17/02/04 18:36	OZ1FF	JO45BO	451km
17/02/04 18:34	SK7MW	JO65KN	661km

17/02/04 18:21	SM7ECN	JO65NQ	681km
17/02/04 18:14	OZ2LD	JO54TU	552km
17/02/04 18:11	OZ9KY	JO45VX	545km
17/02/04 18:10	OZ1BGZ	JO65AP	623km

13cm:

27/01/04 18:49	SK7MW	JO65MJ	660km
16/02/04 11:49	G0RRJ	IO91FE	441km

tnx fr info Uffe

Personality – ZS6AXT

DUBUS wants to introduce active dxers from europe and overseas. Today: ZS6AXT.

OM Ivo Chladek lives in Krugersdorp (KG33VV), 25 km west from the centre of Johannesburg. It is the highest point of the area, few metres below 1800m.a.s.l. Virtually all equipment and antennas for VHF, UHF and SHF are homemade and mostly own design.

Ivo started in Feb. 1948, licensed as club station operator in Dec. 49 (that time it was difficult to get private licenses in OK) and then about 1956 got his license as OK2VCG, changed to OK2WCG, which he holds until now. That time he was qrv on 2m and 70cm tropo, MS and satellites and also preparing 70 cm EME.

In 1968 he left for ZS and here after settling down started on VHF again, satellites. In 1988 he made the first QSOs on 23 cm EME with a 2,6m home-made dish, 11 stations were worked. In 1989 he built a 5m dish and continued on 23cm, later also on 70 cm. So far he has worked 188 stations on 23 cm and 152 on 70cm.

In 1997 Ivo started on 13cm and so far he has worked 31 stations. However, he had to remove the 70 cm antenna from the dish. This is now with ZS6JON, who should be QRV on 70 cm soon.

In 1999 he started on 6cm and so far worked 13 stations. This is all with the homemade 5m dish. Now Ivo is preparing for 3 cm EME with another solid 2,4m dish. In the meantime with ZS6JON they are preparing to break all ZS distance records, in the 13, 6 and 3 cm bands. So far they managed it on 6cm.

OM Ivo was also active on 6m for the last 15 years and has 135 countries worked so far. VHF activity in ZS is virtually non-existent with the exception lately of digital modes which do not appeal to him. Thus since he is QRV on 2m, 70 and 23cm tropo, but he does not operate there.

E-Mail: ZS6AXT@GLOBAL.CO.ZA

E-Mail-Account DJ8ES

Because of spam I have to change my Email-account. Please send your next Tropo-Report to:

E-Mail: DJ8ES@gmx.de

6m News

Reports, Expeditions, Infos

Editor: Joachim Kraft, DL8HCZ

6m-Logbook from DL (Day, Month, UTC):

26.11. 2040-2200 Es from all DL: TS7N JM54
Not too much to report on 6m from DL during the last 3 months. A few Es openings and Auroras occurred in December but nothing special.

But in some regions of the world there is still live on 6m even in this part of the solar cycle. As an example here the list of my TEP observations when qrv as **CT1HZE from southern Portugal (IM57nh) from September to November 2003.** Solar flux was mostly between 90 and 130. Here the list:

8.9. 1900 ZD8VHF 519
11.9. 1555-1615 ZD8VHF 559
1622-1637 TR0A 558 1715 TR0A 519
12.9. 1530-1545 TR0A 519
14.9. 2015-2206 ZD8VHF 559
15.9. 2030 ZD8VHF 519
20.9. 1950-2107 PP1CZ/B 419 2107 PY1RO
21.9. 1558-1700 TR0A 559 1653 7Q7RM 559
22.9. 2150 ZD8VHF 539
23.9. 1540 TR0A 529
2020-2220 PP1CZ/B 529 2023 PY1RO
2030 PY2SFY 539 2220 ZD8VHF
24.9. 2130-2230 ZD8VHF 559 + PP1CZ 529
25.9. 2045-2240 ZD8VHF 559 - 599
26.9. 2205 ZD8VHF + PP1CZ 559
27.9. 1950-2120 ZD8VHF 559
28.9. 1519 TU2OJ IJ76 55
2120-2130 ZD8VHF 559
29.9. 1655 7Q7RM 529
1850-1920 ZS6TWP/B 539
2025-2125 ZD8VHF 559-579
2030 PY1RO GG87 549
2150-2200 PY2SFY GG77 + TR0A 419
30.9. 2020-2220 ZD8VHF 579
2210-2220 PP1CZ 519
1.10. 2200 ZD8VHF569, PP1CZ, PY1RO
2.10. 1545 3C0V 51 hrd
2100 ZD8VHF 599, PP1CZ 529
3.10. 2100-2240 ZD8VHF 539,
2130-2200 PP1CZ
4.10. 2145-2230 ZD8VHF 559, PP1CZ 539
5.10. 2115-2225 ZD8VHF 549, PP1CZ 519
6.10. 2025-2125 ZD8VHF 579, PP1CZ 529
7.10. 2145 ZD8VHF 599
8.10. 2020-2200 ZD8VHF 579 2155 PP1CZ 51
9.10. 2005-2120 ZD8VHF 559
10.10. 2025 ZD8VHF 559

11.10. 2105-2240 ZD8VHF 579, PP1CZ 559
2113 PY1RO 599
12.10. 2150 ZD8VHF 539, PP1CZ 529
13.10. 2135 ZD8VHF 529, PP1CZ 419
16.10. 1814 PY2SFY 529, PP1CZ 559
1830 PY1RO 59, 1854 LU9EHF529
17.10. 2025-2140 ZD8VHF 559
18.10. 2125 ZD8VHF 579, PP1CZ 559
19.10. 2205 ZD8VHF 539
21.10. 2210 PP1CZ (OP himself) 59 59
23.10. 2105 ZD8VHF 599
25.10. 2050 ZD8VHF 579
26.10. 2130 PY1USK 52, ZP6CW 419
2130-2230 ZD8VHF 579-549
27.10. 2235-2250 ZD8VHF 599, PY1USK 54,
LU9EHF 519, PP5LD 52,
PY5DJ 529, PY2PAI 53,
PY2PRS GG66 559
30.10. 1630 TR0A 559
31.10. 2010 PY1RO 59, PY2SFY 529
3.11. 2110 ZD8VHF 559
4.11. 1442 TR8CA 539 TR0A 529
2120 ZD8VHF
5.11. 1615-1730 7Q7SIX 419, TR0A 579
2230 LU9EHF 519, ZD8VHF 579
6.11. 1520 RW1AY/MM JG67ot 59 59
TR0A 559, ZS6TWP 539
2000-2130 ZD8VHF 549
7.11. 1419 RW1AY/MM JG68lj 59 59
2015 ZD8VHF 559
8.11. 1610 TR0A + TR8CA 599
9.11. 1500 RW1AY/mm53, ZS6TWP 529

Nothing until Nov. 14th when observations stopped due to my QSY to DL. Experience from the last years shows that Dec./Jan. are the months with very little chances for TEP. So this timeframe was very bad for the FR/E expedition and resulted in no 6m TEP-QSO from Nov. 24th to Dec. 15th!

The best timeframe for TEP for me was from 1 week prior the September equinox to 6 weeks after. Many days with the ZD8VHF (II22TB) and PP1CZ (GG99) beacon in for hours and absolutely no other signal on the band.

Activity from ZD8 stations is zero now and also from Brazil very low activity now compared with the previous 2 years.

My rig in CT: IC706, 4 ele Yagi, 8m a.g.l., 96m a.s.l., with a free horizon all over the Atlantic from ZD8 and PY to the Caribbean and U.S. BTW even with this superb location it is absolutely impossible for me to work any F2-scatter with only 10 W output. Even 100 W is not enough quite often although I was the only calling station. And also TEP-QSOs with 10W are not possible too often.

I am quite sure that TEP from CT to ZD8 and may be also PY is possible quite often during the maximum of the solar cycle.

As expected almost no TEP from mid January to mid February 2004. But the spring TEP season started again in mid February:

CT1HZE, IM57nh, hrd in 2004:

17.2.	2250	ZD8VHF 559
18.2.	2240-2300	ZD8VHF 559
19.2.	2200	ZD8VHF 559
20.2.	2300	ZD8VHF 539
21.2.	1755-1815	TR0A 539 2120-2215 ZD8VHF
22.2.	2055-2225	ZD8VHF 559
23.2.	2210	ZD8VHF 559
24.+25.2.		NIL
26.2.	2205	ZD8VHF 549
28.2.	2200	ZD8VHF 559
29.2.	1645	TR0A 529, TR8CA 55

Aurora-E from VE4 to Europe!

On January 10th 2004 VE4/KG0VL (Loc. EO26) hrd GB3LER/B from 0059-0110z on 6m with max. 48A, QTF 10. Jeffrey says the sig sounded like a mixture from Aurora and Aurora-E. Before, at 23:30 also TF3SIX (HP84) was heard. No QSOs...

Ak (Boulder) for Jan 10th was 24, so nothing spcial. On Jan. 9th to 11th there were weak to mediocre Aurora openings from N-DL on 2m + 6m.

Expeditions and News:

/MM G0KZG/MM, Andy has changed plans again and plans to be qrv again on 6 + 2m in May/June 2004 from a ship in the North sea.

/MMRW1AY/MM is qrv on 6m from **JG67, 68, 77** with 100w and wire antenna until April 2004.

3B9 Rodriguez: From 16. March 2004 a big group will be qrv for 4 weeks also on 6m under 3B9C.

KP39/49/38 The OH8K group plans to activate these squares in June also on 6m.

KP18/26/27 The OH8K group plant to activate these squares in August also in 6m.

5U Niger: 5U7JB (US-Op) ist qrv on 6m for a longer period. Loc. JK13, QSL via ON5NT.

5V Togo: 5V7BR is qrv on 6m since Nov 17th 2003 with 100w and 5 ele Yagi. Loc.: J08.

7X Algeria: 7X0AD (JM16jr) says that there is good chance now to get a licence for 6m. He states that he will be qrv soon.

9Q Zaire: SM5DIC is qrv as 9Q0AR on 6m until the end of March.

D2 Angola: 4L4FN, Ed, is qrv as D2PFN for a longer period and it was expected that he gets a yagi for 6m soon.

FT1Z Amsterdam Isl.: New licenced Op Sebastien, FT1ZL, is qrv now until the end of 2004 on 6m. Loc.: MF82.

HI Dom. Rep.: HI3TEJ is qrv with 100w + 5 ele on 6m. Loc. FK59. QSL via ON4ANT, now ON4IQ.

S0 DL8YHR still plans to activate West Sahara on 6m (incl. EME) in summer 2004. But there are li-

cence problems with 6m. Obviously also some of the previous 6m operations had no licence, including S01HA... If S0 will not be possible, alternativsey ZB2, Gibraltar, may be activated.

TF TF/SM6CMU will be qrv on 6m from several squares on Iceland from June 3rd to 10th.

J3 Grenada: W7XU plans a QRO 6m activity during the coming summer Es season.

J7 Dominica: J7/W6JKV will be qrv with QRO and good take off to EU from June 24th to July 5th on 6m.

JW Svalbard: JW5RIA is qrv again until June on 6m from Bear-Isl., JQ94.

LA LA/PA2DW will be qrv from several squares in LA during the whole month of July also on 6m.

PJ7 St. Maarten: K4BI + K2ZD will be qrv on 6m from July 2nd to 12th.

S9 Sao Tome: S9TX (ex S92TX) ist qrv on 6m now for about 1 year with 100w + 5 ele Yagi. Loc. JJ30. QSL via W7KNT

SV/A Mount Athos: SV2ASP/A ist qrv now on 6m and made his first Tropo QSOs with SV1's. So far he only has a dipole and 30w but a small yagi will follow soon. He promised to be active in the coming summer Es season.

TN Congo: CT3HK is qrv on 6m as TN3S until October 2005. QSL via Madeira Team, P.O. Box 19, P-9001-901 Funchal, Madeira Isl., Portugal.

Obviously he needs to be a bit more encouragement for 6m operation....

VP2M Montserrat: VP2MEG is qrv now with 100w and a 4 ele yagi on 6m. He is a resident.

VP9 Bermuda: VP9/N0JK will be qrv again on 6m from June 10th to 15th. And VP9/W3CMP will be qrv again on 6m from June 26th to July 5th.

VQ9 Chagos: VK9LA is qrv on 6m for one year.

YA Afghanistan: G4KUX is qrv until April 2004 on 6m as YA4F from Kabul, MM44ON. He is good for DXCC now. Also YA1D is very active on 6m now.

YK Syria: YK1AO, Omar, has announced that he will erect a 6m antenna but nothing heard so far...

XV Vietnam: XV9DT is qrv for some years from Hanoi also on 6m. OL20.

ZA Albania: DL7AJA plans to be qrv from Mai 23rd to June 5th on 6m from JN91.

ZF Cayman: ZF1DC and K4BI (ZF2BI) will be qrv on 6m from June 1st to 9th.

ZL New Zealand: GM3WOJ is moving permantly to ZL and will be qrv on 6m als ZL1CT. He hopes to work many EU's in the peak of next cycle.

Beacons

Check the 6m beacon list at the end of this issue. There are many updates, corrections and proposals included which you will not find anywhere else.

Meteor-Scatter

LA/PA2DWH in Norway 2003

PA2DW, Dick

2m MS-operation during the holidays.... According to my XYL that is not a good idea, so I kept it limited to some hours in the evening. Some indicated my last years operation an "expedition". Although taking the whole family and camping gear to Norway is an expedition in itself, my radio-operation was never intended an expedition!

In the last days before departure I decided to take some radio-gear with me on our holiday-trip, so it was a real ad-hoc decision. I took my FT817 and the TE-systems 1452G (normally driving my GS35b final) with me and as I had no time to arrange a proper antenna, I decided upon an old 4 el Flexa yagi. I made a direct connection to my car-battery, so could run nearly 300W, completely independent from the mains! See picture below.



The first tests I ran in JO58, although near a beach with seemingly good take-off, with PA3BIY. He could not hear a single ping hi! It was until we came to JP41 the day after and I drove up a high spot, that we made a successful QSO.

We stayed 2 days in JP41 and I worked a couple of DL and PA stations. We then went to our destination in JP31, Lusterfjord, nearby the Jostedal Glacier where we wanted to go for Glacierwalking (see picture 2).

On the campsite I asked for a nearby high spot with good visibility to the south and when I replied his question "what is that good for" you could spell his thoughts "this guy is totally". The road up to that high spot was one normally to be taken with a 4-wheel drive so it was actually an expedition for my heavy station car hi! I went up there almost every evening during our stay in JP31, so got used to the steep road and bad pavement after a couple of days.

Underneath a picture with an impression of the place and my "mains-independent station".

I worked a lot of stations from OZ, PA, DL one HB and one OH (OH6MAZ, on backside of my antenna!), but no G's where to be heard at all! I was obviously too much closed-in by the



mountains for that direction. S53VV was obviously too far out for that small 4 el yagi...

But I also missed a lot of other stations, sorry for those who desperately tried.

Mostly I worked in FSK441, some QSO's in high speed CW. After JP31 we went straight east towards JP51 square. On the way I tried some operation from JP42, but that was a very bad QTH for the south and no roads nearby upwards...

In JP51 I had a nice spot on a campsite nearby lake Femund. To the south the take-off was not excellent, but I worked a couple of stations, again mostly FSK441, but two in HSCW.

To the North I had a nice take-off, so I finally worked some aurora! I had experienced only once before a huge (!) pile-up in aurora (that was in 1978, being on a "real expedition" as GM5CJF in IO99/IP90) so it was a special experience! I worked no further south than OZ, but was heard by PA3CEE (Eltje, next time we make it!).

A few days later, right after a MS-QSO with DD3SP, my amplifier gave up. Later I found it was a silly mechanical matter, but being out in the campsite with family etc, it was too hard to find it and fix it...

We went via SM4 back towards home, spending canoeing in Vaermland one week.

Since I had no QRO, I only worked a couple of stations from JO69 in NAC with the bare FT817 and still the 4 el yagi. Best DX was OZ1IEP, not bad!!

I enjoyed the holiday-operation and will surely take my gear again next time and of course a better antenna. The only thing



I would like to ask of some operators is to take things less seriously and not regard me as an expedition! That is the only way my family will let me do it and we will all have fun. This year our holiday plans again are going north, provided that the weather will not be too bad! We plan to be there in whole month of July and go as far up as Lofoten Islands. Watch the DX-cluster for any announcements/info! By the way, I have changed my callsign just after coming back from holidays, so PA2DWH is now PA2DW! Best 73, Dick/PA2DW

RW3AZ/1 Expedition Log

QTH: KO57AV, PSKOVSKAYA OBL.

All QSOs wkd on 144 MHz FSK (all random):

2003-08-08

0420	0459	OK1DFC	JO60TM	27 26 C
0530	0535	DJ8MS	JO63CT	27 37 C 11sec
0600	0640	PE1HWO	JO21GV	27 27 NC 13sec
0620	0630	HA8CE	KN06EN	27 26 C
0645	0700	DL6BF	JO32QI	27 27 C
0830	0909	DJ9CZ	JO31BC	27 27 C
0910	0924	OZ2M	JO65FR	26 26 C
0925	0945	SP6IWQ	JO80HK	27 26 NC
0945	0955	DK1KO	JO53CT	27 27 C
0955	1004	DL9AN	JO63SN	27 27 C
1005	1043	DJ7OF	JO51HG	27 27 C
1100	1120	DJ9FG		27 27 C
1130	1200	DL7UAE	JO62PH	27 27 NC
1915	1925	DL9MS	JO54WC	27 27 C 20sec
1930	1950	DL5ZA	JO51HK	27 26 NC
1950	2000	DK3BU	JO33NO	27 26 C
2033	2040	DJ9MG	JO52TC	27 26 C 22sec
2040	2053	DH3YAK	JO31RS	27 28 C
2103	2106	S57TW	JN75EX	27 27 C
2120	2130	OZ1LPR	JO44WX	27 27 C
2131	2143	OK2BRD	JN99ET	27 27 C
2145	2206	OZ1IEP	JO55XU	27 26 C

2003-08-09

0515	0520	DL5WG	JO52VK	27 27 C
0520	0540	DJ4UF	JO30BR	27 26 C
0610	0615	OE3MWS	JN88DC	27 26 C
0700	0710	DC7OH	JO62OK	27 27 C
0710	0714	DK1KO	JO53CT	27 27 C
0714	0730	DH7FB	JO62SM	27 27 C
0744	0815	DL1EAP	JO31IK	27 27 NC
0816	0830	DL8BDU	JO43AA	27 27 C
0837	0844	DL4DWA	JO61QH	27 37 C
0844	0859	IZ4AIK	JN63HV	27 26 NC
0900	0910	OK1YK		27 27 C
0911	0916	HA6ZB	JN97MM	27 27 C
0925	0934	SP6IWQ	JO80HK	27 26 C
0945	0953	DG0KW	JO64MH	27 27 C
0953	1002	OK1FFD	JO60CF	27 27 NC
1002	1006	DF2ZC	JO30RN	27 26 C 15sec
1008	1016	SM7JUQ	JO65VX	27 27 C
1102	1200	DG5CST	JO60DC	27 27 NC
1202	1206	PA0JMV	JO21PM	27 27 C 20sec



RW3AZ



RA3AH



RA3ABZ

1216	1236	OZ1LPR	JO44WX	27 27 C
1240	1244	PA2CHR	JO22XA	27 27 C 20sec
1248	1303	SM7FMX	JO65KN	27 27 C
1313	1322	DL7FF	JO62TJ	27 27 C
1324	1350	SM6CMU	JO57XK	27 26 NC
1350	1353	OZ1PIF	JO65AN	27 26 C
1400	1411	SP2OFW	JO93AC	27 27 C
1455	1530	PA3CSG	JO21MM	27 26 C
1615	1640	PA0C		27 27 NC
2250	2257	S57EA	JN76HE	27 26 C
2300	2345	IZ4AIK	JN63HV	27 26 NC

2003-08-10

2345	0010	PA3BIY	JO22EB	27 27 C
0010	0018	DL5GAC	JN47UT	27 26 C
0026	0033	OK1FFD	JO60CF	27 27 C
0045	0057	DG6JF/P		27 27 C

2300 2330 PADCOL JQ22MM 37 NC
2330 2359 DLSRBO JN69HG 38 27 NC
2003-08-12
0636 0649 QZ5ASG 27 27 7800
0729 0746 RY1AS 27 26 1225sec
0746 0805 CBEAF 27 27 917sec
0806 0814 PALUMV 27 27 11800
0833 0840 DKSLA 27 27 1148sec
0903 0907 FGCRP 27 26 988sec
0950 0959 ONNAV 27 26 8420
0992 0920 DL8BDU 27 26 8420
0994 0929 DKATE 27 27 8320
0996 0956 SPAMPB 27 26 1717sec
1022 1037 EASAXV 27 26 9280
1056 1125 DQDKW 26 26 111sec
1100 1125 HAT5MS 26 26 209sec
1356 1405 DK7BY 26 26 10880
1406 1425 G3WZT 27 27 10480
1427 1520 F6PHF 27 27 11440
1606 1625 HB8ZTA 27 26 10480
1630 1642 D7FKF 27 27 10380
1731 1742 K1XCDU 27 27 9220
1825 1837 L29M 27 27 7160
1932 1935 DL7DA 27 26 11560
1936 1947 DLGZAU 27 26 81560
1957 2003 PAKSM 27 27 10820
2025 2047 D7HFH 27 26 1026sec
2053 2102 GWGLWE 26 26 11300
2103 2113 PASECU 27 27 10700
2122 2229 SM7XON 27 27 1025sec
2153 2207 DL3AS 26 27 10280
2207 2223 3VMPI 27 27 8160
2225 2237 51WKH 27 27 101sec
2241 2301 LASKO 26 26 965sec
2322 2333 F1DUZ 27 26 10640
2333 2348 PAZDW 26 27 18800
2003-12-13
0000 0006 L2TBEP 26 27 15200
0008 0015 9A1CALP 27 26 11580
0019 0030 LZMM 27 27 1025sec
0046 0952 QZ2M 26 27 5620
0113 0128 FNSR 26 27 10760
0140 0154 D0TEL 27 27 978sec
0156 0207 D0HFG 27 26 147sec
0246 0354 SVAT 27 27 10340
0255 0312 EASZF 27 27 7380
0324 0358 WZHAJ 27 27 52sec
0508 0515 EASDXU 27 27 9360
0521 0528 SMHPV 26 27 1717sec
0549 0603 DL1RTL 26 27 105sec
0603 0608 DLOVV 27 27 94sec
0645 0653 G7PPT 27 27 933sec
0658 0712 W7CHX 27 27 10760
0722 0727 D5JCZ 27 27 1015sec
0803 0807 ESSRQ 27 27 204sec
1012 1017 D89YH 27 27 10380
1243 1254 DQXBC 26 27 9360
1330 1344 OHZZZ 26 26 102sec
1533 1513 YL3DW 26 26 1067sec
1705 1710 G4LTR 26 26 11380
1726 1733 LA3BO 26 26 10300

Ops: Alex, RW3AZ, Harry, RA3AH and Vladimir, RA3ABZ

YZ7MON BCC Meteor Scatter Contest Log 2003

YZ7MON (KN05FG)

wkd on 144 MHz in FSK441:

2003-12-11

2003 2013 SM7DLZ 27 27 1055sec

2013 2021 DL3LST 27 26 10300

2021 2031 F1EBK 27 26 825sec

2048 2053 PE1AHX 26 27 10360

2053 2058 D9NYC 26 27 1015sec

2108 2119 SP2JON 26 26 1015sec

2123 2129 LA0BY 26 27 9280

2129 2137 PA0BAT 27 27 9280

2159 2166 DGLY 27 27 927sec

2147 2155 PE1HX 27 27 9800

2228 2247 SM7FAX 26 27 11640

2003-12-12

0007 0014 DL1G5T 27 26 10760

0014 0021 ON4KHG 26 26 6200

0054 0100 PA1BWM 26 27 11380

0100 0122 DL4EBV 27 27 10380

0130 0141 D9JYE 27 27 102sec

0203 0215 DLOAO 27 27 172sec

0404 0416 PE1HWO 27 27 10480

0519 0544 DL1RHW 27 27 10520

0613 0617 LZ2FO 26 27 102sec

0125 0135 QETSON JN88FF 27 26 C
0530 0548 DO1JNP 27 26 C 10sec
0630 0655 S3VVJ JN65UM 27 27 C
0715 0715 D3JLE JO44TN 27 27 C
0750 0710 SM6CMU JO57XK 27 27 C
0800 0830 QZ2TF JO46PE 27 27 C
0830 0837 DG0GDP JN39H 27 27 C 20sec
0845 0900 DL8EP JN39H 27 27 C 20sec
0910 1015 RK2FWA KO44TF 27 37 C
0915 0922 HAT5MS JN67UO 27 37 C
0932 0945 D9GBEM JO46UO 27 37 C
0945 0950 DL0KRG JN49CC 27 37 C
0950 1000 JN6BWX JN77PK 27 27 C 15sec
1033 1040 IEDXW JN65QO 27 26 C
1121 1151 9A3JH JN7540 27 27 NC
1151 1215 DL1RHW JO62HG 27 27 C
1305 1305 DG0GDP JN887? 27 26 C
1305 1320 D66AH JN887? 27 26 C
1330 1400 DLSMCG JN58KH 27 27 C
1425 1439 D7FOG 27 27 C
1715 1727 DL8EBW JO31NF 27 27 C
1715 1727 DL8EBW JO31NF 27 27 C
2330 2359 LZ2FO KN13XK 27 26 NC
2003-08-11
0000 0040 IZ4AK JN33HV 27 26 C
0040 0055 PA3BZL JN2358 27 26 C
0114 0140 PA2DWH JO22SD 27 27 C
0140 0159 D9DEN JO31JK 27 27 C
0210 0230 DL1EAP JN58Q 27 26 C
0250 0255 IZ0DXW JN65QO 27 27 C
0655 0710 DL1DWI JN75MT 27 26 C
0718 0722 SS1MQ JN75MT 27 26 C
0750 0806 OK1WCF JO80JU 27 26 C
0810 0830 IW1BCV JN44FS 27 26 NC 2079km
0830 0900 IZ4AK JN45PB 27 27 C 2013km 6sec
0907 0927 JN79LN 38 27 C 2sec
0930 0955 I8BCI JN7BKH 27 27 NC 2050km
0955 1005 D9BVB JO30JK 27 27 C
1054 1130 I8BCI JN7BKH 27 27 NC 2050km
1157 1202 DL9MCC JN88UA 27 26 C
1202 1245 IZ0DOR JN55EN 26 27 C
1250 1300 H45VO JN37NJ 27 27 NC
1328 1400 D9JYE JO43HV 27 27 NC
1450 1500 SP4MPB KO00GS 27 27 C
1503 1510 D9JYE JO43HV 27 27 NC
1750 1820 DM1CG JN37NJ 27 38 NC
1820 1840 H45VO JN37NJ 27 27 NC
1924 1931 G4LOH JN95UA 27 37 C v strong
1951 2002 DC9YC JO31P1 27 27 C
2031 2036 I43HWT JN55ST 27 27 C
2036 2100 D2JE JN48HV 38 27 NC
2100 2104 I8BCI JN7BKH 27 27 C 2050km
2115 2125 PE1L JN435X 27 27 C
2157 2201 D9RES JO435X 27 27 C
2201 2207 ON4KHG JO10XO 27 27 C
2219 2224 HA5CBA JN70TM 27 27 C
2229 2247 PA1T JO33JF 27 27 C
2240 2300 D2JE JN48HV 38 27 NC

Ops: Alex, RW3AZ, Harry, RA3AH and Vladimir, RA3ABZ

YZ7MON BCC Meteor Scatter Contest Log 2003

YZ7MON (KN05FG)

wkd on 144 MHz in FSK441:

2003-12-11

2003 2013 SM7DLZ 27 27 1055sec

2013 2021 DL3LST 27 26 10300

2021 2031 F1EBK 27 26 825sec

2048 2053 PE1AHX 26 27 10360

2053 2058 D9NYC 26 27 1015sec

2108 2119 SP2JON 26 26 1015sec

2123 2129 LA0BY 26 27 9280

2129 2137 PA0BAT 27 27 9280

2159 2166 DGLY 27 27 927sec

2147 2155 PE1HX 27 27 9800

2228 2247 SM7FAX 26 27 11640

2003-12-12

0007 0014 DL1G5T 27 26 10760

0014 0021 ON4KHG 26 26 6200

0054 0100 PA1BWM 26 27 11380

0100 0122 DL4EBV 27 27 10380

0130 0141 D9JYE 27 27 102sec

0203 0215 DLOAO 27 27 172sec

0404 0416 PE1HWO 27 27 10480

0519 0544 DL1RHW 27 27 10520

0613 0617 LZ2FO 26 27 102sec

Meteor Shower Calendar

by Guido Jünkersfeld, DL8EBW

April to June 2004

Name	Date	Max.	ZHR	RA	DEC	v (km/s)	Source
ursa-Majorids	01.04.-02.04.	01.04.	20				BMS50
beta-Pavonids	11.03.-16.04.	07.04.	13	308	-63		5IY97
April-Piscids	08.04.-29.04.	20.04.	low	007	+07		IMO04
Lyrids	16.04.-25.04.	23.04.	18	271	+34	49	IMO04 *1
pi-Puppids	15.04.-28.04.	24.04.	var.	110	-45	18	IMO04
delta-Piscids	24.04.-24.04.	24.04.	low	011	+12		IMO04
alpha-Scorpiids	16.03.-12.05.	03.05.	10	240	-27		5IY97
eta-Aquariden	19.04.-28.05.	05.05.	60	338	-01	66	IMO04 *2
Halleids	03.05.-21.05.	08.05.	25	338	-01		BMS50 *2
Piscids	06.05.-10.05.	08.05.	30	012	+19		BMS50 *2
epsilon-Arietids	24.04.-27.05.	09.05.	low	044	+21		IMO04 *2
nu-Piscids	11.05.-14.05.	12.05.	20	016	+27		BMS50 *2
May-Arietids	04.05.-06.06.	16.05.	low	037	+18		IMO04
Sagittarids	15.04.-15.07.	19.05.	5	247	-22	30	IMO04
o-Cetids	05.05.-02.06.	20.05.	14	028	-04		IMO04
nu-Scorpiids	19.05.-20.06.	25.05.	9	261	-37		BMS50
eta-Pegasids	20.05.-09.06.	31.05.	9	333	+27		BMS50
alpha-Scorpiids	01.06.-17.06.	05.06.	15	251	-22		BMS50
lamb.-Sagittariids	04.05.-09.07.	8.-15.06.	12	275	-25		BMS50
Arietids	22.05.-02.07.	07.06.	60	044	+24	37	IMO04 *3
zeta-Perseids	20.05.-05.07.	09.06.	40	062	+23	29	IMO04 *3
Ophiuchids	19.05.-02.07.	10.06.	20	269	-23		BMS50
June-Lyrids	11.06.-21.06.	16.06.	5	287	+35		5IY97
gamma-Draconids	.06- .09	17.06.	4	269	+48		BMS50
Lyrids	12.06.-21.06.	17.06.	8	282	+44		BMS50
Taurids-Perseids	25.06.-20.06.	18.06.	10	054	+31		BMS50
Sculptorids	10.06.-27.06.	25.06.	8	005	-35		BMS50
54-Perseids	22.06.-30.06.	26.06.	30	068	+34		BMS50
Corvids	25.06.-03.07.	26.06.	13	192	-19		BMS50
beta-Taurids	05.06.-17.07.	28.06.	24	086	+19		IMO04
Sagittariids	07.06.-18.07.	30.06.	8	285	-25		BMS50

Kommentar / Comments: *1 Der Lyriden-Schauer ist sicher als der erste gute MS-Schauer nach 3 Monaten MS-Abstinenz zu beurteilen - wie schon in den letzten Jahren, wird aber auch dieses Jahr nicht mit einem Großereignis gerechnet und so wird es für HSCW/FSK- aber nicht für SSB MS reichen. - The first good shower after a 3 month gap, but like in the last years do not expect a big event. Probably good for HSCW and FSK but not enough for SSB-MS.

*2 Der Zeitraum 3.-12. Mai ist für Meteorscatterfreunde und dies speziell in den südlichen Breiten mit einer der guten Zeiten im Jahr. Tatsächlich können in den eta-Aquariden gleiche Weiten erzielt werden wie bei den berühmten Leoniden im November! - The time frame from May 3rd to 12th is quite nice for the MS enthusiasts especially in the more southern latitudes. Distances like in the famous Leonides in November could be reached.

*3 Obwohl es im Zeitraum Anfang Juni nicht soweit geht, bestechen diese Schauer aber durch die Häufigkeit der Reflexionen - teilweise kommt es durch die Überlappung von Radianten bis zu mehr als 60-80 Meteoriten pro Minute (ZHR). Die Reichweite ist aber begrenzt auf unter 2000km. - Max. QRBs are lower (below 2000 km) but plenty of reflections are possible due to the overlap of radiants from several good showers. (60-80 reflections per hour).

Legend / Erklärungen: see DUBUS 4/2003, page 89

Aurora Reports

Editor: Stefan Heck, LA0BY
LA0BY@darc.de

Editors note:

This time we have chosen to add one 50 MHz report. But normally we will not include 50 MHz unless the reported contacts are really very special. That is from very Southern locations or over unusually large distances.

73, Stefan (LA0BY)

50 MHz reports

IW0GPN in JN62FB wkd on 50 MHz:

2003-11-20

1900	G0JHC	IO83PQ		10	1713 km
1908	G0JHC	IO83PQ	SSB	10	1713 km
1914	M0DDT	IO91	SSB	10	1464 km

tnx for info, Alex

144 MHz reports

9A2SB in JN95 wkd on 144 MHz:

2003-11-20

1555	G4RRA	IO80BS	57A	55A	340	1755km
1605	SP2OFW	JO93AC	59A	59A	5	845km
1608	DK8ZJ	JO54AG	59A	52A	5	1147km
1612	SP4JWD	KO03FR	59A	55A	10	924km
1632	SP4MPB	KO03MT	57A	52A	350	940km
1638	OZ6OL	JO65DJ	59A	55A	345	1184km
1845	PC5C	JO33GD	57A	55A	350	1212km
1850	PA5KM	JO11UL	59A	55A	350	1275km

Rig: IC202+100W, 10el. DL6WU

tnx for info, Zlatko

DG2BCP in JO43DC wkd on 144 MHz:

2003-11-20

1625	G8ISY	IO91VC	59A	59A	621 km
1628	G1KHX	IO81MI	59A	59A	789 km
1629	G3EDD	JO02DE	59A	58A	549 km
1631	ON6MG	JO10VK	59A	59A	428 km
1632	PA3GFY	JO32SB	59A	59A	126 km
1633	G3BNE	JO01BJ	59A	59A	587 km
1635	ON4SN	JO21FB	59A	59A	346 km
1637	DO9OAM	JO41XM	59A	59A	209 km
1640	DK1FUN	JN49MR	59A	59A	379 km
1641	DD6BKR	JO72CM	59A	59A	403 km
1644	F5RMK	IN98OJ	59A	59A	825 km
1646	OE9NPI	JN47UJ	59A	59A	642 km
1650	F3BKM	IN98ME	59A	59A	849 km
1653	S53J	JN75EV	55A	55A	913 km
1656	F6CGS	JN18FW	55A	55A	617 km
1658	G4RUW	IO91IJ	59A	59A	679 km
1700	OE3FVU	JN78VE	58A	59A	759 km

1704	F5FT	JN18DO	59A	59A	653 km
1709	G4XG	IO93XE	55A	55A	555 km
1710	G4VPD	IO92BJ	57A	59A	688 km
1712	M0CKM	IO90QT	59A	59A	661 km
1713	G8RSA	JO01BO	59A	59A	578 km
1715	HB9KAI	JN47CG	59A	59A	648 km
1717	OK1CID	JO80CA	55A	57A	645 km
1720	F6EBH	JN19WT	59A	59A	477 km
1722	GW8ASA	IO81EM	59A	59A	828 km
1721	SM7XWX	JO69MO	59A	59A	779 km
1724	F1EBE	JN09NN	57A	59A	634 km
1726	F1ORL	JN08VP	59A	59A	673 km
1727	F8BRK	JN99VF	59A	59A	908 km
1729	F1BZG	JN07VU	58A	56A	742 km
1730	SM6OPW	JO58RG	57A	59A	607 km
1732	G0VOK	IO83RG	53A	53A	721 km
1737	F1BQX	JN18DP	58A	57A	650 km
1739	GW8ASD	IO83LB	57A	57A	756 km
1751	DD1UN	JN49RO	55A	55A	397 km
1754	DC6NY	JN59MI	57A	57A	459 km
1757	G7KTP	IO92RB	55A	53A	607 km
1800	HB9BQ	JN37VN	53A	55A	617 km
1806	F1ORH	JN19BK	53A	55A	591 km
1807	PE1BT	JO22KL	58A	55A	240 km
1809	F6IRS	JN38VN	58A	57A	506 km
1813	F4AJQ	JN18QM	59A	59A	615 km
1840	SP9JDP	JN99HW	57A	55A	796 km
1855	EI2AI	IO51PQ	55A	55A	1161 km
1857	G4YVP	IO93VI	55A	55A	566 km
1918	IW2NOR	JN45ON	57A	59A	841 km

RIG: TS-2000, GS35B, 4 x 3,6 WL DJ9BV, all QSO in SSB
tnx for report, Peter

DJ8MS in JO63CT wkd on 144 MHz:

2003-11-20

1409	OH5LK	KP30ON	56a	55a	1170 km
1840	HA1YA	JN87GF	55a	55a	793 km
1840	IV3DXW	JN65QQ	55a	56a	907 km
1851	S54G	JN65TM	57a	57a	927 km
1855	S51DI	JN76VL	53a	55a	854 km
1859	HB9DFG	JN37SM	52a	57a	772 km
1901	OK2PTC	JN89JM	52a	59a	572 km
1903	HA7WJ	JN97MP	55a	55a	836 km
1905	9A3JH	JN75AD	53a	57a	972 km
1909	OK2AF	JN89AR	53a	55a	525 km
1910	S57Q	JN76PB	57a	59a	889 km

tnx for info, Toralf

DK8ZJ in JO54AG wkd on 144 MHz:

2003-10-29

2047	SK0UX	JO99bm	55A	55A	30
2052	OZ6EI	JO45tt	55A	55A	30
2058	SK6DK	JO67eh	55A	57A	30
2110	SK6BA	JO64hm	55A	59A	SSB
2120	LA6MV	JO59vi	59A	59A	SSB
2130	SK6HD	JO68sd	57A	59A	30
2140	SP2MKO	JO93cb	55A	55A	30
2142	SM5DFF	JO88cn	57A	57A	30
2207	LY2BAW	KO25ca	53A	53A	30

220	SP4JWD	K003H	53A	55A	30	1446	SSAM	JN86d	52A	53A	80	2131	H43UU	JN95p	55A	56A	80	1634	SP2FA	J094MA	54A	55A
223	SP9AL	JN98nt	52A	57A	30	1454	UREXJ	K070K	59A	59A	80	2134	H45NF	JN97rw	55A	55A	80	1635	G3EDD	J002DE	59A	59A
224	OK2WJ	JN98q	52A	57A	30	1456	OK1M4Z	J070A	55A	55A	80	2142	OK2BEJ	JN97j	52A	53A	80	1654	F1PUP	JN15XL	57A	57A
242	PA5KM	J011w	55A	58A	30	1503	SS4G	JN68m	53A	55A	80	2156	GRAU	J094f	59A	59A	30	1655	OK2VSO	JN98XK	53A	57A
243	DJ8MS	J033a	57A	59A	30	1508	OK2PQI	JN68h	53A	55A	80	2206	SM7SPG	J066nd	59A	59A	30	1656	OK2BTRA	JN98ET	57A	59A
243	OK1CZ	J070ee	59A	59A	30	1510	PA2PVC	J031gw	57A	57A	80	2206	SK6DH	J067hw	59A	59A	30	1658	SP6T2	J090BK	57A	59A
238	HB9QD	JN47H	59A	59A	30	1512	RAX3R	K085p	52A	57A	80	2213	SP6LC	J070w	55A	57A	80	1659	F2LQ	JN68EB	59A	58A
313	OK2NY	JP20cm	52A	59A	30	1516	YL2PA	K028ba	52A	59A	80	2216	DF2CZ	J030m	55A	57A	80	1701	OK2PZA	JN68CS	57A	57A
3003-10-30						1518	YL2CQ	K017em	57A	59A	80	2219	OK2PXL	JN68pc	55A	56A	80	1703	US5WQD	KN19SG	45A	59A
800	LA8WF	J059p	55A	57A	30	1521	SM1NC	K097de	59A	59A	80	2222	OK5BK	JN68q	53A	55A	80	1708	OK3PLW	JN78NV	45A	59A
805	LA7VY	J059q	53A	57A	30	1523	OK1FPR	J080ee	59A	59A	80	2226	OK2MCM	JN68r	53A	55A	80	1710	F5FT	JN78VJ	59A	59A
813	SMUKAK	J089XK	57A	58A	30	1528	US4XAC	K085nc	52A	59A	80	2238	DL9MS	J054wc	53A	55A	80	1712	F1DUJ	JN074J	59A	59A
817	SM6FV	J078r	55A	55A	30	1530	PA3YD	J030as	57A	57A	80	2240	DL9MS	J054wc	53A	55A	80	1713	SP8BCJ	K012NC	57A	59A
906	PA3AYD	J021m	53A	55A	30	1533	OK3WQ	J030as	55A	59A	80	2255	G4M4VX	J078a	55A	57A	80	1716	K4DRY	JN64BL	59A	59A
911	DL1EJA	J013db	55A	55A	30	1537	OK1CJ	J070ob	59A	59A	80	RIG-4x12el; 750 Watt. tx for report, Martin						1717	F18OX	JN61BP	59A	59A
957	LA4QC	JP20p	55A	55A	30	1540	OK1M2M	JN68p	59A	59A	80							1718	DL6FH	JN68AK	59A	59A
958	OH1L	JP00ox	55A	57A	30	1541	HA5AB	JN67m	55A	57A	80							1720	DF3DL	JN68TB	59A	59A
903	DL9MCC	JN57??	55A	53A	30	1543	OK2ZW	JN68j	55A	59A	80							1723	F10OC	JN68QJ	59A	59A
1005	OH1NOR	VP00A	58A	59A	30	1548	OK1YA	J0917r	59A	59A	80							1724	F1D1T	JN27KH	59A	59A
1011	M0KCEJ	J075m	53A	55A	30	1551	SP2LBI	J0917r	57A	59A	80							1726	F4CYH	JN26BH	59A	59A
1017	LA5EFP	J084W	55A	57A	30	1555	SP6RQB	J071af	53A	55A	80							1727	DG2KBC	JN58MI	57A	57A
1026	SP3EJ	J084W	55A	55A	30	1559	YL7UW	K095ee	53A	59A	80							1730	MI4HY	I083EV	55A	54A
1029	LA3EDA	JP90q	55A	57A	30	1602	SM2JL	JN75gm	52A	57A	80							1731	MI4HY	I083EV	55A	54A
1033	SM7EOC	J067p	53A	59A	30	1605	OK4NEZ	JN67cc	53A	55A	80							1815	F1DES	J020BV	59A	59A
1035	SM6DOK	J067p	53A	59A	30	1607	OK4NEZ	JN65kg	55A	59A	80							1816	DL4EBV	J031HK	59A	59A
1039	LA8BY	J098W	57A	55A	30	1609	PA2SB	JN65gm	52A	59A	80							1818	DL5ZGP	J051MV	59A	59A
1042	SM4RPP	J079H	55A	55A	30	1613	OK2PM	JN65gm	55A	55A	80							1819	DL7K7F	J031YA	55A	59A
1050	SM3UJF	JP90H	55A	56A	30	1616	SP6FLC	J090ea	55A	55A	80							1825	DC2NY	JN59MI	57A	56A
1114	OK1FID	J080a	53A	55A	30	1618	F8FLB	J000KX	55A	59A	80							1828	SP2CAS	J093BO	42A	42A
1123	SP3EJ	JN67em	55A	57A	30	1620	G3RDE	I082qe	57A	57A	80							1837	DL6YEH	I081XW	59A	59A
1125	HB9QDM	JN75r	55A	58A	30	1623	G3RDE	I082qe	57A	57A	80							1844	G4ERP	I081XW	55A	57A
1128	HB9QDM	JN75r	55A	58A	30	1625	PA1CMS	JN68fm	52A	55A	80							1847	SS3VJ	JN65VW	55A	56A
1150	DL7AJQ	J062mo	55A	53A	30	1629	OK3KZA	JN68fm	52A	55A	80							1853	OK2VMB	JN68WV	32A	53A
2006	SM4CLQ	J011se	55A	55A	30	1632	DL7GW	J062cc	55A	55A	80							1855	F5QWVR	K021IM	54A	57A
212	UTBAL	K061wp	55A	59A	30	1636	KA2DF	JN68p	59A	59A	80							1858	F5QWVR	K021IM	54A	57A
218	OE5XBL	JN68p	55A	55A	30	1638	PA3TJ	JN68p	55A	59A	80							1859	F5QWVR	K021IM	54A	57A
219	OK1YVM	JN67cc	59A	59A	80	1639	OK5AL	JN67cc	59A	59A	80											
222	PA2M	J021p	59A	57A	30	1641	OK1YVM	JN79q	59A	59A	80											
226	OK4KHG	J010m	59A	55A	30	1643	OK2AF	JN68ar	59A	59A	80											
231	GP0RC	I093pa	55A	55A	30	1644	OK1WV	J070b	59A	59A	80											
238	SP4JWD	K003H	55A	54A	30	1646	HA5GK	J097th	57A	57A	80											
3003-11-17						1648	HA5GK	JN68q	55A	59A	80											
702	SM8BOS	J098q	55A	55A	10	1650	HA6NN	JN68q	55A	59A	80											
1000-11-20						1650	HA6NN	JN68q	55A	59A	80											
358	SM7WVG	J087p	59A	57A	80	1652	HA5MO	JN67q	55A	59A	80											
359	SM7WVG	J087p	59A	59A	80	1654	OK2BFI	JN68q	55A	59A	80											
402	YL2BJ	K025er	59A	59A	80	1655	HA5MO	JN67q	55A	59A	80											
410	DL6JQ	JN49H	53A	53A	80	1655	HA5MO	JN67q	55A	59A	80											
416	DL6JQ	JN49H	53A	53A	80	1655	HA5MO	JN67q	55A	59A	80											
420	RV3PF	K093cd	55A	52A	80	1657	LZ7FO	KN13xx	55A	57A	80											
426	SP4MPP	K003H	55A	52A	80	1700	Y08DQD	JN68p	55A	57A	80											
440	DL7BU	J051ut	55A	57A	80	1702	OK3TZO	JN68p	55A	59A	80											
424	DL7BU	J051ut	55A	57A	80	1706	OK2KOS	JN69d	59A	59A	80											
427	DL7BU	J051ut	55A	57A	80	1708	SP6FG	JN69d	59A	59A	80											
433	YL2BJQ	K015xh	59A	59A	80	1711	OK2SA	JN69d	55A	59A	80											
430	R3DRJC	K055f	55A	59A	80	2058	OE5XBL	JN68q	55A	59A	80											
436	OK1VHF	J078ba	53A	55A	80	2100	DF2JQ	J031H	59A	59A	80											
438	HA6NQ	JN68q	53A	57A	80	2103	DF7KQ	J030p	59A	55A	80											
440	OK3GTZ	JN68i	55A	59A	80	2122	SS1ZO	JN68p	59A	57A	80											
442	UTBAL	K061wp	55A	59A	80	2125	HA5GK	JN67q	55A	55A	80											
444	DK3SWD	J062n	55A	57A	80	2129	OM5ZW	JN68i	55A	55A	80											

DL1SUN in JO35PN wkd on 144 MHz:

2003-11-20

1245 DF2JQ J031H 57a 57a

1251 L2BJ J025ER 55a 57a

1254 SP2RS J094BK 55a 53a

1302 PA4PS J093CD 55a 55a

1313 RA3JPF J021PM 55a 59a

1328 HA6NQ JN68W 55a 57a

1331 DF7K J030G 55a 59a

1340 SP4MPB J003HT 55a 56a

1403 SM7WT J065Q 55a 59a

1405 YL85FA J026W 59a 59a

1409 SM5EFP J079W 59a 58a

1434 DF1CF JN57FP 57a 57a

1455 OK1VHF J070EB 59a 59a

1458 UA3DHC K095AW 59a 59a

1505 YL2CQ K017EM 59a 59a

1507 MM4UF J092JP 55a 59a

1508 FDVGP JN19AJ 59a 59a

tx for report, Norbert

DL6VEH in JO32VA wkd on 144 MHz:

2003-11-20

1621 SP2RS J094GK 53a 55a

1623 G3ALEW J071PS 53a 55a

1643 SS1AT JN75GW 53a 836 km

1653 UX0CQ K096AK 51a 53a

1654 H1DMP JN35UB 57a 57a

1749 SS4Z JN65TM 53a 55a

1904 SS1ZO JN68D 57a 57a

1915 IZ4HEH JN54WL 55a 59a

1920 HA8V JN68HT 55a 59a

tx for info, Andreas

DL8BDU in JO43AA wkd on 144 MHz:

2003-11-20

1616 SS4M JN68CL 59A 54A

1620 SS3J JN75EV 59A 55A

1622 SM4BDQ J090FG 55A 57A

1624 HB9EZA JN38BE 57A 54A

1627 SS4G JN65TM 59A 57A

1923 SP8HIV KN00LA 55A 55A
1924 OESXBL JN68PC 55A 55A
1929 CS1ZD JN68DR 55A 55A
1932 OK1VHF J071EB 55A 55A
1937 SP8A J081MC 55A 55A
1938 OK2KOS JN69ST 55A 55A
RIG: FT847 + 2 x CX500 + 2x110 - MGF 3002. Inx for info

F0E/p in IN78QG wkd on 144 MHz:

2003-11-20
1711 OK1DCX J080BJ 1516 km
1716 OK4HKG J010XQ 482 km
1723 E15FK J051RT 672 km
1732 DESKE JN78DF 1396 km
1738 G3KEQ J011BH 480 km
1747 OZ7OL J065DJ 1440 km
1756 SP4JWD J063JR 1845 km
1800 P4SKM J011WL 705 km
1808 E2H1 J051WQ 455 km
1819 DG9YH J032QJ 963 km
1822 DF7DF J030GU 855 km
1850 DL2DX J061VC 1357 km
1852 SP7OGK K001AM 1793 km
1859 PA2M J021JP 770 km
1904 OK1WVK J0700A 1450 km
1907 SS4M JN68CL 1575 km
1911 H48V JN68HT 1892 km
1923 OESXBL JN68PC 1325 km
1926 F5HRY JN18EQ 517 km
2230 GM4BY J085JV 853 km
Inx for report

G3XDY in J0120B wkd on 144 MHz:

2003-11-20
1923 H48V JN68HT 1514 km
1927 SS1ZD JN68 1239 km
2114 J2FAK JN45 974 km
Inx for report, John

G4ASR in IO81MX wkd on 144 MHz:

2003-11-20
1917 H45CV JN97NJ 59A 080° 1661km
1918 H45CV JN97GO 57A 59A 080° 1611km
1919 IV4DWM JN65QO 57A 57A 080° 1381km
1920 IK4PMB JN54LL 57A 59A 080° 1322km
1922 DF832A JN38L 57A 57A 080° 745km
1924 HB982A JN36 57A 57A 080° 945km
1925 IK4DRY JN64BF 55A 55A 080° 1409km
1925 IK4DPM JN35UB 55A 55A 080° 1097km
1927 F5VHX JN04FT 57A 59A 080° 1130km
1928 DJ7GK JN59WH 57A 55A 080° 1223km
1930 IK3XLO JN58RG 55A 55A 080° 1037km
1931 DO3MK JN58JP 52A 55A 080° 1037km
1934 HB80KM JN37SL 57A 57A 080° 944km
1935 SS1MQ JN75NT 57A 57A 080° 1464km
1938 DK7FK JN49H1 55A 53A 080° 866km
1940 OM3GTZ JN69T1 55A 55A 080° 1611km
1942 DL4DWA J061QH 53A 53A 080° 1127km
1946 FB8EG JN25J5 55A 55A 075° 892km

2220 DL9MS J041VC 55A 55A 10 676 km
2240 OK1VIT JN70XV 55A 55A 30 309 km
2243 DF22C J030RN 55A 55A 30 418 km
2259 DJ9CZ J031BC 55A 55A 30 528 km
2316 DF23Q J031HL 55A 55A 30 530 km
2332 DL7FF J062J1 57A 55A 30 505 km
2338 PA0MW J027PM 52A 55A 40 603km
2341 DL8YE J031MH 52A 52A 40 496 km
Inx for report, Hans

F1F1H in JN23GS wkd on 144 MHz:

2003-11-20
1724 DM1GQ J0311C 59A 59A 15 SSB
1725 ON6ZK J020KW 59A 59A 15 SSB
1730 G1PCK J091V 59A 59A 15 SSB
1740 ON7CL J020U 59A 59A 15 SSB
1743 GR8ZY J091VC 59A 59A 15 SSB
1848 PE1AHX J021OS 59A 59A 15 SSB
1857 ON4JL J020V 59A 59A 15 SSB
1902 G8HGN J010FO 59A 59A 15 SSB
1911 G8WLY J010M 59A 59A 15 SSB
1933 ON2AE J010W 59A 59A 15 SSB
1935 ON4MM J011U 59A 59A 15 SSB
1937 DK5WJ J030AS 59A 59A 15 SSB
1942 ON1AEN J010W 59A 59A 15 SSB
Inx for report, Michel

F5SD in JN23RF wkd on 144 MHz:

2003-11-20
1733 M8BXH J091XR 55A 55A SSB 1031km
1912 F6HHR J010AV 55A 55A 890km
1932 EHQD J051WU 55A 55A 1398km
1941 DF22C J030RN 55A 55A 829km
1944 ON1AEN J010M 52A 55A SSB 820km
1954 DK5DQ J010PH 55A 55A 905km
2123 DL1RL J032PH 55A 55A 1167km
2125 DF847 N.C. 55A 55A
RIG: IZ21TE, 11alls F9FT, PA 2x150 (400W)
Inx for report, Dave

F6CRP in IN96KE wkd on 144 MHz:

2003-11-20
1637 DF50E J040CX 55A 53A
1641 LY2IC K014MW 55A 55A
1644 SP4MPB J030HT 55A 59A
1647 SP4BY J031OD 55A 55A
1655 SP2WJ J093AC 55A 57A
1656 MCVDR J083LV 57A 59A
1658 M0VDR J083LV 55A 59A
1659 DL2AF J050W 55A 55A
1701 OK2PM JN9940 51A 55A
1703 PA1LA J032LS 59A 57A
1745 H48V JN68HT 59A 59A
1749 E8GQ J051TV 59A 59A
1754 FDDO JN26F 54A 55A
1756 GAUF J033F 55A 57A
1757 SG9YH J030Z 56A 53A
1807 SP4UW J030FR 56A 56A
1848 OK1DCF J080BJ 55A 59A

2323 ON4PS J020KQ 53A 57A 30°
2324 LY2BL J015XV 53A 59A 40°
2327 HB9DF JN75M 51A 55A 40°
2329 HB80KM JN73L 51A 55A 40°
2330 RV3WR JN70LM 59A 59A 40°
2331 JH3GW JN83VN 51A 53A 40°
2332 RA3GDK J081BL 51A 59A 40°
2333 H45CV J081BL 51A 59A 40°
2339 GA4EP J091J 53A 55A 40°
2345 G4MTH J092J 51A 55A 40°
2350 LY2BAW K025KA 55A 59A 40°
2003-11-01
0007 URLX J070WK 51A 55A 60°
0009 DJ5J J042HA 51A 55A 40°
0013 SMTCOA J097JN 52A 55A 40°
0017 OK2TT JN68KU 51A 55A 60°
0018 DL1ALB J050KW 52A 57A 60°
0036 H48V JN75M 52A 57A 80°
0038 H48V JN68HT 55A 59A 80°
0038 S5TWW JN75EX 53A 55A 80°
0039 URSXK JN28EW 51A 55A 70°
0043 OK1L J070WN 51A 59A 70°
0046 OK1FD J080Q 55A 59A 80°
0047 SPLTC J070WV 53A 55A 70°
0050 DL2AN JN48BC 52A 53A 70°
0052 G3PQ J010BU 55A 57A 30°
0054 G3BEL J080Q 55A 59A 60°
0056 YU8G K050FT 51A 59A 60°
0057 SP8S K020MF 51A 55A 60°
0100 DO3SP J072EN 51A 59A 50°
0101 SPLB J070U 54A 55A 60°
0106 HB9BZ JN68BE 55A 57A 70°
0123 H48V JN68W 55A 57A 70°
RIG: TS-890S, LT2S, 300W, 17el F9FT
Inx for report, Norbert

DL3WCC in JN58UA wkd on 144 MHz:

2003-11-20
1901 DK2AR J042UA 57A 58A 340 467 km
1906 E2H1 J051Q 52A 57A 300 1547 km
1911 F60Z JN19NE 55A 55A 330 644 km
1922 GA2PU JN15UE 51A 55A 330 393 km
1924 DL1ANA J050TV 52A 57A 330 324 km
1927 G3PW J083PM 53A 57A 330 1180 km
1929 OK2SBL JN69HQ 53A 55A 330 539 km
1933 G7RAU J090R 57A 59A 330 998 km
1939 DJ7EB JN68GD 51A 54A 330 63 km
1944 OK1TEH J070FD 53A 55A 330 310 km
1948 OK2BEJ JN79H 51A 55A 330 263 km
2047 G6ATZ J074AJ 59A 59A SSB
2055 G4OVC J064X 53A 55A 330 1518 km
2059 G4OVC J093UK 53A 55A 330 1035 km
2108 E15FK J051RT 53A 55A 330 1504 km
2113 J0DPU J032LS 55A 59A 330 626 km
2116 GARRA J091F 57A 57A 330 967 km
2120 H48V JN68L 53A 59A 330 572 km
2122 DF81K J030J 57A 59A 330 424 km
2127 DO3MA JN68P 55A 59A 330 71 km
2204 OZ2ZS J055RT 55A 52A 10 865 km
2213 OZ1CL J065Q 55A 57A 10 860 km

217 F5EPJ JN19XV 53A 56A 40°
235 SP4JWD J063JR 52A 55A 70°
251 RA3AO K08SSP 52A 57A 40°
255 SS7ZO JN68DR 52A 57A 60°
257 DL2JN JN48C 51A 57A 60°
300 DL5GAC JN47UJ 52A 55A 70°
301 HB9DF JN75M 53A 56A 70°
303 DL1GTG JN58A 51A 53A 70°
304 HA1BC JN87PU 51A 53A 70°
305 FL3L JN26MV 51A 55A 45°
308 PA3JCN J033AH 52A 57A 30°
313 SNEEFT J079WV 53A 57A 30°
332 LA8Y J059FW 51A 57A 20°
344 SM5FNU J080CR 52A 55A 30°
348 SM0LQ J089XV 51A 55A 30°
351 G3WZT I090 52A 55A 35°
352 G4RCK J091ON 52A 55A 30°
354 SM0WV J089VL 52A 55A 30°
357 YL3DQ K026CW 52A 55A 40°
2003-10-30
002 OK2PM JN98A0 51A 55A 60°
004 OK2BDS JN79HF 51A 55A 60°
009 YU7EY JN79HF 51A 55A 80°
112 HA1AD JN87RT 41A 57A 80°
115 H45CV JN75PM 51A 58A 80°
116 H45CV JN75PM 51A 58A 80°
118 H45CV JN75PM 51A 58A 80°
126 SS1ZD JN68DR 52A 55A 80°
155 RV3WR J071M 57A 59A 60°
159 SP4BY K013OD 52A 53A 60°
02 SP6GV J081LC 52A 59A 60°
04 YU8G K050FE 41A 55A 60°
06 ON7CMK JN68NR 41A 55A 60°
11 OK1ZHS J070VA 51A 57A 60°
12 ON5CM JN68MD 51A 57A 60°
16 HA7WJ JN75PM 41A 56A 70°
22 URSX JN70WK 54A 59A 60°
23 H48V JN75U 51A 59A 60°
28 SP4SLO J091BS 41A 55A 60°
37 CE15OW JN68F 53A 55A 60°
40 OK2YJ JN68W 51A 55A 60°
42 OK1MHZ J070VA 51A 55A 60°
57 E5SPC K038AJ 52A 57A 30°
59 RX1AS J089FX 53A 58A 30°
15 SM3BEI J091JP 54A 57A 0°
03-10-31
09 DL6YEH J031TX 53A 55A
19 DHTFB J0625M 53A 52A
45 HALCK JN54PD 52A 59A 80°
43 LY2BJ K015NV 41A 57A
10 UTRAL K061WP 55A 59A 40°
16 DL8KR J030NQ 51A 56A
18 YL2HA J030J 51A 57A
19 ON4CL J011SE 53A 55A 30°
21 G48BK J091DP 51A 53A 30°
23 DK2VA JN38L 52A 57A 40°

1946 F6G1H JN19EV 55A 54A 080° 444km
1953 HA8V KN06HT 55A 57A 080° 1792km
2016 EH4D IO51WU 55A 55A 035° 355km
2105 DP2JQ JO31HL 53A 55A 065° 862km
2110 OK2BSL JN99HQ 53A 55A 060° 1511km
2110 SP6WQ JO80U 55A 59A 060° 1378km
2115 OK2BSH JN89W 55A 55A 060° 1471km
2116 OK1TEH JO70DB 55A 59A 060° 1231km
2118 OK1TEH JO70DB 55A 59A 060° 1231km
2120 F8CWU JN07JK 52A 55A 050° 578km
2125 FR8FW N941TR 59A 59A 050° 828km
2220 DL7JK JO41SX 55A 59A 060° 875km
2223 FG6JN JO01ST 53A 55A 060° 875km
2235 SK6DK JO67EH 59A 57A 020° 1147km
2238 G4MVE JO68FR 59A 55A 035° 531km
2242 G4MXX JO78TA 55A 55A 010° 678km
tx for report, David

G4KXLX in IO93FB wkd on 144 MHz:

2003-10-29
1810 DL8BDU JO43AA 53A 52A
1820 E5FK IO51RT 53A 53A
1834 OK1YV JN79JX 54A 54A
1841 OK1YV JN79JX 55A 55A
1856 OE5BLX JN67AW 54A 55A
1913 OK3MRM JN87WV 57A 57A
2003-11-20
1745 OK1DCF JO80BJ 57A 57A
1809 OK1VHF JO70EB 56A 55A
1818 IO2XW JN50DQ 55A 56A
1834 9A1EZA JN86AF 57A 56A
1851 DL2BHG JO51MW 54A 52A
1854 FL8UX JN90OK 58A 57A
1856 F3HRY JN18EC 59A 59A
2218 OK2PXX JN69AJ 55A 55A
Rig: Iq-910H, 100W, 9-ale F9F1, tx for report, Jonathan

G4RRA in IO80BS wkd on 144 MHz:

2003-11-20
1541 LY2BJ KO25ER 53A 55A 1943 km
1541 LY2CJ KO14MW 57A 57A 1906 km
1546 HA0HQ KN07SU 59A 59A 1861 km
1622 OH1FA KP10PL 55A 55A 1996 km
1703 EU3AJ KO22DE 55A 56A 1944 km
1718 HA8BR KN06PJ 57A 57A 1851 km
1744 HA8V KN06HT 59A 57A 1838 km
1945 Y02S KN06PS 55A 55A 1932 km
tx for info, Paul

3AXBF in IO91QD wkd on 144 MHz:

2003-11-20
605 LY2BJ KO25ER 1721 km
612 ES6RG KO28WA 1855 km
632 LY2C KO14MW 1681 km
642 EU3AJ KO22CE 1708 km
715 HA8NA JN98VC 1503 km
831 USSWU KO20D1 1746 km
834 SP9HV KO00LA 1523 km
938 YU7EW KN05HP 1676 km

2102 SP9GJH KO02RU 1517 km
tx for report, Mike

G8HGN in JO01FO wkd on 144 MHz:

2003-11-20
1507 G4T0E IO87HF 52A 59A 15
1516 PD0RC JO33AD 55A 51A 25
1517 PE0PK JO27JL 55A 59A 25
1519 G4MUR JO27JL 57A 58A 25
1525 O21ZD JO46H 55A 55A 25
1527 PA0LW JO22NS 56A 56A 25
1529 PA0HFH JO22SA 56A 59A 25
1536 FIDES JO20BE 54A 57A 30
1541 DK3WO JO30AS 59A 59A 45
1543 DG0GAM JO41XM 59A 59A 45
1544 DG0NA JO63UX 57A 57A 45
1545 G4Y1R JO42NB 57A 57A 45
1546 DL1MP JO31CH 55A 53A 45
1550 DJ9J JO440J 57A 58A 50
1554 FC1P IN97DD 57A 58A 45
1556 PD0RT JO23RE 57A 59A 45
1557 DB8NT JO50VJ 56A 56A 45
1559 D02MN JO41EQ 54A 55A 45
1605 D06OM JO62QD 55A 57A 45
1606 ON6G JO10VX 55A 55A 45
1609 DG1A JO30VJ 57A 59A 45
1610 DG6AO JO42WQ 59A 57A 70
1615 SP9NCJ KO12NC 55A 55A 55
1617 DB8C JO31BH 56A 57A 55
1618 PA0GY JO23SH 59A 59A 55
1620 DL1RW JO62GH 56A 59A 55
1622 DK5TX JO41WJ 55A 55A 55
1623 DL8YU JO41HS 57A 59A 55
1624 DG00BU JO50WV 57A 57A 60
1630 DG2BCA JO42AD 56A 57A 60
1645 DG5CT JO60DS 57A 57A 60
1647 HA6ZB KN07BV 57A 57A 65
1648 SP7CN JO91QD 57A 57A 65
1651 DL1RM JO62HD 56A 56A 65
1653 DH2AK JO41SM 57A 59A 65
1655 OK1KAO JN86KL 54A 54A 65
1656 DF2WE JO31GO 59A 57A 65
1702 DL2IAN JN49BC 59A 59A 65
1703 G2F2AJ JN49FS 59A 59A 65
1706 DL8EV JO31HK 49A 59A 65
1710 F4DXX IN87LH 55A 55A 65
1715 G4FQD JO40HM 59A 59A 65
1716 DG2BCG JN86J 57A 59A 65
1718 K1EGC JN86J 59A 59A 65
1719 DL4VIA JO61QH 56A 59A 65
1720 DG1NSE JN98NO 55A 59A 65
1738 F6XI JN19BB 55A 59A 65
1815 OE5OL JN86SI 56A 56A 65
1821 G4USY IN89EJ 57A 59A 65
1832 DL5ET JO41EV 56A 59A 65
1834 DL5ET JO41EV 56A 59A 65
1835 D01LK JO53CQ 55A 57A 55
1837 DL3YBZ JO33KU 56A 56A 55
1902 F1FH JN23GS 56A 59A 75

1903 DF7AP JO51FI 57A 59A 75
1904 F1JRD JN13WO 57A 55A 75
1906 FO2XO JN28PP 56A 55A 75
1908 D01NF JN89AM 56A 59A 75
1910 OK2PMX JN88W 54A 59A 75
1912 DH8NB JN49WS 56A 59A 75
1913 OK2YO JN89XK 56A 59A 75
1915 SP8HD JO80K 54A 59A 75
1916 D01WKS JN8WK 57A 59A 75
1921 DB1EPO JO31HG 54A 57A 65
1923 F4CHY JN28BH 57A 59A 70
1924 S1MJC JN75NT 57A 59A 70
1925 F1BBL JN06PP 54A 55A 75
RIG: FT-447, 150 W, 2 x 15-ale, all contacts in SSB
tx for info, Bob

G8IYV in IO91VC wkd on 144 MHz:

2003-11-20
1559 DJ2VD JO31PF 59A 59A 60
1600 PE1AHX JO21OS 59A 59A
1604 M0CIN IO75OP 59A 59A
1607 PD2JVE JO22VJ 59A 59A
1616 DB8WV JO33TH 59A 59A 60
1618 PE2DB JO22MD 59A 59A
1627 DG2BCP JO43DC 59A 59A
1630 DK3EC JO41GU 59A 59A
1636 DG5CST JO80DS 59A 59A 70
1639 ON7CL JO30JU 59A 59A
1646 DB8KJ JO30CM 59A 59A
1656 D01FOX JO31PU 59A 59A
1700 DK9TF JO31NF 59A 59A
1709 G4USY JN89XE 59A 59A
1715 DK9AT JO43XO 57A 58A
1720 HA6ZB KN07BV 59A 55A 75
1722 DL5EY JO41EV 59A 59A
1730 F1JRD JN13WO 59A 59A 70
1743 F1FH JN23GS 55A 57A
1746 PA0LW JO22NS 57A 57A
1748 SP9JUD JN89W 57A NC
1752 DB4ET JO31GO 55A 57A
1752 PD0RT JO23RE 59A 57A
1753 DK7LZ JO30MR 57A 57A
1756 F6EET JN18AM 55A 55A
1800 D090AM JO41XM 55A 55A
1801 F0KKT JN18H 57A 55A
1802 FIDES JO20BE 59A 55A
1804 F1EBK JN18BW 55A 55A
1805 DB8KH JO31FO 59A 57A
1807 SP9XMU KO02LG 55A 55A
1809 SP9PMP JO00FB 57A 55A
1810 DG2DAA JO62GU 59A 59A
1811 F4DSD JN23JO 59A 55A
1814 HA8U JN75SA 59A 59A
1814 HA8V KN06HT 59A 59A
1816 DG8AK JO30AV 59A 59A
1818 F1BGNP JN08K 59A 55A
1820 SP9MYC JO90OG 59A 57A
1823 IW4BET JN54OL 55A 57A
1825 F1EYB JN23 55A NC

GM0CLN in IO85HV wkd on 144 MHz:

2003-11-20
1611 OK1JL JO70WN 1405 km
1613 SS4M JN86CL 1710 km
1615 DL3AMB JO61EA 1471 km
1616 SP8IOV JO80HK 1458 km
1618 DL4MO JO59N 1087 km
1619 DK6NJ JN59WV 1194 km
1621 HB9DFG JN37SM 550 km
1622 G39NE JO01BJ 1302 km
1624 OK1KDO JN69KL 1431 km
1625 OK1YHO JN78GX 1651 km
1628 OM3JZZ JN88T1 1028 km
1629 DJ1FH JO40JN 898 km
1631 DL1HA JO30JN 1112 km
1634 DL8WF JO61CU 1190 km
1636 DJ4EJ JN48GB 1430 km
1657 OK1DCF JO80BJ 604 km
2232 G4ALY JO70VL 545 km
2234 E4BZ IO51UW 431 km
2256 GW4FRX IO82NA 289 km
2305 GW8ZR IO73TI 298 km
2307 GW6GJ IO73TG
tx for info, Colin

3W7SMV in I0811N wkd on 144 MHz:

003-11-20

950 DM10G J0311C 55a 55a 679 km
957 PD4HDB J0232H 55a 55a 669 km
959 PA3FXO J022GD 55a 55a 621 km
000 PDAR J022XE 55a 57a 669 km
002 PA4AR J0324H 55a 55a 598 km
006 GR1ZY J0191C 55a 59a 598 km
007 ON4MM J011UB 55a 59a 581 km
105 PE1AHV J0200S 55a 59a 695 km
108 DF8XR J0324F 55a 57a 541 km
110 PE1EJC J022PE 55a 57a 541 km
112 PE1EJC J022PE 55a 57a 541 km
113 PE1EJC J022PE 55a 57a 541 km
114 PA1EJC J022PE 55a 57a 541 km
118 PA1EJC J022PE 55a 57a 541 km
119 PA1EJC J022PE 55a 57a 541 km
120 SP9APC J022PE 55a 57a 541 km
125 PE1MJC J022PE 55a 57a 541 km
126 PE1LIS J022PE 55a 57a 541 km
128 G0CUI J022PE 55a 57a 541 km

It contacts in SSB, tx for info, Jamie

IAGNN in J98BVC wkd on 144 MHz:

003-11-20

749 DK8ZJ J054AG 59a 55a my first AU QSO
704 OK10CF J031PG 55a 55a
117 DK5DQ J031PG 55a 55a
IG: T290, 20W (1), 7ade DL6WU
x for report, Andy

A8V in KN06HT wkd on 144 MHz:

003-11-20

306 SM7WT J0650Q 59a 59a 0
107 OK10CF J031PG 55a 55a 0
108 PA0V J0331C 59a 57a 0
111 DL9MS J054WC 59a 57a 0
113 OK2EL J0650J 55a 55a 0
119 DL4MO J044UW 59a 58a 0
122 OK2LPL J044UW 59a 58a 0
123 DL6BF J032D 59a 57a 0
30 LY2BJ J025EF 59a 57a 0
30 SP4MB J033HT 55a 53a 0
34 G4MFK J0191H 59a 56a 330
38 PA5SD J022IC 59a 55a 0
44 SM6GM J057KJ 57a 58a 0
51 DK5EE J041GU 55a 55a 0 SSB
55 SM7FMC J0650J 59a 59a 340 SSB
01 OK2PZ J0650J 59a 59a 340 SSB
08 PE5DQ J033MD 59a 55a 350 SSB
12 G3UTS J094CJ 59a 57a 340
23 DKZAF J042UC 55a 57a 340
39 K8JAF J0650J 55a 57a 10
40 EW6FS J035LH 59a 59a 15
45 SP5GJH J0202R 59a 59a 15
49 SP7AME J063FG 55a 56a 15
51 OK2LPL J0650J 59a 59a 15
54 DJ3V J043VJ 57a 55a 0
54 SP2BNJ 56a 59a

1946 G3JHW I091LC 59a 55a 320

1947 DL3WV J066FL 57a 57a 340
1949 G4KXV J092AQ 55a 55a 330
1954 G4ASR J081MX 57a 55a 330
1956 G4KXV J092PB 59a 55a 330
2125 OK2SBL J092PB 59a 55a 330
2134 PA3ZDL J021HM 55a 55a 330
2156 OK2ZCL J0650Q 55a 55a 0
2206 OK2ZCL J0650Q 55a 57a 0
2206 OK2ZCL J0650Q 55a 57a 0
2210 OK2ZCL J0650Q 55a 57a 0
2253 DF2ZC J030RN 57a 55a 340
RIG: TM255E, PA 900W 4 x 11 el Yagi 35m AGL
tx for info, G4th

HB9BZA in JN36BE wkd on 144 MHz:

2003-11-20

1450 PA3OC J021FW 53a 55a 25°
1458 DF2QJ J031HL 53a 57a 20°
1523 G4W4DGU J071SV 41a 57a 15°
1530 PA0V J031HL 57a 57a 25°
1536 OK2LPL J021HM 52a 55a 30°
1541 DK8ZJ J054AG 53a 57a 30°
1547 PA1LA J032LS 56a 59a 15°
1550 G4KXV J090IR 53a 55a 15°
1602 PE5DQ J031PG 55a 55a 15°
1607 PA0PVP J022MW 55a 55a 15°
1613 PA1GYS J022MW 55a 55a 15°
1618 G4KXV J031HL 57a 59a 20°
1625 DL6BDU J043AA 54a 57a 15°
1634 DF2ZC J030RN 55a 57a 35°
1640 DL4HG J032CM 57a 57a 30°
1740 G4MFK J031HU 53a 55a 30° SSB
1744 F0HPP J019PG 53a 55a 30°
1756 G4DEZ J034AG 53a 55a 10°
1806 DK3WG J072GJ 55a 55a 340°
1810 G4KXV J091ON 53a 55a 340°
1812 PA1LA J032LS 56a 55a 10°
1826 G3JHW J091LC 55a 55a 15°
1854 DM10G J031C 56a 57a 20° SSB
1858 DL6BDU J042ON 57a 57a 20° SSB
1901 PE1AHV J0202R 57a 57a 20° SSB
1923 G4XBF J0191H 57a 57a 15° SSB
1924 G4XBF J0191H 57a 57a 15° SSB
1927 DK5DQ J031PG 56a 59a 30°
1931 PA6AT J031PG 56a 59a 35°
1933 PA6AT J031PG 56a 59a 35°
1936 DF7KZ J021HM 55a 55a 40°/87°
1940 F0HPP J019PG 55a 55a 40°/87°
1945 SP6WU J090HK 54a 57a 325°
1947 OK10CF J066FL 56a 59a 325°
1950 OK2LPL J066FL 56a 59a 325°
1952 DL10GT J066FL 56a 59a 325°
2056 DJ9FE J034H 55a 55a 15°
2120 OK1AD J090E 52a 55a 345°
2124 SP6A J091MC 51a 55a 345°
2124 DL3LT J061F 55a 55a 345°
2126 OK1MZM J098PR 51a 55a 335°
2128 OK4HG J010XO 51a 55a 335°

2300 G7RAU I090IR 52a 55a 30°

RIG: 800 W, 11-el-yagi, tx for info, Robert

HB9DFG in JN37SM wkd on 144 MHz:

2003-10-29

1721 DJ9YE J043HV 58a 55a 030° 713 km
1728 GM1XOI J085AL 53a 55a 005° 1213 km
1822 DF7KF J030PB 57a 52a 030° 378 km
1826 PA3CEC J033J 55a 55a 030° 651 km
1830 PA2DW J022GD 56a 53a 030° 557 km
1834 DK1KR J053J 56a 55a 030° 719 km
1840 DL8DU J043B 56a 54a 030° 613 km
1849 PA1T J033J 58a 59a 030° 637 km
1855 DL2EAP J031K 55a 55a 030° 440 km
1911 DF2QJ J031H 55a 52a 030° 445 km
1944 G7RAU J090IR 57a 55a 040° 734 km
2222 PA5SD J022C 58a 57a 030° 549 km
2225 PA3FQC J021HM 57a 57a 030° 539 km
2234 GM0AI J075U 57a 53a 005° 1243 km
2245 MM0CEZ J087F 53a 57a 005° 1269 km
2258 MM0CEZ J075U 52a 56a 005° 1220 km
2302 DL8AQ J043B 56a 51a 030° 724 km
2003-10-30
0015 E4HQ J051U 57a 53a 340° 1223 km
0035 G3LTF I091W 57a 57a 030° 773 km
0049 G6ATX J074J 59a 57a 010° 1212 km
0045 PE1AHV J021a 59a 57a 020° 502 km
0047 G8GJH J0016 57a 56a 030° 683 km
0055 G4AEP J091J 57a 57a 030° 745 km
0059 DL1DAX J031M 56a 57a 030° 445 km
0103 G3VDY J022C 56a 57a 030° 679 km
0110 G4LSR J081MX 55a 54a 030° 591 km
0115 PA3CQB J032M 57a 57a 030° 325 km
0120 G4BHF J083V 56a 52a 005° 1187 km
0134 PA3DTM J022a 53a 53a 030° 540 km
2120 DK1KR J055J 56a 59a 000° 719 km
2129 DL5DQ J043B 56a 57a 030° 708 km
2131 DK5DQ J031PG 59a 57a 030° 417 km
2137 DJ3AK J051J 57a 57a 030° 482 km
2139 OK10CF J075W 59a 55a 045° 579 km
2142 DL2ARD J060R 59a 55a 030° 484 km
2145 G3MZT J090W 57a 57a 030° 686 km
2147 DL6YH J032a 59a 57a 030° 501 km
2155 ON4MM J011U 57a 55a 030° 482 km
2203 ON4LH J020V 57a 55a 030° 428 km
2209 G4BRK J091H 53a 54a 030° 808 km
2212 PE1GNP J031K 54a 55a 030° 499 km
2215 PE1AQ J022M 57a 55a 030° 576 km
2218 DM10G J031C 59a 57a 030° 403 km
2222 PASVT J022H 53a 55a 030° 568 km
2240 OK2LPL J044W 55a 55a 030° 838 km
2315 DL6LPC J033B 53a 55a 030° 720 km
2328 DL6AQ J043B 56a 51a 030° 724 km
2003-10-31
0025 DL7AHR J062M 55a 53a 030° 691 km
0032 DL1RHW J062H 57a 59a 030° 642 km
0035 PA2M J021J 59a 59a 030° 500 km
0038 DF3OJ J041W 57a 55a 030° 510 km
0043 G4HRRR J073H 57a 58a 030° 1069 km

1838 DF7KF JO30GU 53A 51A 340 1811 km
My first Aurora QSO ever. RIG: IC275E, 2x16 F9FT
tx for report, Mike

LZ3GM in KN32RL wkld on 144 MHz:

2003-11-20
1811 DF7KF JO30GU 56A 53A 340 1811 km
My first Aurora QSO ever. Ind also DF22C, PA0UM, DF2JQ
RIG: 100 W. tx for report, Ivan

OE3FVU in JN78UE wkld on 144 MHz:

2003-11-20
1650 PD2DB JO22MD 59A 59A SSB 882 km
1652 DK3EE JO41GV 59A 59A SSB 661 km
1701 G1POK O91V1 56A 59A SSB 1199 km
1701 DG3BD JO43CE 59A 59A SSB 770 km
1707 PD2JVE JO22VJ 59A 59A SSB 849 km
1707 PE3QJ JO23MD 59A 59A SSB 826 km
1736 FH8PP JN18RG 59A 59A SSB 923 km
1832 GW8ASA O81EM 55A 55A SSB 1435 km
tx for info, Franz

OK2PM in JN99AO wkld on 144 MHz:

2003-10-29
1608 SM6EG JO67CQ 55A 55A 976 km
1618 OZ6OL JO85D 55A 55A 751 km
1628 SK4BX JO79OF 55A 55A 1085 km
1645 SM6YFF JO88GO 55A 55A 1061 km
1655 SM6NET JO88SD 55A 55A 983 km
1705 OZ1CL JO66GQ 55A 55A 770 km
1718 DLJLE JO44TJ 55A 55A 763 km
1721 DK2ZF JO43WJ 55A 55A 704 km
1759 PA1T JO32J 55A 55A 845 km
1830 DL4TC JO63PN 55A 55A 873 km
1834 DK3BK JO33NO 59A 59A 55A 873 km
1840 SM1CXE JO97FJ 57A 57A 866 km
1845 SK6HD JO68SD 59A 59A 943 km
1848 SM7LQ JO77DS 55A 55A 940 km
1854 DJ9YE JO43SV 55A 55A 754 km
1936 LA380 JO59CD 55A 55A 1173 km
1950 DL8BDU JO43AA 51A 55A 791 km
2024 SK6DK JO67EH 55A 55A 934 km
2105 SM5DFF JO88CN 56A 56A 1002 km
2115 SM4VE JO79SD 57A 55A 1073 km
2125 LY2BAW KO25KA 59A 59A 759 km
2159 YL2AJ KO16OY 55A 55A 688 km
2210 DL90BD JO43QN 57A 57A SSB 888 km
2224 DG6JFP JO33DN 57A 57A SSB 857 km
2230 SK6BA JO67HM 57A 57A SSB 949 km
2233 OZ4VW JO45UJ 57A 57A SSB 902 km
2246 SM7MD JO68GP 57A 57A SSB 748 km
2254 SM7DLZ JO68GJ 57A 57A 752 km
2333 ES8RO KO28WA 59A 59A 1065 km
2338 GM0TE JO87FT 55A 55A 1588 km
2344 OZ8ZS JO55RT 55A 55A 819 km
2359 RA3AQ KO85SP 55A 57A 1470 km
2003-10-30
0004 DL8LQ JO43XU 51A 55A 729 km
0019 LY2BJ JO43XU 51A 57A 803 km

1927 DLJST JO6KGU 59A 57A
1930 G3JHM IO91LC 52A 55A
1932 GWAHDB IO83GH 52A 55A
1935 DK1HO JN49JY 57A 57A
1940 DL3NTA JO6QOP 59A 59A
1942 DL4QVA JO61QH 59A 59A
1945 KH4RC JO66WD 59A 59A
1947 DL4VJ JO66WD 55A 55A
1950 DL3JW JO66FM 59A 59A
1951 DLH3DM JO31PH 59A 59A
1953 DL2TG JO61SG 59A 59A
1955 GW4FRX IO82MR 55A 55A
2120 GA6EP IO91NJ 55A 55A
2120 DF5DE JO40CX 55A 55A
tx for info, Rob

LA0BY in JO59FW wkld on 144 MHz:

2003-11-16
1431 GM4LLS IO87TP 53A 53A 330 830 km
1510 SPFGJH KO2ZRU 57A 59A 70 1036 km
1513 RAD3QX KO55JR 55A 57A 70 1285 km
1514 DLZDXA JO61VC 52A 55A 80 1004 km
1607 SP2BRO JO83JA 55A 55A SSB 867 km
1626 DLRLT JO6ZPH 53A 55A 50 886 km
2138 SPAMPB KO0GHT 52A 55A 40 918 km
2003-11-22
1830 PA1GYS JO22WV 55A 55A 280 827 km
1844 DK7HH JO53CJ 53A 55A 290 728 km
1834 DK7HH JO53CJ 53A 55A 290 728 km
2004-01-09
1920 LA1QDA JP20QH 52A 54A 340 285 km
1937 SM7UJY JO65OP 55A 57A 40 504 km
2004-01-22
1503 GM4VYX IO78TA 59A 59A 350 874 km
1655 DK8ZJ JO54AG 57A 59A 20 631 km
2004-01-25
1865 GM4LLS IO87TP 57A 57A 340 830 km
1898 OH2NY K2POCM 55A 59A 0 761 km
1813 OZ1DD JO49VT 59A 59A 20 460 km
1854 DLPOUR JO33LS 59A 59A 20 823 km
1846 DK1KO JO53CT 59A 57A 20 661 km
2004-02-11
1647 PAZDV JO22GD 55A 53A 350 940 km
1648 LAYGA JO48AE 55A 55A 350 239 km
1650 GAWKV IO92AQ 55A 55A 310 1108 km
1652 GAHGJ IO83PL 53A 55A 310 1074 km
1656 DL7JAW JO41GO 55A 57A 320 934 km
1701 SP4MWD KO33FR 57A 57A 45 918 km
1702 LY2BJ KO25ER 55A 59A 45 945 km
1712 PE1AHK JO21OS 57A 55A SSB 965 km
1716 DLNOK JO25BV 55A 59A 0 672 km
1716 DLNOK JO25BV 55A 59A 0 672 km
1721 SP3MGM JO73EE 55A 59A 40 808 km
1728 SP2DQ JO4HGM 57A 56A SSB 773 km
1735 DG8LG JO54VP 59A 59A SSB 594 km
1814 DL6PLC JO53KP 55A 53A 40 700 km
2003-11-20
LZ1YKG in KN31CS wkld on 144 MHz:

1715 G3JHM IO91LC 55A 55A 25
1728 DF7VX JO41KP 55A 58A 25
1730 OH4RG JO10XO 57A 57A 25
1830 DL8NA JO50VF 55A 56A 25
1852 PA5DD JO22JC 55A 56A 25
1855 PAZDV JO22JA 54A 54A 25
1902 DL7BYC JO41CV 53A 53A 25
1906 GW4FRX IO82MR 56A 56A 25
1912 GADEZ JO03AE 57A 57A 25
tx for info, Enrico

IK4PMB in JN54LL wkld on 144 MHz:

2003-11-20
1711 DF22C JO36RN 55A 55A 726 km
1720 PD8RFU JO33DN 55A 59A 962 km
1721 DL8YE JO31MH 59A 59A 814 km
1727 PA2M JO31LH 59A 59A 855 km
1732 GAHGJ IO83PL 55A 55A 1411 km
1739 DK5OD JO31FG 59A 59A 804 km
1743 PA3POC JO31FW 59A 59A 884 km
1746 DF7KF JO30GU 57A 55A 782 km
1752 G3JHM IO90RC 55A 54A 1069 km
1755 GADEZ JO03AE 55A 55A 1254 km
1803 PA3DV JO22JD 55A 54A 975 km
1811 SP1FPG JO73GN 59A 59A 1043 km
1819 GA4BF JO10OD 55A 55A 1030 km
1821 DL1RNV JO62GH 59A 59A 879 km
1844 GA4WV IO92AQ 59A 55A 1316 km
1856 PA8BAT JO31FX 59A 59A 809 km
1859 DF0WD JO42FD 59A 55A 872 km
1919 GA4JR JO11MX 59A 57A 1442 km
1921 GA4BF JO10OD 59A 53A 1030 km
1932 DF5DE IO91HJ 59A 59A 751 km
1935 GA4MF IO91HJ 59A 59A 1197 km
1938 SP2DFW IO93AG 59A 59A 1090 km
1940 GA4RA IO98S 59A 57A 1313 km
1943 DL8NAA JO50VF 59A 59A 642 km
1946 OK1DXQ JO70NO 58S 59R 749 km
1953 OK1CDF JO80B0 55A 55A 748 km
2015 PA3DZL JO41WJ 55A 55A 914 km
2017 DF25RD + Muzek, 8877 PA, 1tel F9FT
tx for info, Anto

IW0GPN in JN62FB wkld on 144 MHz:

2003-11-20
1746 DF22C JO36RN 35 1020 km
1814 DK5QD JO31PH 35 1101 km
tx for info, Alex

IZ4BEH in JN54WL wkld on 144 MHz:

2003-11-20
1900 GAHGJ IO83PL 59A 58A
1905 DF7VX JO41KH 55A 55A
1910 D9MGM JO52TC 59A 59A
1913 DL5VHM JO32VA 55A 55A
1915 DL50HM JO71EK 59A 59A
1919 DL1RTL JO62PH 55A 55A
1925 DF0WD JO42FD 59A 59A

1047 ONAPS JO20KQ 57A 55A 030P 402 km
1053 DO1EOX JO31PU 57A 55A 030P 482 km
1064 MB8XB IO91VR 56A 56A 030P 719 km
1100 PA38TY JO22BB 59A 59A 030P 554 km
1103 PA38TY JO22BB 59A 59A 030P 539 km
1111 DK7FC JO49HS 55A 51A 030P 263 km
1116 GM4NED IO85VJ 55A 55A 005P 1194 km
1118 ON4BBD JO21NA 56A 51A 030P 427 km
124 DH5FS JO61UA 59A 57A 030P 593 km
2003-11-20
558 SM7ALC JO65H 57A 57A 20 974 km
602 GA4BF IO91HD 57A 715 km
605 GA4G JO83PL 59A 59A 20 980 km
607 SM7MT JO65QJ 59A 59A 20 993 km
619 GA4FV IO91JN 57A 20 676 km
621 GM4NED IO85VJ 59A 59A 20 1194 km
632 GA4BYF IO85VJ 59A 59A 20 1187 km
634 DL8BDU JO43AA 59A 20 613 km
637 LY2BJ KO25ER 56A 55A 20 428 km
643 ON4LN JO20VJ 59A 59A 20 471 km
646 GA4RA IO80BS 59A 59A 20 904 km
658 SP4MUP KO33R 59A 59A 20 1155 km
709 PA5KM JO11W 59A 59A 20 587 km
712 PA1GYS JO22WV 59A 59A 20 614 km
716 OK1AHS JO60BB 59A 59A 20 453 km
729 OK1VHF JO70BS 59A 59A 20 600 km
731 DF5DE JO40CX 57A 57A 20 388 km
738 ONTCL JO20J 57A 57A SSB 421 km
740 PE3QJ JO33MD 59A 59A SSB 626 km
752 PAZDV JO22JD 56A 55A 350 557 km
801 DQJSP JO32AP 55A 55A 10 507 km
808 GW3LEW IO71P 59A 59A 350 1025 km
817 DF8IK JO30L 55A 55A 350 371 km
817 PA3FCG JO21W 59A 59A 350 539 km
828 DF22C JO30M 55A 55A 30 629 km
831 PC5C JO33PD 58A 59A SSB 502 km
839 PE1AHX JO2106 59A 59A SSB 627 km
843 DG8BEM JO43CD 59A 59A SSB 627 km
853 PA38TY JO22BB 59A 59A 30 554 km
855 PA8AT JO31TX 59A 59A 30 612 km
859 DQJSP JO32AP 57A 57A 30 777 km
902 GW3FRX IO82NR 59A 59A 30 945 km
925 DF2JQ JO31TH 59A 59A 30 555 km
941 PA3CSG JO21W 58A 58A 30 471 km
952 DL6NDK JO53AV 55A 55A 30 730 km
957 SM9CZ JO54WC 53A 53A 30 792 km
103 MMRCE IO75UA 53A 53A 10 1220 km
117 DF5XBL JO30GU 55A 55A 30 378 km
118 OZ8BK JN86BC 56A 56A 320 434 km
128 SS5M JN65M 55A 55A 320 516 km
139 GRAU IO90R 53A 53A 30 734 km
tx for info, Chris

Z2GSO in JN45NP wkld on 144 MHz:

203-11-20
331 ON4LN JO20VJ 55A 55A 25
337 PA3FOC JO21FW 55A 55A 25
742 DJ3AK JO55GJ 56A 56A 25

1652 URSSKB KN28W 55A 59A 60 1370 km
1821 IWZNRJ JN450W 59A 59A 20 756 km
1853 9AIEZA JN86HG 59A 57A 75 1006 km
1905 IZAKH JN63HV 55A 55A 85 1031 km
1909 HA1YA JN87GF 59A 59A 80 764 km
1922 11JTO JN35UB 59A 59A 80 780 km
IZAKH and HA1YA came in on vertical polarization!
tx for info, Chris

PA3DZL in JO21HM wkd on 144 MHz:

2003-10-29
1916 IAYNO JNS4KP 55A 55A 60⁸ 893 km
2127 IZABEH JNS4WL 55A 55A 65⁸ 960 km
2142 IKZJUS JNS5 55A 55A 65⁸ 817 km
2003-10-30
0234 ES3PC KO38JX 55A 53A 40⁸ 1568 km
0241 SMKAK JO69XK 57A 57A 5⁸ 1272 km
1812 LAYVH JO59EK 53A 51A 20⁸ 951 km
2001 GW3LEW JO81PS 57A 57A 30⁸ 506 km
2003-11-20
1502 SM2CEW KP1SCR 57A 59A 20⁸ 1864 km
1655 URSSKB KN28W 57A 57A 55⁸ 1452 km
1841 US5WU KO20D 59A 59A 50⁸ 1380 km
1846 HA6NQ JN86WA 59A 59A 65⁸ 1159 km
1855 SS1MQ JN75NT 57A 57A 75⁸ 997 km
1857 F8OP JN37JN 55A 55A 75⁸ 872 km
1901 F8OP JN26AM 59A 59A 75⁸ 557 km
1903 HA5PT JN7MK 59A 57A 70⁸ 1133 km
1908 OK1DXW JO7HC 57A 59A 70⁸ 719 km
1913 OK2CP JN89U 55A 55A 70⁸ 990 km
1918 FH1JL JN12KQ 59A 59A 75⁸ 991 km
1920 HA7WJ JN7MP 59A 59A 75⁸ 1122 km
1923 HB8BQ JN37D 59A 59A 70⁸ 1107 km
1925 HA3HV JN66UK 55A 55A 70⁸ 603 km
1933 HB9EZA JN36BE 59A 55A 65⁸ 639 km
1934 FBEG JN25J 59A 55A 70⁸ 983 km
1938 OK2WB JN9F5 59A 59A 75⁸ 1141 km
1942 HA5OV JN7N1 59A 59A 75⁸ 1272 km
2133 HA8V JN69HT 59A 59A 80⁸ 746 km
2146 GW4FRX JO2MR 59A 59A 20⁸ 334 km
2151 MM0CEZ JO7SX 59A 59A 20⁸ 362 km
2234 G4EZ JO40AE 57A 57A 40⁸ 794 km
2252 I2FK JN35UB 55A 55A 85⁸ 939 km
2301 SS5M JN65TM 55A 55A 85⁸ 753 km
2308 11JTO JN35UB 55A 55A 85⁸ 735 km
2009 K1EGC JN35U 55A 55A 85⁸ 753 km
2312 11DMP JN35U 55A 55A 85⁸ 753 km
2314 K3PNB JN54L 54A 55A 85⁸ 954 km
tx for report, Jac

PA5DD in JO22JC wkd on 144 MHz:

2003-11-20
2326 SS5M JN65TM 53A 55A 960 km
2315 I2FK JN450B 56A 57A 849 km
1936 HA1YA JN87GF 59A 59A 1008 km
1931 SS4M JN86CL 55A 54A 1041 km
1926 DL3JN JO60HR 54A 55A 571 km
1924 YU7OP KN05AO 54A 55A 1331 km

2212 HA8NY JN86WC 56A 55A
2215 OK1FRD JO80DD 53A 55A
2216 DL3NCR JN48XQ 59A 57A
2218 JN63HV JN58SE 55A 56A
2221 HA5JUT JN97OM 54A 55A
2225 DK8JL JO54AG 59A 59A
2226 ES3PC JO78DF 55A 57A
2226 ES3PC JO78DF 55A 57A
2240 QZ1HXL JO93AG 52A 55A
2245 G4YTK JO93AG 51A 54A
2249 GW3LEW JO71PS 56A 57A
2343 EL4DQ IO85HV 53A 53A
2345 GM4CLN IO85HV 53A 53A

PA3DZL in JO21HM wkd on 144 MHz:

2003-11-20
1713 OM3KZA JN69F 56A 55A 0-90⁸
1715 FP9AI JN69MT 59A 59A
1716 SPOE JN780G 55A 59A
1719 S57TW JN75EX 59A 59A
1720 LZ2FO JN13XQ 59A 59A
1721 94JH JN13XQ 59A 59A
1723 IO3XW JN65Q 55A 55A
1726 SP6RBG JO71SF 54A 54A
1728 OKYVK JO70DB 56A 55A
1729 OK1YN JO70DB 55A 59A
1733 HA5PT JN67MK 56A 55A
1735 HA0HO JN67MK 55A 55A
1740 HB8DM JN47SL 59A 57A
1741 IK2GSO JN37N 57A 57A
1743 F8OP JN26AM 59A 59A
1749 IK1SPR JN34TQ 56A 59A
1750 SS1MQ JN75NT 53A 55A
1753 11DMP JN35UB 54A 55A
1807 SS4G JN65TM 53A 55A
1836 USSWU KO20D 53A 59A
1840 OK2MWR JN69H 55A 55A
1845 SP8A JO81ME 53A 55A
1847 HA5CV JN97PM 53A 56A
1848 OK2KOS JN69CT 55A 55A
1903 PB4FUN JO33KC 56A 53A SSB
1907 F04HD JN28D 59A 59A SSB
1909 F0DZG JN28D 59A 59A SSB
1910 OK1FD JO62H 59A 59A SSB
1914 SP9QJ JN99H 57A 59A SSB
1915 SP9MCY JO90G 56A 59A SSB
1918 HA6NO JN86WA 56A 57A SSB
1919 SP9QJ JN99H 57A 59A SSB
1920 F0TA JO90FB 57A 57A SSB
1931 OM5JM JN99E 53A 55A
1943 F6DRO JN03TJ 55A 57A
2017 HB9EZA JN68BE 55A 57A 270⁸
2146 GM4VLV IO86J 53A 55A 340⁸
2149 GM3WJ IO86F 41A 51A 330⁸
2203 SM7SPG JO66MD 57A 57A 30⁸
2224 SK6DK JO67HE 57A 57A 0⁸
Rig. 300W, 9-ele DK7ZB. tx for report, Gellian

PA2CHR in JO22XA wkd on 144 MHz:

1620 SP2BDR JO83UA 59A 59A 60 804 km

043 G7RAU IO90R 57A 57A 1378 km
053 EH4Q IO51WU 52A 55A 1848 km
101 GA4SR IO81MX 52A 55A 1493 km
120 GM4VLV IO86J 52A 55A 1600 km
128 GM4BYF IO85V 52A 55A 1583 km
134 GM4VXV IO78TA 54A 55A 1731 km
222 ES3PC KO38JX 53A 55A 1124 km
236 LY3OD KO24QH 55A 55A 717 km
ig. 80W, PA0MS antenna. tx for report

JN4K7G in JO10XO wkd on 144 MHz:

2003-07-26
333 GM4VXV IO78TA 51A 54A 340⁸
331 GA4JC IO91VJ 55A 55A 340⁸
331 G3UJF IO83UJ 52A 55A 340⁸
338 DL4KUG JO64PB 52A 55A 340⁸
341 DL2BA JO64X 53A 55A 340⁸
347 LY3OW KO25KA 55A 59A 25⁸
311 GM0VHR IO89L 54A 57A 340⁸
311 GM0VHR IO89L 54A 57A 340⁸
322 GM4VXV IO78TA 55A 55A
327 SK6DK JO67HE 53A 55A
2003-08-21
115 SM7EOE JO86FP 57A 59A 350-60⁸
726 EW6FS KO35LB 55A 56A
732 GM1XO IO85AT 57A 59A
2003-10-29
051 ES8RQ KO28WA 55A 59A
02 G4SJO IO64SK 55A 53A
30 RA3LE KO64AR 55A 57A
54 SP4JWO KO39FR 52A 56A
01 DJ1TO JO62SL 52A 56A
05 G4GMS IO82XT 51A 52A
08 O2BTV JO55XE 54A 57A
16 GM4LLS IO87IP 56A 57A
15 LY2BJ KO25ER 57A 59A
03-10-30
04 QZ3N JO46S 55A 56A
15 OK1CZ JO70EC 51A 41A
19 OK1FD JO69C 52A 55A
20 E8GF IO62R 51A 52A
28 DL3DXX JO61VC 51A 55A
23 OK1AD JO80E 51A 55A
40 SP9WU JN99H 54A 56A
12 OK2PWY JN89KW 54A 59A
14 OK2DZ JO51MV 56A 59A
16 DM2TX JO51MV 56A 59A
18 DL1GT JN58AO 52A 57A
11 OM5CM JN98DF 52A 57A
15 K1TTR JN34TC 53A 55A
8 OK1PM JN34TC 56A 59A
9 11DMP JN35UB 55A 55A
9 948WV JN86HE 56A 55A
8 DH6FS JO61UA 55A 55A
5 SK6DK JO89X 51A 54A
4 K1EGC JN35U 55A 55A 45⁸
8 DL3JN JO60P 52A 59A
0 DL1VDL JO61WB 57A 59A

1922 DJ0VZ JN49NA 55A 59A 463 km
1921 OK2BOS JN79WF 55A 57A 849 km
1920 HA5OV JN87N1 57A 59A 1159 km
1918 9AICMS JN86FM 56A 56A 1052 km
1916 IZCPO JN457T 57A 55A 802 km
1915 HA8BUT JN97MI 58A 57A 1156 km
1913 SN7OL JO91 55A 56A 982 km
1911 HA3HV JN86UK 56A 59A 1132 km
1910 OK1YVK JO70DA 57A 59A 770 km
1909 F8OP JN26AM 57A 59A 620 km
1907 OK2CP JN89U 54A 55A 1328 km
1905 EL3AI KO22CE 57A 55A 619 km
1903 DL2DWR JO61TK 57A 59A 619 km
1901 DL2TG JO61SG 57A 57A 617 km
1900 HA5PT JN87NK 55A 55A 1151 km
1859 SS1MQ JN75NT 57A 59A 1032 km
1857 YU1LA KO4FR 57A 59A 1418 km
1855 HA7WJ JN97MP 53A 57A 1139 km
1854 11JTO JN35UB 57A 55A 812 km
1852 IK2GSO JN45NP 56A 55A 976 km
1850 OK2PM JN99AO 53A 55A 796 km
1849 YU7KB JN84XX 54A 55A 1370 km
1848 F8CJU JN807K 57A 58A 591 km
1847 F8RAB JN80RJ 54A 57A 471 km
1845 DL2JWN JO60W 53A 55A 569 km
1844 OK2KOS JN99CT 56A 57A 979 km
1843 UT5DL KN18EP 55A 59A 1308 km
1841 OK1YN JO70DB 55A 59A 707 km
1840 LY2CI KO14WV 55A 57A 1302 km
1837 SP3QDM JO82EF 54A 55A 796 km
1835 DL5ZGP JO51WV 56A 57A 435 km
1833 SP3JBL JO91BR 55A 59A 921 km
1832 DL5AXJ JO50SS 53A 53A 497 km
1831 DL8WF JO61CU 55A 55A 515 km
1741 944EW JN85G 55A 55A 1285 km
1739 S57TW JN73EX 55A 59A 979 km
1738 SS1NK JN76HB 52A 55A 985 km
1654 URSSKB KN28W 57A 55A 1453 km
1642 SP2BNJ JO94HK 54A 59A 961 km
1620 OK1VHF JO70EB 59A 59A 713 km
1615 OK1SC JO70DB 56A 59A 788 km
1610 SM7BYB JO76X 57A 59A 1493 km
1552 11DMP JN35UB 57A 57A 812 km
1551 SP2RS JO94GK 55A 55A 896 km
1550 OK2PTC JN89JM 56A 55A 896 km
1546 FV3VH JN04FT 57A 57A 867 km
1545 DM3SWD JO62WL 54A 57A 558 km
1543 OK1KDO JN69KL 53A 59A 644 km
1539 HA8V JN86HT 55A 59A 1289 km
1537 942AE JN86HF 56A 55A 1083 km
tx for info, Uffe

P00EBF in JO21LQ wkd on 144 MHz:

2003-11-20
1505 GM4VXV IO78TA 59A 59A
1510 GAUXC IO82BC 59A 59A
1522 GM4TCE IO87HF 59A 58A

G3XDY in JO202B wkd on 432 MHz:

2003-11-20
1539 PA3ECU JO202C 57A 59A 35° 134 km
1541 DF90X JO24HD 59A 59A 40° 283 km
1543 DL9EX JO49RE 59A 59A 40° 377 km
1548 DB3BV JO24C 55A 55A 35° 583 km
1614 DL3UN JO60LX 57A 58A 40° 180 km
1630 DJ9DL JO31NC 57A 57A 40° 624 km
1637 DL7FF JO62TJ 57A 57A 40° 928 km
1641 SP2CNW JO34AD 55A 59A 45° 608 km
1645 DK9QR JO88BV 59A 59A 45° 608 km
1647 DK3WG JO72GJ 57A 57A 45° 685 km
1647 SS1ZO JN68DR 59A 59A 65° 1000 km
1711 9A2SB JN55GM 55A 59A 65° 1219 km
1713 HB9EZA JN38BE 59A 59A 65° 603 km
1715 HB9DKM JN37LL 55A 59A 65° 474 km
1717 OE1SOW JN88FF 51A 55A 65° 922 km
1721 OM1TL JN88ND 57A 59A 65° 970 km
1723 PA4FP JO33BC 59A 59A 65° 203 km
1725 OE2LDM JN67NW 55A 55A 55° 734 km
1735 DF10Q JO32TD 55A 55A 55° 217 km
1808 EISFK JO51RT 54A 54A 315° 907 km
1823 DJ4TC JO33PN 55A 57A 330° 628 km
tx for report, Joe

PA5DD in JO222C wkd on 432 MHz:

2003-11-20
1822 KG9TF JO31NF 57A 56A 193 km
1807 EISFK JO51RT 52A 52A 200 km
1729 DF10Q JO32TD 56A 55A 867 km
1728 FS9HX JN41FT 55A 55A 116 km
1725 DL1EA JO31DS 55A 55A 570 km
1722 DL1RG JO62MW 56A 57A 435 km
1721 DM2DXG JO51MW 56A 56A 415 km
1720 DK1VI JN49HH 55A 58A 601 km
1718 DK2WV JN58 53A 59A 781 km
1715 FC8RP JN6KE 57A 57A 1453 km
1713 UR5SKB JN28HW 55A 57A 743 km
1712 OK1AG JO70NN 54A 55A 744 km
1710 HB9DKM JN37SL 57A 59A 715 km
1702 OK1TEH JO70ED 57A 59A 953 km
1701 OE3JPC JN4TEW 59A 59A 1009 km
1700 HLCK JN54PD 55A 55A 1247 km
1658 9A2SB JN55GM 54A 59A 665 km
1656 HB9EZA JN38BE 57A 57A 570 km
1655 DL5DSM JO71EK 52A 55A 670 km
1651 SS1ZO JN68DR 57A 59A 1028 km
1644 DJ2VZ JO31PF 59A 59A 204 km
1635 DL1SUN JO33PN 59A 59A 473 km
1633 DL3JN JO60LX 57A 58A 585 km
1632 D4MB JN30NR 57A 57A 227 km
1630 DJ9DL JO31NC 59A 59A 201 km
1629 DL7FF JO62TJ 58A 59A 609 km
1607 DL2WY JO43NL 57A 55A 336 km
1558 SP2CNW JO34AD 55A 59A 908 km
1534 OZ1CZ JO46DE 57A 57A 542 km
tx for info, Uffe

PD0EBF in JO21LQ wkd on 432 MHz:

2003-11-20
1633 DL5YET JO41EV 55A 52A tx for info, Hank

G3XDY in JO202B wkd on 432 MHz:

2003-11-20
1656 FC6RP JN66 675 km
1900 9A2SB JN65 1458 km
1903 DB3BV JO24C 55A 55A 35° 583 km
1907 OK2MWR JN88 1239 km
1909 SS1ZO JN65 682 km
1914 DL6NAA JO37 1174 km
1914 HB9DKM JN37 808 km
1917 FV5VX JN04 653 km
2052 DK6AS JO52
tx for report, John

GW4DGU in IO71SV wkd on 432 MHz:

2003-11-20
1907 9A2SB JN55GM 55A 55A 1821 km
1917 SS1ZO JN68DR 55A 55A 1603 km
tx for info, Chris

HB9EZA in JN38BE wkd on 432 MHz:

2003-11-20
1648 DL6BF JO32QI 52A 56A 35° 165°
1650 G3LOR JO20ZF 56A 59A 35°
1653 PA0EZ JO22FC 54A 56A 30°
1656 PA0EZ JO22FC 57A 57A 30°
1659 PASDOL JO22MT 52A 57A 30°
1709 DJ2VZ JO32MF 52A 57A 30°
1713 PA3DZL JO21HF 52A 57A 32° 175°
1718 PA0BAT JO31FX 54A 55A 35°
1724 DJ6MB JO30NR 55A 55A 25°
1906 G3LOR JO20ZF 55A 55A 30°
1910 DF5UJ JO43XO 53A 41A 30°
RIG: 800 W, 11-dec-yagi. tx for info, Robert

PA2CHR in JO22XA wkd on 432 MHz:

2003-11-20
1659 SS1ZO JN68DR 59A 59A 75 952 km
1703 HLCK JN54PD 55A 57A 75 961 km
1917 OK2MWR JN59H0 55A 55A 65 927 km
Used 15 deg elevation for OK2MWR. tx for info, Chris

PA3DZL in JO21HM wkd on 432 MHz:

2003-10-30
2340 G4BRX IO91BP 55A 55A 45° 449 km
2346 G3DYX JO20EB 55A 55A 45° 242 km
2003-10-31
0000 DK6VS JN39AE 59A 59A 65° 310 km
0033 DF8IK JO3JUT 59A 57A 65° 170 km
0034 HB9EZA JN38BE 54A 57A 65° 603 km
0121 OK2MWR JN67EW 57A 57A 65° 932 km
0124 DK3QR JN68BV 55A 57A 60° 608 km
2003-11-20
1512 G3LOR JO20ZF 57A 58A 20° 237 km
1519 DL1SUN JO53RO 55A 59A 35° 516 km
1522 DJ6MB JO30 55A 55A 35° 402 km
1524 DF6NA JN49XS 55A 55A 45° 423 km
1533 G4RKG IO91ON 57A 59A 35° 374 km

1600 OK1TEH JO70FD 59A 59A
1602 DF90X JO24HD 55A 55A
1605 DO1CTL JO60LT 59A 57A
1607 UA3PTW JO90BS 57A 52A
1612 DG6CZ JO60DS 55A 55A
1615 LY2CI KO14WV 55A 59A
1616 SP2MKO JO24C 55A 55A
1623 DB3BV JO42AC 55A 55A
1627 SNRBI JO57XQ 57A 55A
1630 SNRSG JO60ZF 53A 53A
1633 G3LOR JO60ZF 55A 55A
1635 SS1ZO JN68DR 59A 59A
1637 SS1ZO JN68DR 59A 59A
1641 OM1TL JN88ND 59A 59A
1642 LK4L JN54DD 59A 55A
1644 OE3QP JN88BV 59A 59A
1645 DK9QR JN55GM 59A 59A
1646 9A2SB JN55GM 59A 59A
1649 OK1AG JO70NN 55A 55A
1713 PA4FP JO33BC 59A 59A
1738 DL1EA JO31DS 53A 55A
1739 EISFK JO51RT 53A 51A
tx for report, Norbert

DL80BE in JO42XI wkd on 432 MHz:

2003-11-20
1623 UA3PTW KO93 54A 54A
1628 DF6NA JN49 55A 56A
1632 SP2MKO JO93 59A 57A
1633 DK6VS JN49 53A 53A
1636 SP2CNW JO93 59A 55A
1639 LY2CI KO14 59A 52A
1644 SS1ZO JN68 59A 56A
1700 OE2LDM JN67 59A 59A
1703 OE3JPC JN67 59A 59A
1710 OK1TEH JO70 59A 56A
1714 LK4L JN54 59A 59A
1720 9A2SB JN54 59A 56A
1726 OM1TL JN88 59A 59A
1804 EISFK IO51 55A 56A
tx for report, Jürgen

F6CRP in IN96KE wkd on 432 MHz:

2003-11-20
1715 PA5DD JO22IC 57A 57A
1717 DF2JQ JO31HL 51A 57A
1720 DF6NA JN49XS 55A 54A
1729 DK6VS JN39NF 55A 55A
1730 EISFK IO51RT 55A 55A
1741 G4RKG IO91ON 55A 55A
1812 DJ4TC JO36PN 55A 55A
1857 G4BRK IO91DP 59A 59A
1858 G3DYX JO20EB 59A 59A
1900 G3RBJ IO91FE 51A 55A
1904 GW4DGU JO71SV 57A 55A
1913 SP8WQ JO30RK 55A 55A
1918 DL1EA JO31DS 55A 55A
RIG: FT847 + 100 w amp + 2x 21 el - MGF 1302
tx for info

1713 DK3UHV JO61JO 55A 55A 345 924 km
1716 DG3SLM JO61EG 55A 55A 345 871 km
1726 PE2DB JO22MD 55A 51A SS8 1344 km
1734 DK5DO JO21PG 55A 53A 340 1162 km
1750 OH4NL JO20IV 55A 55A 330 1310 km
1755 DF2ZK JO30RU 55A 55A 330 1117 km
1757 DF7FK JO30RU 55A 55A 330 1169 km
1759 DF2JQ JO31HL 55A 55A 330 1214 km
1818 OK1CDF JO88BJ 59A 59A 330 625 km
1904 G4ZBI JO03AI 57A 55A 335 1706 km
1908 SP3JH JO91BR 55A 57A 335 700 km
1923 GRAU JO90IR 57A 55A 330 1771 km
1924 G4XBF IO91DO 55A 52A 330 1615 km
2127 DL2KRO JO53CT 55A 55A 0 1174 km
2240 DK1AR JO60AR 55A 55A 5 850 km
2252 OK1VT JN79IX 55A 55A 5 653 km
tx for info, Pista

432 MHz reports

3A2SB in JN95 wkd on 432 MHz:

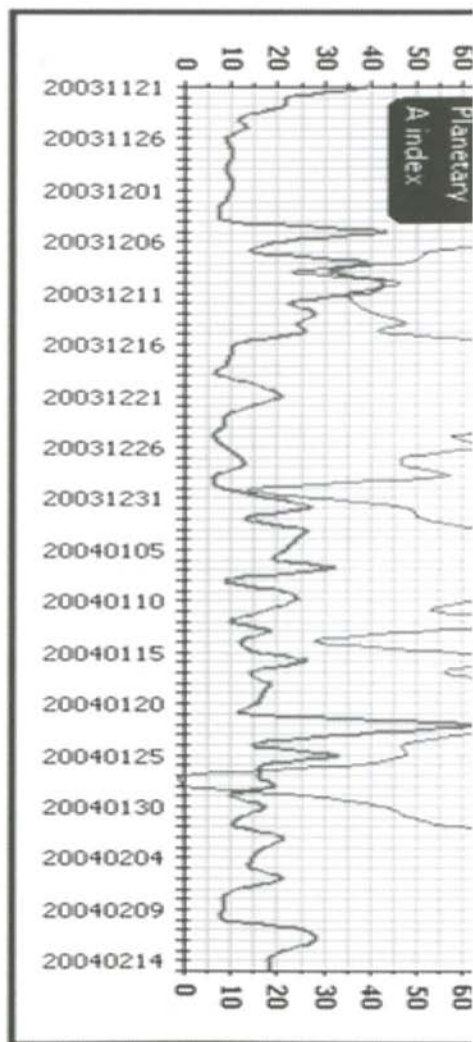
2003-11-20
1649 DL1SUN JO53PN 59A 59A 20 1006 km
1657 PA5DD JO22IC 59A 54A 350 1247 km
1706 DK3WG JO72GJ 59A 57A 40 814 km
1710 PA3DZL JO21HM 59A 55A 350 1221 km
1715 PA0EZ JO22OF 59A 59A 350 1225 km
1718 DL80BU JO42XI 59A 53A 30 984 km
1725 G4RKG IO91ON 59A 57A 350 1568 km
1730 G3LOR JO20ZF 55A 55A 350 1456 km
1900 G3XDY JO20EB 55A 55A 350 1459 km
1910 G4BRK IO71SV 55A 55A 340 1821 km
1910 G4BRK IO91DP 55A 55A 345 1632 km
3g. Ham transmitter 50 W, CF-300 Preamp. 23 el. DL8BV yagi
tx for info, Zinko

DL1EJA in JO31DS wkd on 432 MHz:

2003-11-20
726 PA5DD JO22IC 55A 55A 115 km
726 G4RKG IO91ON 52A 55A 489 km
752 OZ1CZ JO46DE 55A 53A 527 km
805 EISFK IO51RT 53A 51A 1018 km
838 DL3YEE JO42GX 55A 55A 203 km
856 DK6VS JN39NF 55A 55A 289 km
900 DL4WV JO63SO 59A 59A 532 km
918 FC6RP JN6KE 55A 55A 823 km
tx for report, Oliver

DL1SUN in JO53PN wkd on 432 MHz:

2003-11-20
516 PA3DZL JO21HM 58A 55A
538 OZMM JO66FR 55A 55A
540 DJ6MB JN30NR 59A 59A
545 SP2CNW JO34AD 59A 59A
550 DF6NA JN49XS 55A 55A
553 DL4OL JO52BC 55A 59A
555 DJ2VZ JO31PF 57A 57A
556 DL6BF JO32QI 57A 57A



Planetary A Index November 2003 to February 2004. Only one strong peak on Jan. 22/23 and mediocre ones on Dec. 5th and 10/11.

Source: <http://dxlc.com/solar>

Reports of SK4MPI/B, JP70nj, via Aurora on 2m:

Date Time	from area
25.11. 1445-1600	OH5 (KP30), N-DL (JO53)
30.11. 1510-1730	SM3, SM2 (KP15)
5.12. 1450-1915	OH6 (KP02), N-DL (JO42)
6.12. 1440-1605	OH6, OH2 (KP30)
7.12. 1740-1845	SM3 (JP82), SP4 (KO03), PA (JO33)
8.12. 1130 + 1605	OH5 (KP30)
1640-1935	N-DL (JO53)

2200	OZ (JO65)
9.12. 1540	OH6 (KP02)
10.12. 0945	OH5 (KP30)
1525-1735	OH1 (KP00), N-DL (JO32)
11.12. 1400-1645	OH6 (KP02), GM (IO87), PA (JO33)
12.12. 1420-1730	OH5 (KP30), N-DL (JO53)
13.12. 1430-1540	OH6 (KP30)
14.12. 1820-1830	SM3, N-DL (JO53)
15.12. 1355	OH5 (KP30)
16.12. 1729	OH5 (KP30)
20.12. 1505-2010	SM2 (KP15), N-DL (JO53), SP2 (JO94), OH1
22.12. 1605-1700	OH5 (KP30), OH6 (KP02)
23.12. 1525	OH6 (KP02)
27.12. 2145	SM3 (JP82)
31.12. 1555-1905	OH6 (KP02), PA (JO33), DL (JO53)
1.1. 0145	SM0 (JO89)
5.1. 1640-1815	N-DL (JO53), UA1 (KP50)
7.1. 1235-1420	LA (JP33), N-DL (JO53), PA (JO33)
1910	OH6 (KP02)
9.1. 1550-2100	OH6 (KP02), N-DL (JO53), UA1 (KP50), LA (JP33), SP (JO94)
10.1. 1755	OH6 (KP02)
11.1. 1745	N-DL (JO53)
12.1. 1605	OH5 (KP30)
13.1. 1710	OH5 (KP30)
14.1. 1825	OH5 (KP30)
15.1. 1440	OH6 (KP02)
16.1. 1130 / 2130	OH6 (KP02) / OH8 (KP24)
17.1. 1815-1829	OH6 (KP02), OH8 (KP24)
18.1. 2110	UA1 (KO59)
19.1. 1404-1645	OH5 (KP30), NDL (JO53), PA (JO32)
20.1. 1310 / 1500	OH5 (KP30) / SP4 (KO03)
22.1. 1215-1845	OH5 (KP30), DL (JO53), SP2 (JO94)
2240	N-DL (JO53)
23.1. 0950	OH5 (JO32)
1610-2000	N-DL (JO33)
24.1. 0115	N-DL (JO33)
25.1. 1430-2000	OH6 (KP02), N-DL (JO53)
26.1. 1440	OH5 (KP30)
27.1.* 1350	OH5 (KP30)
28.1. 1740	OH5 (KP30)
30.1. 1230 1510*	OH5 (KP30)
2.2.* 1225	OH5 (KP30)
3.2. 1240	OH5 (KP30)
1755*	SM2 (KP04)
10.2.* 1725	OH5 (KP30)
11.2. 1300-1900	OH5 (KP30), N-DL (JP53, 33)
12.2. 1230-1630	OH5 (KP30), PA (JO33), UA1, OH6
13.2. 1540	OH6 (KP02)
14.2. 1320	OH6 (KP02)
1950-2200	N-DL (JP53), SM2 (KP15)
15.2. 1030 / 1550	SM0 (JO89) / SM2 (KP25)
17.2. 1945	SM2 (KP15)
18.2. 1655	OH5 (KP30)
23.2. 1800-1915	OH5 (KP30), SM2 (KP15)
27.2. 1940-2015	OH1 (KP10), N-DL (JO53)
1.3. 1510-2000	OH1 (KP10), N-DL (JO53), SM2

* report of LA2VHG JP53eg

New Products

Editor: Joachim Kraft, DL8HCZ

New PA from DanAmps Dual 50 & 144 MHz PA

DanAmps (a Danish company owned by OZ2ELA and OZ1J) plans to sell a new Dual 50 & 144 MHz Linear Amplifier starting from autumn 2004.

This amplifier is fully microprocessor controlled with automatically band switch.

Conservative rated 1500 Watt RF output with no time limit in all modes will make also the WSJT users happy.

Here some more features and specifications:

T/R-QSK shift with RF output vacuum relays for high reliability.

3,5 kW+ power supply with unique uncore transformer.

Frequency Coverage: 50-52 MHz and 144-148 MHz one band at any time.

Power Output: Minimum 1500 Watt at any mode no time limit with approx. 60 Watt drives.

Input VSWR: < 1,5:1 within amateur bands.

Harmonics: More than 50 dB below rated output.

IMD distortions: Better than 35 dB below rated output.

Protection: Cover interlocks, inrush power-on current, all supply voltages and currents, excessive drive, load mismatch, RF/DC flash over, severe mistuning, SWR faults, tube temperature, primary and secondary fuses, cooling fans.

Tube cooling: Variable DC controlled forced airflow depend on tube temperature for minimum noise.

PSU cooling: Variable DC controlled forced airflow depend on transformer temperature for minimum noise.

ALC: Grid-current derived, 0 to -11 Volts via panel adjustable.

RS 232 port: All measurements data accessible on personal computer via RS 232 standard port.

Primary Power: 190 - 264 Volt (210, 220, 230 & 240 V nominal taps, 10% tolerance), 50-60 Hz, single phase, 20 A maximum 360 - 440 Volt (400 V nominal tap, 10% tolerance), 50-60 Hz, dual phase, 12 A maximum (internal manual switching).

Tubes: Two GU74B/4CX800A Svetlana High Performance Ceramic Metal Tetrode.

Dimensions: 44,5 (W) X 20 (H) X 43,0 cm (D)

Weight: approx. 35 Kg.

Price: Around 4100 Euro.

For more info you can email info@danamps.com

And there is a website www.danamps.com

Neu von WiMo USB CAT-Interface

Endlich ein CAT-Interface für USB! Viele moderne Notebooks haben nur noch USB- und keine COM-Schnittstelle. Bei Desktop PCs ist die einzige COM-Schnittstelle häufig schon belegt. Es können beliebig viele USB-CAT Interfaces an einen PC angeschlossen werden. Die USB-Interfaces unterstützen Win 98/ME / 2000/XP, Linux 2.4.0 und höher, Apple Macintosh OS-8/OS-9/OS-X. MS-DOS wird NICHT unterstützt!

Nach der Treiberinstallation erscheint jedes USB-Interface als COM-port in der Systemkonfiguration. Deshalb arbeitet das Interface mit jeder Software, die auf normale serielle COM-Schnittstellen zugreifen kann.

In den USB-Interface sind die erforderlichen Pegelwandler für die Funkgeräte bereits enthalten, so daß die Original-Pegelwandler der Hersteller (z.B. CT-62, IF-232, CT-17) nicht mehr benötigt werden. Zur Steuerung von CW und PTT sind zwei weitere Interfaces eingebaut so dass auch dafür kein extra Interface mehr benötigt wird.

Im Gegensatz zu den einfachen USB/seriell-Adaptern hat das MicroHAM USB-Interface ein einstellbares Metallgehäuse. Alle Eingänge sind über 3-polige LC-Filter entstört, das Interface selbst produziert keine Störstrahlung.

Stromversorgung erfolgt unter Einhaltung der Spezifikation über den USB-Anschluß, zusätzliche Stromversorgung ist nicht erforderlich.



Fig.: MicroHAM USB-Interface

6 verschiedene Modelle für YAESU / ICOM / KENWOOD / Ten-Tec / Elecraft K2 sind erhältlich. Der empfohlene Verkaufspreis beträgt 42.-EUR, Lieferung erfolgt komplett mit Treiber-CD und Anleitung.

Bezug: WiMo Antennen und Elektronik GmbH, Am Gäxwald 14, 76863 Herxheim. Tel. 07276 96680 FAX 07276 6978

<http://www.wimo.com>
email: info@wimo.com



Fig. left: Danamps PA

News & Comments

Editor: Joachim Kraft, DL8HCZ
info@DUBUS.de

Beacons

Beacons for Mauretania (5T)

5T5SN will erect beacons for 6m on 50.028 MHz and for 2m (close to 144.300 MHz) very soon. The Call will be 5T5SN/B. May be also 5T5SIX and 5T5VHF. Locator is IK28AC. This project was initiated by DUBUS. DF7KF will supply 2m-TX, PSU and keyer, W3CMP will supply the 6m-TX. Tnx also to DL8YHR for coordination and help! (DL8HCZ)

New 2m Beacon for EA8

In order to bring more light into the speculations about the Es - Doublehop-Es - Tropo-Es propagation discussion very soon a 2m beacon will be qrv from La Palma Island (IL18BS). Pwr around 50W and 2 yagis, one to EU another to USA! This is another project initiated by DUBUS. DF7KF will supply TX, PSU and keyer. DL7AJA will supply antennas and will bring the beacon to EA8. DL6FAW coordinates everything on La Palma. Many thanks to all for supporting a really most important beacon project! (DL8HCZ)

LA9SHF new on 10 GHz

LA9SHF is qrv on 10368.890 from JO59FS with 1.3 W power and 10dB horn. QTF is 180. So good for OZ/DL. (LA8GKA)

DB0FAI QRP

The 2m high power beacon DB0FAI, Locator JN58IC, is running around 30WERP only since a couple of months now. Reason is a PA failure and 2 spare transistors BLY94 are needed. It is quite difficult to find them. Anyone who could help, please contact: DL5MCG@amsat.org

GB3MCB Info

The lease of the tower has been extended now to April 2005. So the 2m and 6m beacon continue to be qrv from IO700J. The 70cm beacon is QRT because the antenna was lost. The 23cm beacon has been stolen.... (G3YJX)

New 6m beacon on VP5

W3CMP is organizing a new 6m beacon on Turks and Caicos Islands (VP5). Locator FL41.

DB0XL QRT

The 10 GHz beacon DB0XL (JO53HU) is QRT now. It is planned now to erect this beacon on 10368.805 and 24048.805 GHz in the Schwerin area, JO53QP, at the DB0SI QTH (90m asl). Callsign will also change to DB0SI.

New 24 GHz Beacon QRGs in DL

Here the proposed new QRGs for the 24 GHz beacons (144 MHz down in frequency) in DL. Several beacons are already qrv on the new QRG.

DB0SI	24.048.805 JO53QP (ex DB0XL)
DB0VI	24.048.815 JN39NK
DB0JL	24.048.820 JO31MC
DB0FGB*	24.048.833 JO50WB
DB0KI	24.048.840 JO50WC
DB0GW	24.048.850 JO31JK
DL0WY	24.048.853 JN67AQ
DB0ARB	24.048.860 JN69NC
DB0JK	24.048.865 JO30LX
DB0HW	24.048.875 JO51GT
DB0TUD	24.048.885 JO61UA
DB0ECA	24.048.895 JN57UU
DB0CU	24.048.900 JN48BI
DB0NCO	24.048.905 JN59MI
DB0HEX	24.048.910 JO51HT
DB0RG	24.048.912 JO51GO
DB0AJ	24.048.920 JN57 (QRT)
DB0KHT	24.048.930 JO40FE
DB0JO	24.048.940 JO31SL
DB0KNL	24.048.950 JO40RW
DB0MFI	24.048.960 JN58KR
DF0ANN	24.048.965 JN59PL
DB0AS	24.048.985 JN67CR

* confirmed to be qrv on new QRG

KL7GLK beacon on 144.289 MHz

Larry, KL7GLK, writes:

The 2 metre beacon KL7GLK is currently running low power, 4-5 watts into a turnstile antenna overlooking the Chesapeake Bay (Grid FM18, no KL7GLK is not in Alaska!), with a clear view to the horizon, over water, of 20 km (at its antenna height, on a cliff).

On the tower, and facing towards Northern Europe, is a pair of 13 element vertical polarized beams with a theoretical gain of 15 db. Its not been actually measured. In its "European Mode" the beacon can put 100 watts, and this has been directly measured at the antenna input, into this array, giving something on the order of 2-3kw ERP. The power could be higher, but our government limits the input to 100 watts in this type of operation without special authorization. The beam width is fairly narrow, but by the time it would/could reach Europe the spread should be considerable. Of course, it has never been heard there so all this is to be taken with 'a grain of salt' - lots of theory but no results.

It is not operated in this mode (high power) often since it can annoy the devil out of local 2 meter hams in its beam path. It can be accessed and switched from low power to hi-power directly by a uhf control link, but to this date that occurs only when there is a possibility of it being heard in Europe. If 50 Mhz is open to Europe it will be turned on providing someone with access to the beacon control link, and aware of the 50 Mhz opening, will activate it. Once activated, it will transmit for 1 hour then automatically resume low power operation unless reset.

Unfortunately those few of us associated with the beacon have had waning interest in VHF over the years (I personally can date my 6 meter days continuously back to 1964 as ON5YE and 9A1LJ/M1Q- the old San Marino calls- the torch will have to be carried by younger blood I'm afraid, if serious 2

meter transatlantic studies are to take place.

But still, in a haphazard manner, on a sporadic "ham radio" basis, the beacon does transmit towards Europe. If 6 opens this season, who knows?

I often monitor 50.110MHz (very easy and painless since the squelch can be set and it rarely breaks open and therefore doesn't annoy the family) and so do others who can also activate it. 73's Larry, KL7GLK

P.S. There is also a 6m beacon running with 4W ERP and a turnstile antenna on 50.075 MHz, callsign also KL7GLK.

New Beacon List

On the following pages you will find a completely new beacon list for 6m, 4m, 2m and 70cms. Many hundred hours were spent to check the data and verify them. If you find differences to other lists you can be almost sure that the data from the DUBUS list are the correct one's because in most cases we checked the data personally with the beacons's keeper! We will try to make the same effort for the microwave beacons during the next 3 months. So please send any corrections and updates to info@dubus.de TNX! DL8HCZ

Stations, OPs and Expeditions

Cayman Islands, ZF1A.

ZF1DC and K4BI, along with at least 3 more ops will be doing an ARRL June VHF QSO Party contest effort on 6M, 2M & 2M EME from an excellent site on Grand Cayman. The contest crew will be arriving several days prior to the contest (2nd full weekend in June) to set up and staying a few days afterward. They will pay attention to the EU path on 6. Callsign has yet to be finalized, but the C.A.R.S. has kindly offered us ZF1A should we wish to use it. Jim, K4BI / ZF2BI

Bermuda, VP9

N0JK will be returning to VP9 for the ARRL June 2004 VHF QSO Party. Probable dates are June 10 - 15. The purpose of the trip is the contest, but he will keep an eye out for Europe. Station will be 100 watts and the 6M5X on 6M, 100 watts and Jr. Boomer on 2M. Jon, N0JK

W3CMP will be going back to VP9 from June 26th to July 5th, 2004. Will be at VP9GE's again. Purpose of trip is to use up rest of QSL cards I had printed, make sure 2 mtr VP9DUB beacon is operating properly and to try to work areas missed last time, e.g., UK and Northern Europe. I will ask everyone who has worked VP9 previously to standby so those who haven't can get a crack at it. Should have "adequate" antenna and power on 6 mtrs; hope to have a bigger antenna on 2 meters and have WSJT running properly this time. 73 Chris Patterson, W3CMP

3B9, Rodriguez Isl.

From March 16th the 3B9C expedition will be qrv also on 70cm EME. With AO-40 qrt now this may result in more EME....?

PY5CC in GG53

PY5CC, Peter Sprengel, well known 6m DXer is QRT from GG54RE since July 2003 and moved to GG53QW. Peter expects to be qrv again in mid 2005.

SM6CMU/TF Expedition

SM6CMU, Ingo, will be qrv from TF from June 3rd to 10th on 2m for CW/FSK-MS and on 6m as well.

LA/PA2DW Expedition

Dick, PA2DW (ex PA2DWH) plans to be qrv on 2m FSK and CW-MS during the whole month of July from several squares in Norway. The trip will go up to Lofoten Islands. This is a holiday operation and not a "big" expedition.

9A4QV Info

Adam, 9A4QV, who was qrv /mm on 6m from the Mediterranean sea in 2002 and before reports that he has been working offshore in Libya for one year now. So far it was impossible to get a licence from the 5A authorities.

Microwave from a Balloon

A group of Swedish Ops plans to become qrv on all microwave bands from a hot air balloon in summer 2004 from the area JO78/79/89. All bands from 23cm to 47 GHz are planned and it is hoped to reach altitudes up to ~4000m(ASL).

The people involved at the moment are: SM0SBI, SM4LMV, SM7EYW (SK7MW Op), SM4VLM and SM4XDJ who is the organizer for this project. All comments/questions regarding this project can be posted to microbaloont@osa.t.se. All people above are recipients to this email address. A website for this project is online at: www.osa.t.se/microbaloont/

PC7CW/MM

Theo was qrv again /MM since the end of February for 6 weeks from the mediterranean sea. This time he only took his FM handheld for 2m/70cm. So only stations close to the coastlines had good chances for QSOs. Last year Theo worked 700km QSOs on 70cm FM with his handheld and rubber duck off the CT coast.

OH5NYI/MM

DL2SB will be qrv again for one day in July and for one day in August as OH5NYI/MM from the Baltic see between DL and OH on 2m with 50w and HB9CV. Squares like JO85 will be crossed. QRG is 144.300 and 144.288. In the weeks between the /MM trip he will be qrv as OH5NYI from KP31MC on 2m and 6m. No MS operation!

G0KZG/MM Update

Andy, G0KZG/MM, has changed plans again. He will be qrv from a new ship in Mai / June from the North Sea. He hopes to be qrv again on 144 MHz Meteor-Scatter and Tropo, running the usual high power station. Since the beginning of February he was /MM also, but had only a very little rig.

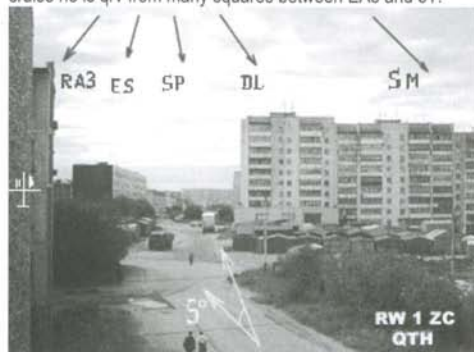
Bornholm Expedition

DL1EJA and others will be qrv from Bornholm (JO75 and 74) from June 10th to 23rd. They plan to operate on 6m, 2m, 70cm and may be 23cm.

RW1ZC/MM qrv again

RW1ZC/MM is qrv again /MM on 2m from March to July off the 5T-coast. In March and July at the start and the end of the

cruise he is qrv from many squares between EA8 and 5T.



RW1ZC take-off from Russia to EU

RW1ZC who is also active on 2m MS from his home QTH has sent the above picture to illustrate his difficult local take-off.

SV/A Mount Athos qrv on 6m now!

Monk Apollo, SV2ASPI/A, made his first 6m tropo QSOs on January 8th. 1st SV/A – SV was SV1DH, of course. He was working with 30w and a dipole. A short yagi was expected to be erected soon. Apollo promised to be qrv in May/June for the Es-season. With his IC756PRRO and IC706 there is still a chance also for a 2m activity in the future.

OH8K Expeditions 2004

OH8K (OH90) has at least following DXpedition-plans for the summer 2004:

June 2004: **KP49, KP39 and KP38**, and in August 2004: **KP18, KP28 and KP27**, both mainly random FSK.

We have looked hill-top QTHs, so QRBs up to 2200km should be possible. Rig: 144 MHz 2x16ele and 50 MHz 5ele, 2x IC-7400. More details later. 73 de Aulis, OH6JW

S0 2004

DL8YHR, Frank, still plans to become qrv in June/July 2004 from S0, Western Sahara on 2m and 6m via EME. So far it was not possible to get a licence for 6m. Also the recent expeditions and S01HA did not have a 6m licence....

Alternatively Frank is checking the possibility of activating Gibraltar, ZB.

Helgoland qrv

From April 23rd to 25th there will take place an amateur radio meeting on the island of Helgoland (Loc. JO34we). The special station DA0HEL will be qrv on April 24th on 2m and may be 70cm (no MS operation!). The OP can be reached via mobile phone on this day at: +49 178 5791777

CT1HZE qrv im summer 2004

DL8HCZ will be qrv again on 2m and 6m in from Mai to mid July from IM57nh. I will be looking especially for Es-openings. Hope to work many of you (again). Joe, DL8HCZ/CT1HZE

23cm from Ireland

EI5FK, IO51, plans to be qrv on 23cms soon with 10w and one yagi antenna.

JP32 on 2m

LA2YMA is qrv on 2m via FSK-MS from the rare JP32lk square. Rig is 40w and a 9 ele. Yagi.

JP93 on 2m

SM3JBO is qrv on 2m FSK-MS from JP93ih.

Azores on 2m

CU2IJ, Luis, is a new station who will be active on 2m FSK-MS soon with 70W and 12 ele Yagi. Locator is HM77

HB0 qrv in April

HB0/HA5KHC/p will be qrv from Liechtenstein (JN47) on 2m and 70cm Tropo from April 1st to 8th.

OY qrv in Mai

ON5DRE and ON4QJ will be qrv from Faroer Islands on 2m via Tropo from Mai 3.-13.

T98KFN activity

CT1FKN is qrv now from Bosnia-Herzegovina until August 6th as T98KFN on 2m and 70cms.

IO65 qrv in Mai

MM0BQI/p will be qrv the first week of May on 6m from IO65.

Comments

15 W ERP for German beacons?

In the end of December 2003 the German authorities have released a new draft for new amateur radio regulations in Germany ("AFUV 2004"). According to this draft all unattended and automatic stations have to be limited to 15 Watt ERP in the future. Also beacon transmitters would be affected if this becomes law. Several beacons for propagation research (Aurora, FAI, Meteor-Scatter) work with more power and do not make sense anymore with 15W ERP.

This is only one very bad item of the draft. Many other points worsen the future of amateur radio in DL either. One of them is the following. (DL8HCZ)

No general 6m allocation in DL

According to the above mentioned draft it is not planned anymore to release the 6m band to all German radio amateurs on a secondary base. In a previous draft from spring 2003 this was planned. The new draft suggests to keep the status quo with a few thousand special permits for 6m.

With almost all countries around DL giving free access to 6m and more and more also permitting 4m, I can only state "Poor Germany!" (DL8HCZ)

10 GHz for Novice Licences in DL

According to the above mentioned draft of new amateur radio regulations in DL it is planned to release the complete 10 GHz band to the Novice class (DO prefix) with max. 10 WEIRP. So far the DO's were allowed to work only on 2m and 70cm. At least one positive news in the draft.... Anyway it is not sure if and when this will be realized.

Moon Station

According to the latest plans from U.S. president George W. Bush a manned station on the moon will be realized within the next 10 years or so. Unless this is not just propaganda for this years elections this will generate a good chance to erect beacons on all VHF/UHF/SHF frequencies on the moon. May be a group in the U.S. should be formed in order to realize this as a long term project? (DL8HCZ)

Echolink QRGs in DL

There is some pressure now in Germany on the DARC radio club to name frequencies or a frequency range for FM-Echolink stations. May be also duplex frequencies will be announced although this makes not much sense. Their input QRGs may be very close below 432.000 MHz. So this may cause some trouble for EME-Ops.

DD0VF comments on WSJT/Toplists:

Jürgen - DK3WG unterbreitete ja den Vorschlag, WSJT-"QSO" bei den Toplisten auszuschliessen und es wurde um Kommentare gebeten.

Ich kann Jürgens Meinung natürlich nur voll und ganz unterstützen. Es wäre schön, wenn wenigstens 1 Topliste ohne Maschinen-QSO übrig bleibt. Für die, die es brauchen, kann ja eine "nur wsjt-toplist" ins Leben gerufen werden.

Sicher fühlen sich dann diejenigen auf den Schlipps getreten, die ohne Spaß und zähneknirschend WSJT machen, damit sie noch paar neue Felder abbekommen....aber was soll's.

Es wird nicht mehr lange dauern, bis eine der nächsten Versionen ein "Single-Ping-QSO" in 500ms erledigt...was soll dann noch der Vergleich in einer Topliste?!

Auch wenn (offiziell) die Software keine automatischen QSO unterstützt, ist das sicherlich das kleinste Problem....

Es wird ja immer behauptet, das die neuen digi Betriebsarten Newcomer schafft...aber es gibt auch die andere Seite....mir hat das alles so ziemlich den Spaß am Amateurfunk genommen, so dass ich nur noch gelegentlich qrv bin - 1 x im Monat EME natürlich in CW ☺

73 de Steffen - DD0VF, JQ61WB 4x7el 1.5w 550W

(Ed.: This was the only comment in favour the exclusion of WSJT-QSOs from the toplist. Many others asked us to keep digital modes included. So the final decision is that there will be no exclusion of WSJT QSOs from our toplist)

Dates

Nordic VHF Meeting 2004

The Nordic VHF-meeting will be held on June 11-13th, 2004. The meeting will take place on Åland islands (OH0), on the island of Eckerö. Organizer is the RATS, Radio Amateur Technology Society in Finland.

Microwave Update 2004

Microwave Update 2004 will be hosted by the North Texas Microwave Society and will be held in Dallas near the DFW airport on October 14, 15, and 16. A formal announcement with all the details is on www.ntms.org and www.microwaveupdate.org (W5LAU/WA5VJB/WA5YWC)

G3WOS 6m BBQ in August 2004

G3WOS, Chris Gare, organizes a 6m BBQ on August 7th. For more info go to http://www.gare.co.uk/6m_bbq.htm. Location is close to London.

VHF-Meeting Weinheim 2004

The next Weinheim VHF Meeting (UKW-Tagung) will take place now definitely on **August 28th/29th 2004**. Same location like in 2003. For more info check www.ukw-tagung.de

News

3.4 GHz band in 9A

On December 26th 2003 the 3.4 GHz band was released to Croatian radio amateurs. Max. Output 10 Watts in the range 3400 to 3410 MHz.

Most Wanted Squares 2004: 2m + 6m

Like in the previous years DL8EBW, Guido, made together with the German magazine Funk-Telegramm the usual poll for the most wanted squares on 2m in EU.

Here a preliminary result after 105 entries:

- | | | |
|----------|----------|-----------|
| 1.) KN38 | 2.) KO67 | 3.) KO42 |
| 4.) KO41 | 5.) KO31 | 6.) KN20 |
| 7.) KO40 | 8.) KO43 | 9.) KO44 |
| | | 10.) KO79 |

For 6m IW0GPN made a similar poll. Here the preliminary result (21 participants):

- | | | |
|----------|----------|-----------|
| 1.) KO22 | 2.) JO79 | 3.) JP30 |
| 4.) IM79 | 5.) JP21 | 6.) KN68 |
| 7.) JM29 | 8.) JN96 | 9.) KN49 |
| | | 10.) IM78 |

MS-VHF-Database: Data request

Hallo dear VHF meteoroscatter friends,

as you possible know, the VHF-DX-Group DL-West has built up a VHF-DATABASE, to spread infos about activities and locations of active DXers. A part of our big VHF-DATABASE is the METEORSCATTER-DATABASE

Therefore it is necessary to ask from time to time about your details and data-updates. We will try to spread around a new issue of these MS-DATABASE (version 2.05) in May 2004 and a total update of our VHF-DATABASE (version 2.10) in second half of the year 2003! For more details check:

<http://www.dl8ebw.de/DATABASE-Project/DATABASE-project.html>

Please fill in your details like in the sample further down in this info, or send info about call / locator of new active stations on VHF/UHF... Please be so kind to send it back to the editor of this DATABASE:

DL8EBW, Guido Juenkersfeld, c/o VHF-DX-Group DL-West via PR-BBS-System: DB0LJ.#RPL.DEU.EU

via Cluster-System: DB0SUE-7 via e-mail: guy@dl8ebw.de
The MS-DATABASE will be given out as text file (zipped) and sorted file (zipped) for use with the WSJT programm from K1JT and will be free of charge to everybody!!!

BE CAREFULL!!! These MS-DATABASE files can NOT be used for using it into logging programmes (VQ-Log) or into DX-Clusters (therefore you do need our VHF-DATABASE, check announcements before!)

A sample for a complete input to the MS-DATABASE:

Call: Rufzeichen: AB1CDE
OLD CALL: Altes Call: DE3EF
LOC-WW: Lokator: JO55AB
Activity: Aktivitaet: 50-144-ms-FSK441-JT65-WWConvers...
RIG: Equipment: 100Watt 11el MGF1302
NOF/CW: CW-QRG: 144.155 30000pm DTR
NOF/SSB: SSB-QRG: 144.355
Update: Update: 02/04

NOF CW/SSB is the normal-operating-frequency which is used as sked frequency at MS, EME, contest etc.!

For MS work you also can determine the max. speed and the way of decoding (eg. MSDSP, DTR ...) you will use!

73 de DL8EBW, Guy, c/o VHF-DX-Group DL-West

No more 6m in India

According to VU2ZAP all 6m licences in India are expired and there will be no new ones for the time being. The licences were given out in order to give a possibility to the VU amateurs to take part in the good conditions of the sunspot maximum which is over now.

220 MHz Band in Canada

Due to the high pressure from the industry the 220 to 222 MHz amateur radio band will probably taken away from the amateurs very soon. Instead the will probably a secondary allocation from 219 to 220 Mhz. The U.S. amateurs had this pleasure already in 1991.

10 GHz in Hong Kong

The VR2-amateurs get a new allocation for amateur radio from 10.450 to 10.500 GHz.

DF5AI Websites

Volker, DF5AI, has relaunched his websites. There are now two sites: One for propagation studies and another for his beamfinder program. For more info check www.df5ai.net

GHz Info- and Skedpage

There is a new page on the web where infos and skeds for activities on the GHz band may be posted, especially before contests. Check www.amram.de

W6PO Silent Key

On January 11th 2004 Bob Sutherland, W6PO became a Silent Key. Bob's PA designs for the VHF bands with the 8877 tube is still famous all over the world. W6PO worked for Eimac for 48 years and was active on EME since the 70s. - - Am 11.1.2004 verstarb Bob Sutherland, W6PO. Weltbekannt wurde W6PO durch seine PA's mit der Röhre 8877 für die VHF-Bänder. Das „W6PO-Rack“ wurde so zu einem festen Begriff. W6PO war seit 48 Jahren bei der Firma Eimac beschäftigt und er war seit den 70er Jahren via EME aktiv.

4m-Band-News

4m band released to all OZ hams

On February 1st 2004 several segments in the 4m (70 MHz) range were released to Danish radio amateurs. Special licences are not longer issued anymore. Max. Output power is

25 Watts. The segments (secondary allocation) are:

QRG	Bandplan
70012,5 – 70062,5	CW + Beacons
70087,5 – 70112,5	SSB + CW
70312,5 – 70387,5	all Modes
70412,5 – 70487,5	FM

4m band in Croatia

In Croatia the 4m band was released to amateur radio on December 26th 2003. Max Output is 10 Watts in the range 70.000 to 70.450 MHz. Already QRV are 9A2EY, 9A3AB, 9A4EP and 9A4QV all from JN75.

4m beacon in Greenland

There are plans to built a 4m beacon for OX. QTH will be in HP15EO at the location from OX3SIX. The beacon may be operational in 2005 or 2006

4m band in Finland

There is coming information from Finland (OH) that there is a good chance for access to 4m in 2004.

Last

DUBUS Rating

On the websites from eham.net one of the most important portals for amateur radio worldwide I have found the following comments about DUBUS Magazine. Reading this was really great joy for me. What could be a better recommendation for DUBUS magazine? (DL8HCZ)

NEOP Rating: 5/5 Nov 9, 2003 02:35

Excellent VHF/UHF magazine Time owned: 3 to 6 months

This is an excellent VHF/UHF magazine, based on the last year of issues which is the only ones I have seen. It is written at a high level, much above QST, Worldradio, or CQ, but is still very understandable. Probably like a combination of QEX and NCJ for VHF. Contains excellent information on all frequencies above 50mhz, including the upper microwaves, and it is neat to see that people are actually playing up there.

It is also nice to read a European perspective to amateur radio. You quickly realize that the proportion of EU hams on VHF/UHF greatly outnumbers the proportion in the US. After reading an issue of DUBUS, I feel like selling all my HF gear and getting a 1296mhz radio.

If you are really curious about VHF/UHF propagation, equipment, and operating, this magazine is highly recommended.

NI6G Rating: 5/5 Jun 7, 2003 18:19

Time owned: more than 12 months

Out of all the amateur magazines that I subscribe to, I look forward to reading DUBUS the most. The technical articles are of the highest caliber and the activity reports are exciting. When I pursue this journal, I feel that I am vicariously living the life of a European weak signal operator. The interest level in VHF DXing seems to be much higher there than in the US. I've read many accounts that were exclusive to DUBUS and not found anywhere else including the internet.

Redaktionsschluss für DUBUS 2/2004

ist der 5. Juni 2004.

Deadline for DUBUS 2/2004 is June 5th 2004

Inhalt / Contents 2003

Reports (Technique, Operation Technique, Propagation, Software)

A Simple TWT Tester	Marko Cebokli, S57UUU	I/8
Locking VCXOs to 10 MHz – Any Frequency, Universal Version	Luis Cupido, CT1DMK	I/15
Locking VCXOs to a 1pps GPS Output for the Microwave and mmWave Local Oscillators	Luis Cupido, CT1DMK	I/18
Analysis of the OK1DFC Septum Feed	Paul Wade, W1GHZ	I/22
Information and practical hints for the construction of a Septum Feed	Zdenek Samek, OK1DFC	I/39
A 145/435 MHz Transmission Line Diplexer designed with Harmonica (Part 1)	Gyula Nagy, HA8ET	II/9
A User's View of the DSP-10, B. Larkin, W7PUAs innovative 144 MHz DSP-TRX	Roger Rehr, W3SZ	II/20
SAM – Simple Aurora Monitor. Buiding your own magnetometer	K. Hansky, DL3HRT	II/36
L and S band LNA with enhancement mode pseudomorphic HEMT	Goran Popovic, AD6IW	III/9
A short backfire antenna (2.4 GHz)	Sergio Cicerone	III/17
A 145/435 MHz Transmission Line Diplexer (Part 2)	Gyula Nagy, HA8ET	III/29
Receiver Dynamic Range	Leif Asbrink, SM5BSZ	IV/9
A 1W Power Amplifier for 24 GHz	Philipp Prinz, DL2AM	IV/40
New Semiconductors for DB6NT's Mixers	Jürgen Dahms, DC0DA	V/45
Update / Corrections for the 13/2cm-LNA (DUBUS 3/03)	Goran Popovic, AD6IW	IV/47

Parts and Products

Microwave Parts LT5504, LTC5507, LTC5505	Henning Weddig, DK5LV	I/48
Sources for Microwave parts	Henning Weddig, DK5LV	I/21
Microwave Parts MMIC multipliers for 10 GHz	Daniel Upström, SM6VFZ	II/54
Microwave Parts LT5512	Henning Weddig, DK5LV	III/41
Microwave Parts ECP053/103/203/100/200	Henning Weddig, DK5LV	IV/47
New Products	DL8HCZ	I/86, II/94, III/92

News, Activity Reports, Contests etc.

2m-EME News	DM2BHG	I/62, II/62, III/70, IV/71
70cm + up – EME News	DL7APV	I/67, II/69, III/74, IV/76
European EME Contest 2003, Results	F6HYE	II/57
FAI News	F6DRO	I/70, II/74, IV/70
FAI on the Edge	DF5AI	I/71, II/74, IV/70
Microwave Europe	GM4PLM	I/58, II/55, III/43, IV/48
Microwave USA	WA5VJB	I/61, II/56, IV/56
Microwave Japan	JH6RTO	IV/54
6m News	CT1HZE/DL8HCZ	I/75, II/79, III/84, IV/87
Aurora Reports	LA0BY	I/76, II/82, III/88, IV/89
Meteor Scatter Reports	DL8LAQ	I/77, II/91, IV/88
Meteor Shower Calendar	DL8EBW	I/85, II/93, III/91, IV/89
Es-News	DL8LAQ	II/86, III/46, IV/57
Tropo News	DJ8ES	I/73, II/75, III/78, IV/82
Toplist	DL8LAQ	IV/107
News & Comments	DL8HCZ	I/88, II/95, III/93, IV/104
Lunar Weekend Calendar 2004	G3SEK	IV/80
European EME Contest 2004, details	F6HYE	IV/81
DUBUS Poll, Results	DL8HCZ	I/94
6m Digital Meteor Scatter at 4U1ITU	F6IRF	II/80
RW1ZC/MM Report	RW1ZC	III/82
GM4COK/MM Report	GM4COK	III/86
A Review of the 2003 Rain Scatter Season	G4KLX	IV/51
DX Opportunities in 144 MHz Double Hop	DF5AI	IV/66
Sporadic E- Ground Reflection of Radiowaves by Lakes and Rivers		
DUBUS Contents 2002	DL8HCZ	I/97
VHF North America	DL8HCZ	III/45

DUBUS Beacon List 50 MHz and 70 MHz, March 2004, by DL8HCZ

QRG	Call	Locator	Pwr	Antenna	QTF	m asl	Mode	Status	Last Report	Remarks
50000	GB3BUX	IO93BF	25	2x turnstile	Omni	460	A1	TQRT	204	
50000	9A1CAL	JN86EL	1	7-ele	Vary	15	A1	24	0803	
50001.0	VE1SMU	FN84	25	3-el Yagi	90		A1	24	1203	
50001	IW1RCT	JN44FH	2	3 El Yagi	180			22-06z	0104	IW1RCT QRG 50007.5?
50001.4	BV2FG	PL05RA	3	5/8 GP	Omni		A1	24	0903	Probably QRT now
50001	BX2AB	PL05SA	10	GP	Omni		A1	?	0703	Probably QRT now
50001.3	IW3FZQ	JN55UW	8	1/4 GP	Omni		A1	24	1003	
50001.5	IK1PLD/B	JN43CW	0.1	Vertical	Omni		A1	24	0503	IK1PLD
50002.5	7Q7SIX	KH75	5					24	0903	
50004.0	IOJX	JN61HV	10	5/8 Vert	Omni		F1	24	204	
50004.7	4N0SIX	KN04FU	1	Dipole	NNE/SSW	400	F1	24	104	
50005	ZS2SIX	KF25UX	25	Dipole			F1	24	0204	ZR2DX
50005	9M2TO	OJ05DJ	50	5/8 GP	Omni		A1	24	0104	
50006.2	PY0FF	HI36TD	30	Ringo Vert	Omni		A1	24	0203	
50007.8	DU1EV	PK04MP	10	1/2 GP	Omni		A1	24	104	
50008	XE2HWW	DL44UC	10	Vertical	Omni		A1	24	204	
50008.0	K0GUV	EN26	8				A1	24	1103	
50008	I5MXX	JN53JU	10	3/4 Vert	Omni	20	A1	24	0304	I5MXX
50009.5	P43L						A1	QRV?	1103	
50009.5	PY2SFY	GG77GA	4	5/8 GP	Omni		A1	24	1103	
50010.8	SV9SIX	KM25NH	30	Vert. Dip.	Omni		F1	24	0903	
50010	JA2IGY	PM84JK	10	5/8 GP	Omni		A1	24	0204	JA1VOK
50011	TG9RB/6	EK44SM	30	V-Ringo	Omni		A1	24	0104	TG9RB
50012	OX3SIX	HP15EO	100	Dipole	E-W	1500	F1	24	1103	
50012	HP1RCP	FJ09HD					A1	24	0303	
50013	CU3URA	HM68QM	5	5/8 V	Omni		A1	TQRT		
50013	LZ1JH	KN22TK	1	GP	Omni	520		24	0204	LZ1DP
50014	V73SIX	RJ39RJ	70	Pan Loop			A1	24	0204	
50014.5	9M6SMC	OJ85AX	2	Turnstile	Omni			24	0304	
50015	9Y4AT	FK90ER	25	3-el Yagi	N/NW		A1	24	1203	
50015.3	SR3SIX	JO92DF	5	1/4 GP	Omni		A1	24	0902	
50015.5	LU9EHF	FF95	15	Inv. Vee			A1	24	1103	
50015.5	VA2MGL	FN47UQ	6	1/2 Vert	Omni		A1	24	0703	
50016	SV5SIX	KM46CF	0.2	dipole	Omni			0204		SV5BYR
50016	GB3SMG	IO91QP	25	V-dipole	Omni	195	A1	Proposal		
50017	JA6YBR	PM51RT	50	Turnstile	Omni		A1	24	0204	JA1VOK
50018	VE4ARM	EN09MW	50	3-el Yagi	NE		A1	24	104	
50019	IZ1EPM	JN34WR	15	5/8 Vert	Omni	400	A1	24	1103	
50019	CX1CCC	GF15	5	GP Omni				24	0304	Temp QRT
50020	GB3SIX	IO73SG						Proposal		
50020.7	IK5ZUL	JN52					24	104		
50021	S55ZRS	JN76MC	8	1/4 GP	Omni	1219	A1	24	1103	
50021.3	OZ7IGY	JO55VO	30	Turnstile	Omni	92	A1	24	204	
50023	XE1KK	EK09QC	20	Ringo	Omni		A1	24	204	
50023	LX0SIX	JN39AV	10	Horiz Dipole	N/S			24	104	
50023.5	ZP5AA	GG14	5	Vertical				24	1203	
50023.0	JA1ZYK	QM05	10	Hentenna				24	1203	
50023	SR5SIX	KO02LL	2	Dipole	SW/NE	100	A1A	QRT	0104	soon qrv agn. SP5XMU
50024.7	UN1SIX	MM83KE	12	GP	Omni	3400		OP?	0402	
50025	9H1SIX	JM75FV	7	5/8 GP	Omni		F1	24	104	
50025	YV4AB	FK50	15	Ringo			A1	24	1203	
50025	VE4SPT	EO26PI	100	3 el yagi	N		A1	INT	0304	KG0VL
50025.4	OH1SIX	KP11QU	40	4 Turnstiles	Omni	157	A1	24	0104	
50026	VP9DUB	FM72OH	15	Horiz Loop	Omni		A1	24	0204	
50026	SR9FHA	KN09BW	4	GP	Omni		F1a	24	0803	
50027	CN8MC	IM63NX	20	J-Pole	Omni			24	0104	

50027	JE7YNQ	QM07	2	T'stiles	Omni	A1	24	0204	JA1VOK
50027.8	SR6SIX	JO81HH	7	GP	Omni	F1	24	0603	
50028	XE2ED	DM10	15	1/2 Vert.	Omni	A1	24	204	
50028.0	SR8SIX	KN19CN	10	Dipole		A1	?	0203	
50028	5T5SN/B	IK28AC	25		Omni			Proposal for summer 2004	
50029	P29BPL	QI30	25				24	NonOp	
50029	CT.....	IM59						2/04	Proposal CT1DDN
50031.0	ZD8VHF	II22TB	50	Dipole	Omni	723	24	2/04	
50031.5	CT0SIX	IM59	10E	Dipole	E/W	F1	24	2/04	
50032	JR0YEE	PM97	2	Loop		A1	24	0204	JA1VOK
50033	OH5RAC	KP30HV	26E	2dBD	200	145		Proposal for spring 2004	
50033	VE2RCS	FN25	5			A1	24?	1203	
50033	VE7FG	CO83SW	50	1/4 GP	Omni	1051	A1	24	104
50034	YU1EO	KN04ML	1	GP	Omni		A1	24?	1003
50034	OX	GP47						Planned	
50035	ET3VSC	KJ99IA					24	0703	Call will change to ET3BN
50035	HC8GR	EL59	30	Halo			A1	24	1203
50035	OY6SMC	IP62MB	15W	Vertical	Omni		F1	24	1203
50036	VE4VHF	EN19LM	35	Vertical	Omni		A1	24	0104
50036	OA4A/B	FH17	13	Vertical	Omni		A1	?	0204 hrd from LU
50036.6	SR2SIX	JO93BC					A1	QRT	0702
50037.5	ES0SIX	KO18CW	15	Horiz Dip.	E/W	35	A1	24	1203
50038	VA2MGL	FN47UQ	5	Quad-Turn.	Omni		A1	24	0104
50038	C21SIX	RI39LL	25	Dipole	E-W		F1	24	0104
50039	VO1ZA	GN37	10	1/4 Vert	Omni		F1	24	0803
50039	FY7THF	GJ35	10	Vertical	Omni			24	0104
50040	ZL3SIX	RE66	65	2xDipole	NW/NE		A1	24	0104
50040	SV1SIX	KM17UX	25	Vert Dip	Omni		F1	24	0104
50041	VE6EMU	DO33	35	4-el Yagi	NNE		A1	24	0104
50041	ON0SIX	JO20EP	5	X-Dipole	N-S		A1	24	1103
50042	YF1OO	IO33	15	GP				EXP	1103
50042.5	GB3MCB	IO70OJ	40	Dipole E/W			F1	24	0104
50043	VE6ARC	DO05	25	GP	Omni		A1	24	0104
50043.3	ZL1VHF	RF73MB	20	1/2Vert@4m	Omni		A1	24	0204
50044	ZS6TWB	KG46RD	15	5/8	Omni		A1	24	0204
50044.5	LZ2CM	KM13NE	0.5	GP	Omni	1500	A1	24	0204
50045	OX3VHF	GP60QQ	20	GP	Omni	20	F1	24	1203
50045.6	JW5SIX	KQ26MM	10	Dipole	Omni		A1	24	0204
50046.5	4N1ZNI	KN03WH	1	GP	Omni	200	F1	24	1203
50046.5	VK8RAS	PG66	20	X-Dipoles	Omni	300	A1	24	0104
50047	JW7SIX	JQ68TB	10	3-el Yagi	S	10	A1	TQRT	0204
50048.2	VE8BY	FP53RS	25	Vert Dip	Omni		F1	24	0104
50048.5	TR0A	JJ40	15	5-el	N		F1	24	0204
50048.7	JW9SIX	QJ94LM	10	GP	Omni		A1	24	0204
50049	VA3BCN	FN03	3	Horiz Dipole			A1	24	0701
50050	LZ2CC	KN22GS	5	GP	Omni	1500		24	0204
50050	VA2WW	FN46	5				A1	?	0603
50050	FO5DR	BH52	50	Dipole				NonOP?	
50050	ZS6DN	KG44DE	100	5-el Yagi	SE		24	0204	ZR2DX
50051	PY1AA	DG87IC					24	0603	
50051	LA7SIX	JP99EC	30	4-el Yagi	190		A1	24	0104
50052	SK2CP	KP07MU	10	5-el Yagi	NW	478	A1	24	TNonOp / Proposal
50053.5	PI7SIX	JO22NC	12	Dipole	N/S	40	A1	24	0803
50053.6	XE3RCM	EL50	1	Dipole			A1	24	0102
50055	OZ6VHF	JO57EI	1	Vertical	Omni		A1	24	0204
50055	V44KAI	K87QH	2	5/8 Vert	Omni		A1	24	1203
50056.5	VA7SIX	CN89	10	Ringo	Omni		A1	OP	1203
50057.0	VK8VF	PH57	100	Delta Loop	Omni		A1	TQRT	0104
50057	VK7RAE	QE38	25	X-Dipoles	Omni	220	A1	TQRT	0104
									WIA location change

50057	TF3SIX	HP94BC	110	GP	Omni	A1	24	0104	
50058	K3DEL	FM28FN	5	SQLoop	Omni	33	A1	INT	1203
50057.5	VK4RGG	QG62	10	Turnstile	Omni	A1	24	0104	W3OR Intermittant WIA
50058	HB9SIX	JN47QG	0.2	J-Pole	Omni	A1	24	1103	
50058	IQ4AD	JN54DT	8	GP	Omni	A1	24	1103	
50058	LZ2TU	KN13SF	5	Dipole				0803	Probably QRT now
50058	OE3XLB	JN87AT	10	5-el Yagi			24	0204	
50059.3	PY2AA	GG66	25	GP@60ft	Omni	850	A1	24	0802
50059.5	VE3UBL	FN03	10	Turnstile	Omni	A1	24	1003	Probably QRT now
50059.5	K4TQR	EM63OM	4	Dipole	NW/SE	A1	24	1203	
50060	KE4WBO	EL96VV	10	Loop horiz	Omni	20	A1	24	0104
50060.0	KD4NMI	FM05RT	100	Vert.Dip.	Omni	A1	24	0204	grv when OP not on air
50060.0	GB3RMK	IO77UO	40	Dipole	N/S	270	F1	24	0104
50060.0	W5VAS	EM50	100	SqualoOmni		A1	24	0204	
50060.0	WB0RMO	EN10	25	Squalo		A1	24	1103	
50060	K5AB	EM01				A1	24	0104	
50060.5	KD4AOZ	EM83				A1		1002	
50061	K9MU	EN44	7	Dipole	N/S	A1	24	0204	
50061	N7SCQ	CN98				A1	?	0903	
50061.7	KH6HME	BK29	20	Dipole		A1	24	0104	
50062	GB3NGI	IO65VB	10		Omni	518		02/04	Proposal
50062	K8UK	EN82JP	2	Omni		A1	24	0104	
50062	KB6BKN	CM88	30	3 Ele	E			0703	
50062	C6AFP	FL16	1	5/8 Ringo	Omni	A1	24	0204	
50062	T99YVZ	JN84UD						0603	
50062.5	W7HAH	DN28NB	25	Stack Halos	Omni	A1		0104	QRT
50062.5	WB4WOR	FM05BT	2	GP@3m	Omni	A1	24	0602	
50063	W9JN	EN54DN	1,10	Halo	Omni	A1	24	0104	
50063	LY0SIX	KO24PS	7	Dipole@5m				0901	QRT
50063	KF4ODI	EM56	10	Loop@40'		A1	24	0902	
50063	KE4SIX	EM83	5	GP	Omni	A1	24	0104	W4WTA
50063.5	NL7Z	BP51	20	Vert Dip	Omni	A1	24	1202	
50063.5	K7UV	DN31XM	40E	H. Loop	Omni	A1	24	0104	K7UV
50064.0	W3VD	FM19NE	7	Squalo	Omni	A1	24	0104	
50064	GB3LER	IP90JD	30	Horiz.Dip.	N/S	100	F1	24	0204
50065	KL7/KG0VL	BP75XW	100	Horiz.	Omni	A1	24	0204	KG0VL
50065	HK3BZO	FJ24					?	0603	Probably QRT now
50065	GB3IOJ	IN89WE	10	Dipole	E/W	115	F1	24	1103
50065	VQ9SIX	MI62	10	Vertical	Omni	A1	24	1103	
50065.0	W0IJR	DM79	20	2-ring halo	Omni	A1	24	0204	ex Call KA0CDN
50065.2	KH6HI	BL01	50	Halo	Omni	A1	24	0104	
50065.5	W0MTK	DM59	4	4 Veess	Omni	A1	24	0104	
50066	W5GPM	EM26						0703	
50066	VE9MS	FN65RR	0.25	Ringo	Omni		F1		QRT
50066.0	VK6RPH	OF88	10	U Dipole	Omni	366	A1	24	0104
50066.2	W5SIX	DM54WA	1	5-el Yagi	060	A1	24	0104	WIA
50066.5	KA5FYI	EM10	1			A1	INT	0702	
50066.6	WA1OJB	FN54	10	2-el Coll.	EU	A1	24	0803	
50066.7	KQ4E	EM86	10	halo	Omni	A1	24	0104	
50067	WZ8D	EM79	15	6-el@75'	S-Am	A1	seasonal	1003	
50067	OH9SIX	KP36OI	35	2x X-dipole	Omni	192	A1	24	0104
50067	N7DB	CN85RM	5	halo	Omni	A1	Days	0203	
50067.5	W3HH	EN90	10	Horiz Loop		A1	24	0104	
50067.5	N8PUM	EN65BX	10	Dipole	Omni	A1	24	0204	
50067.7	K0EC	DM69	45	1/2V+1/2H	E-W	A1	24	0204	
50068	LZ0SIX	KN12QP	20	GP	Omni	670	A1	24	0204
50068	N9UM	EN52WG	10	Loop@8m		A1	INT	0104	LZ1DP
50068.5	K2ZD	FN20	20	5/8 V	Omni	A1	24	1203	new Call KG9PF, soon 24h
50068.8	LU1DMA	GF05PH	1	5el Yagi			INT	0304	

50068.9	K6FV	CM87UL	100		varies 55/275	A1	24	0903	
50069	W9NFL	EN55	8	5 Ele		A1	24	QRT	proposed qrv in spring 04
50069	W7BAS	CN87TG	10	5-el	090	A1	24	1201*	
50069.5	N0UD	DN87	50	3-el@6m	N	A1	24	0104	
50070	W1RA	FN41	15	Vert. Dipole	Omni	A1	24	0703	
50070	EA3VHF	JN01EA					?	1203	
50070	WA7X	DM49HO	100	J-Pole	Omni	A1	24	1203	
50070	N3LEL	FN00	1	Dipole		A1	24	0104	N3LEL
50070	K4AR	EM76...				A1		0104	QRT for QTH+Call-change
50070	W4CLM	EM95				A1	24	1203	
50070	LU4DRM	GF05	15	6dbi Vert	Omni	A1		1002	QRT
50070	KG6JAI	?						0703	
50070.5	W5HN	EM13SJ	0.8	Halo	Omni	A1	24	0104	W5LUA
50070.6	K0ETC	EM27	10	Turnstile	Omni	A1	24	0104	
50071	W7WKR	CN98PI	25	X Dipole	Omni	A1	24	02/04	qrv agn from spring 04
50072	KB8GC						Test	0603	
50072.5	W4IO	EM81	1	Vertical	Omni	A1	24	0803	
50073	PP2SIX	GH53MP	7	GP	Omni	A1	24	1103	
50073	ES6RQ	KO28WA	10	6-el Yagi	240	A1	24	Intermittant	ES6RQ 1/04
50073	W7KMA	FM18GR	1	Dipole N	EU	A1	24	0203	
50073	NN7K	DM09	1	Squalo@15'	Omni	A1	24	0104	
50073	WR7V	CN87	10	Halo	Omni			0503	
50073	K0KP	EN36WT	100	Delta Loop	Omni	A1	24	1003	
50074	FG1JD	FK96ID	10	Dipole		A1	24	0204	FG1JD
50074	KD4HLG	EM73	3/30	AR6		A1	24	0204	
50074	W5RP	DM91						0204	
50074	N7LT	DN45	3	1/4 GP	Omni	A1	24	0303	
50074	K8PLF	EN81						0204	
50074.8	YV5LIX	FK60NM	5	5el Yagi	045	A1	24	1102	
50075	VR2SIX	OL72BH	12/4/1.2/0.4	1/4GP	Omni	450	24	0903	
50075	N4RT	EM60	20	Vertical	Omni	A1	24	0104	N4RT
50075	KA7BGR	CN82	40	2x2el	SW/NE	A1	24	0204	
50075	YO3KWJ	KN34BJ	10	vertical	Omni		24?	0703	
50075	NL7XM/2	FN20WO	1	Squalo	Omni	A1	24	NonOp	
50075.0	KL7GLK/3	FM18SW	4	Turnstile	Omni	A1	24	0204	KL7GLK
50076	WR9L	EN61BD	10	Turnstile	Omni	A1	24	0104	
50076	OH5SIX	KP30HW	50	2x XDipole	Omni	130	TEST	0104	Temp. QRV -> .033
50077	W8MI	EN75					OP?	0602	
50077	K4AHO	EL98FR	7.5	Vert@30'	Omni	A1	24	0204	
50077	W9FZ	EN43				A1	24	1203	
50077.5	VK4BRG	QG48	3	Turnstile	Omni	A1	TQRT	0104	WIA - Attended keyer
50077.6	N0LL	EM09OW	14	2 halos	Omni	A1	24	0204	
50078	AC3A	EM28QV	10	Loop	Omni		24	0104	
50078.0	OD5SIX	KM74	8	1/4 vert.	Omni		24	0903	
50078	PT7BCN	HI06RF	5	GP	Omni		TQRT	0902?	
50078.2	TI2NA	EJ79	50	Vertical	Omni	A1	24	0204	
50078	W8SCA	EM79UN	6	Halo	Omni	A1	EXP/24	1102	
50079	KK7CQ	DM33	10	Squalo		A1		Proposal	
50079	JX7SIX	IQ50RX	10	Slope dip	Omni	269	A1	24	0204 Temp. QRT
50079	W4CHA	EL88	30	Quadloop	Omni	A1	?	0104	
50079.5	XE2KIB	DL99				A1		1002	
50080	K0UO	EM07	10	Omni		A1	24	0204	
50080	ZS1SIX	JN96FB	10	Inv. V Dipole	Omni	F1	24	0204	ZR2DX
50080	FK8SIX	RG37FR	15	GP	Omni	F1	24	0204	
50080	WD8OST	EN76	10	6el@45'		A1	24	1003	
50080.0	UU5SIX	KN74AL	10	dipole		1300	24	0204	
50080.0	4X4SIX	KN72JB	3	dip@15m		50	A1	24	0903
50080.0	W3CCX	FM29JW	4	halo@500'	Omni	A1	24	1103	
50080	PP1CZ	GG99UQ	3	5-el	N		24	0903	

50081	LZ2SIX	KN33						0503	Probable QRT now
50081.5	LU3EMK	GF05				A1	Not24	0302	
50082.5	LU8DCH		1.5	5/8	Omni	A1	24	1203	
50083.6	LU7YS	FF49IU	5	GP	Omni	A1	24	0903	
50084.0	UT5G	KN66LS	10	GP	Omni	A1	24	0204	
50084.5	K5NZ	EM20	10					0104	QRT
50087	VK4RTL	QH30	10	Vee	Omni	582	24	0104	WIA
50088	HC2FG	EI97	5	J-pole	Omni	?		1103	
50090	ZZ2LER	GG66	10	dipole	Omni	A1	24	0104	
50090.5	WP4F	FK68WL	5	GP	Omni	A1	24	0104	
50203	F6IKY	JN35AT	0.75	5 Ele.	360	700	IRR	0603	
50283	VK3RMH	QF22OH	10	Double Loop	Omni	A1	24	TNonOp	
50288	VK2RHV	QF57	25	Turnstile	Omni	A1	24	0104	
50293	VK3RMV	QF02	15	9-el@37m	N	305	A1	24	0104
50297	VK7RST	QE37	15	Dipole	E/W	F1	24	0204	WIA (ex VK3SIX)
50304	VK6RSX	OG89	50	Hor. dip.	Omni	50	A1	24	0204
50306	VK6RBU	OF76	25	3-el Yagi	80/260	10/100	A1	24	0104
50308	VK6RTW	OF84	10	Horiz		A1	24	0104	WIA
50315.0	FX4SIX	JN06CQ	25	2xHor.Dip	Omni	A1	24	0204	
50321	ZS5SIX	KG50	8	Halo		A1		0204	QRT ZR2DX
50386	F1GTU	JN05IF	5	Loop Omni		A1	?	0104	
50480	JH8ZND	QN02UW	10	5/8 GP	Omni	A1	24	0204	JA1VOK
50485	JH9YHP	PM86	2	13-el SE		A1		0204	JA1VOK
50490	JG1ZGW	PM95VP	10	7el Yagi	S	A1	24	0204	JA1VOK
50499.5	5B4CY	KM64PR	15	¼ GP	Omni	F1	24	1203	
50520.3	SZ2DF	KM25	1kw	4x16 el el.	30	330		EXP/SUMMER	
51030	ZL2MHB	RF80					RTTY	24?	0104
52275	ZL2MHF	RE78NU	10	V ½ Dip	Omni	F1	24	1203	
52315	VK3RSX	QF13IR	10	Turnstile	Omni	A1	24	0104	WIA / VK3PA
52346.5	VK4ABP	QG26	10	¼ Vert.	Omni	480	A1	24	0104
52420	VK2RSY	QF56MH	25	Turnstile	Omni	240	A1	24	0104
52450	VK5VF	PF95	30	Turnstile	Omni	680	A1	24	0104
52490	ZL2SIX	RE68	10			F1	24	0503	WIA

70000	GB3BUX	IO93BF	20	2 x Turnstile	Omni	456	A1	24	07/03	
70007	GB3WSX	IO80QW	2,25	5 Ele		70	105	F1A	24	03/04
700 . .	OX3 . . .	HP15EO								G3ZXX
70010	ZD8VHF	II22TB			N	723			24	03/04
70016	GB3SMG	IO91QP	20	Dipole		195				01/04
70020	GB3ANG	IO86MN	100	3 Ele Yagi		160	370	A1	24	09/03
70021	OZ7IGY	JO65FR	25	3 Ele Yagi		225	40	A1	24	02/04
70025	GB3MCB	IO70OJ	40	4 Ele Yagi		45	320	A1	24	09/03
70028.5	S55ZMB	JN76VK	5	4 Ele Yagi		310	250	A1	24	08/03
70114	5B4CY	KM64PR	15	4 Ele Yagi		315	30	A1	24	08/03
70312.5	SJV900	JO89								Proposal SM0TSC

There are NO 4m beacons from South Africa qrv anymore.

INT= Intermittent Operation TQRT = Temporary QRT EXP= experimental UC=Under Construction
The 6m list is based on G3USF's list.

DUBUS 2m and 70cm beacon list, March 2004, by DL8HCZ

QRG	Call	Locator	m asl	Antenna	QTF	Pwr	Info/Keeper	Last Update	Remarks
144.060	LU1DMA	GF14		Yagi	N	7		02/04	LU1DMA
144.141	EW6DS/B	K045IG					EW6DS	11/03	UA3ARC
144.146	LZ2CC	KN22GS	1480	GP	Omni	5	LZ2CC	02/04	LZ1DP
144.170	KH6HME	BK29GO	2501	2x7 Ele.	to CA	1200e		07/03	
144.280	VE2FUT	FN35		Horiz.		10		10/03	
144.285+/-	VP9DUB	FM72OH		4 Ele	EU	40	DL8HCZ	02/04	Proposal
144.285	VE1CBC	FN63						10/03	
144.289	KL7GLK	FM18		turnstile	omni	5	KL7GLK	02/04	
144.289	KL7GLK	FM18		2 x 13 ele	EU	100	KL7GLK	02/04	Temporary QRV
144.300	ZD8VHF	II22TB		Yagi	EU	50	DL8HCZ	03/04	Proposal
144.302+/-	5T5SN/B	IK28AC		Yagi	EU	50	DL8HCZ	03/04	Proposal
144.341	UT5EC	KN78MM	180	4 x Yagi	Omni	10		01/04	
144.350	RN6MT	KN97LN	120	Ring Dipole	Omni	5	UR5LX	01/04	
144.362	UT5G	KN66LS	80	Dipole	N/S	5		01/04	
144.370,5	UR4YWW	KN28WG	280	Turnstile	Omni	4		01/04	
144.387,5	RA3IM	KO56UM	230	Discone	Omni	9		02/04	
144.400	VO1ZA	GN37JS	90	11 El Yagi	EU	250	VO1NA	10/03	
144.402	UA3FI	KO76WF	160	1/4 V	Omni	5		12/03	DK3WG
144.402	OY6BEC	IP62OA	40	11 El Yagi				12/03	LA0BY
144.403	EI2WRB	IO62IG	248	Vertical	Omni	100e	EI8JA	10/03	DL8HCZ
144.405	EA1VHF	IN53RE	615	Halo	Omni	25	EA1DKV	02/04	CT1HZE
144.405	F5XAR	IN87KW	165	9 eleYagi	290	400	F6ETI	01/04	F6HTJ Temp. QRT
144.406	CT1ART	IM67AH	560	Halo	Omni	5	CT1EPS	03/04	CT1HZE Dirty Signal
144.407	GB3SSS	IO70IA	50	8/8 slot Yagi	300	200	G3AGA	11/03	Transatlantic Beacon
144.409	F5XSF	IN88GS	145	9 el Yagi	90	50	F6DBI	01/04	F6HTJ
144.410	DB0SI	JO53QP	90	Big wheel	Omni	10 TX	DL1SUZ	01/04	DL8HCZ
144.410	RB4NA	LP30VF		9 Ele Yagi	15	5		12/02	
144.411	IK1LBW/B	JN44VC	745	Turnstile	Omni	1	IK1LBW	01/04	DL8HCZ
144.412	SK4MPI	JP70NJ	520	4 x 6 el Yagi	45/315	1500	SM4HFI	02/04	SM6CEN
144.414	DB0JW	JO30DU	238	5 el Yagi	22	50	DL9KAS	01/04	
144.415	ZS2VHF	KF25UX			Omni	10	ZS2FM	02/04	ZR2DX FSK
144.416	PI7CIS	JO22DC	40	Dipole	E/W	50	PA0CIS	02/04	
144.417	OH9VHF	KP36OI	310	10 dBd	200	200		02/04	actually on .415
144.418	ON4VHF	JO20HP	180	Clover leaf	Omni	15	ON7ZV	09/03	
144.419	I2M	JN55AD	46	Big wheel	Omni	10	IK2THZ	01/04	DL8HCZ
144.420	ZS1VHF	JF96						12/03	ZR1AEE QTH
144.420	DB0RTL	JN48PL	480	Big wheel	Omni	15	DL8SDL	01/04	DL8HCZ QRT!
144.421	OZ...	JO46/47		2 x Yagi	300	250	OZ7IS		Transatlantic Beacon Proposal
144.422	DB0TAU	JO40HG	326	4x4 el Yagi	Omni	15	DL3DC	02/04	
144.423	PI7HVN	JO22WW	52	Halo	Omni	10	PA3FHY	12/03	PE2KP ex PI7FHY
144.424	IN3A	JN56NB	225	GP	Omni	0.1	IN3IYD	01/04	DL8HCZ
144.425	F5XAM	JO10EQ	99	Big wheel	Omni	14	F6BPB	02/04	F6HTJ
144.426	EA6VHF	JM08SQ	150	Halo	Omni	25	EA6FB	11/03	
144.427	OK0EJ	JN99FN	1323	4 el Yagi	270	0.3	OK2UWF	12/03	DL8HCZ
144.427	PI7PRO	JO22NA	10	Halo	Omni	10	PI4VRZ	02/04	
144.428	DB0JT	JN67JT	785	4 x Dipole	0	30	DJ8QP	03/99	Probably QRT
144.429	IQ3MF/B	JN65RW	130	2xTurnstile	Omni	4	IV3HWT	01/04	DL8HCZ ex IV3HWT/IV3A
144.430	GB3VHF	JO01DH	268	2x3el Yagi	315	40	G0FDZ	02/04	
144.431	9A0BVH	JN85JO	489	V Dipole	Omni	1		03/95	9A2MP QRT
144.432	9H1A	JM75FV						02/04	9H1BN QRT!
144.433	OH7VHF	KP52IJ	165	2xBigWheel	Omni	50		01/04	OH6DD
144.433,5	HB9OK	JN45MW	1601	Halo	Omni	3,5	HB9DUR	02/04	
144.434	DB0LBV	JO61EH	232	2xDipole	Omni	0.4TX	DL1LWM	09/03	
144.435	HB9H	JN46KE	1600	2xDipole	Omni	0.3/3		06/98	HB9PQX Probably QRT
144.435	SK2VHG	KP07MU	535	16 el Yagi	180	800e	SK2CP	02/04	SM6CEN

144.437	LA1VHF	JO59MS	358	X300 (Vert)	Omni	20	LA4PE	01/04	LA0BY	Temp. QRT
144.438	OK0EO	JN89QQ	602	Ring dipole	Omni	0.05/0.12	OK2VLX	12/03	DL8HCZ	Planned
144.439	OZ3VHF	JO55				0.1		02/04		
144.440	DB0RG	JO51GO	730	V Dipole	Omni	1	DL4OAN	12/03	(D)	
144.440	ZS0VST			8 el Yagi	0	30	ZS4NS	03/00	ZS5JF	Planned
144.441	LA4VHF	JP20LG	30	2x8el Yagi	0	380	LA6LU	01/04	LA0BY	
144.441	CT0TWO	IM59OM		Vertical	Omni	0.25		02/04	CT1HZE	
144.442	IAA	JN54QK	300	4xDipole	Omni	1/10	IK4PNJ	01/04	DL8HCZ	QRT
144.443	OH2VHF	KP20BB	65	2xYagi	20/230	130	OH2LNM	01/04	OH6DD	
144.444	DB0KI	JO50WC	1025	Dipole	Omni	2.5	DC9NL	01/04		
144.444	IQ5LU/B	JN53GW	1000	Big wheel	Omni	6	IW5BHY	01/04	DL8HCZ	ex I5A
144.445	GB3LER	IP90JD	108					06/02	G3UUT	QRT
144.446	OK0EB	JN78DU	1084	Miniwheel	Omni	0.07/0.007	OK1APG	01/04		
144.447	SK1VHF	JO97CJ	65	2xCloverleaf	Omni	10e		02/04	SM6CEN	
144.448	HB9HB	JN37OE	1300	3 el Yagi	345	120	HB9AMH	03/04		
144.449	IOA	JN61		2 x Big wheel	Omni	10	IW0BCF	01/04	I5WBE	QRT
144.450	DL0UB	JO62KK	120	4 x Dipole	Omni	10 TX	DL7ACG	01/04	DL8HCZ	
144.450	F5XAV	JN23GX	100	Big Wheel	Omni	20	F5IHN	01/04	F6HTJ	
144.450,7	RB3PA	KO84UF		8 Ele Yagi	360	1		10/03		
144.451	LA7VHF	JP99EC	320	10 el Yagi	190	20	LA5TFA	01/04	LA0BY	
144.452	OK0EC	JO60CF	778	3 el Yagi	90	0.7	OK1VOW	12/03	DL8HCZ	
144.453	GB3ANG	IO86MN	370	4 el Yagi	160	20	GM4ZUK	02/04		
144.453	IOG	JN63IB	1200	4 x dipole	Omni	10	IW0QIT	01/04	DL8HCZ	QRT
144.454	IW0UMP/B	JN40QW	350	Turnstile	Omni	1	IW0UGR	01/04	DL8HCZ	ex IS0A
144.455	OH5ADB	KP30NN	65	Dipole	135/315	0.1		01/04	OH6DD	
144.456	DB0GD	JO50AL	930	Dipole	0/180	1 TX	DG6ZX	03/99		QRV?
144.457	SK2VHF	JP94TF	300	2 x 10 el Yagi	0/225	100e		02/04	SM6CEN	
144.458	F1XAT	JN15AO	913	Big wheel	Omni	25	F1HSU	01/04	F6HTJ	
144.459	LA5VHF	JP77KI	260	2x6 el Quad	15/180	100	LA1UG	04/01	LA0BY	QRT
144.460	HG1BVA	JN86CW	370	2 x Big Wheel	Omni	5	HA1YA	02/04	HA1YA	
144.460	VK6RPH	OF88	337	Horizontal	Omni	10		01/04	WIA	
144.460	TF3VVV	HP84							Proposal	
144.461	SK7VHF	JO65KJ	25	2xCloverleaf	Omni	10e		02/04	SM6CEN	
144.462	IQ0RI/B	JN62MH	1010	3 Ele Yagi	330	25e		08/03		
144.463	LA2VHF	JP53EG	710	12 el Yagi	15	500	LA1K	01/04	LA0BY	
144.464	IFA	JN81EC	685	Big wheel	Omni	8	IF7NW	01/04	DL8HCZ	
144.465	DF0ANN	JN59PL	630	V Dipole	Omni	0.3 TX	DL8ZX	01/04		
144.465	CN8MC	IM64		5 el Yagi	N	30TX	CN8LI	02/04	CT1HZE	QRT
144.466	OZ4UHF	JO75LD	130	Big wheel	Omni	10	OZ1HTB	01/04	OZ2TG	
144.467	HB9RR	JN47FI	904	4 x Dipole	Omni	0.0225		09/03	HB9RR	
144.467	I8A	JM78WD	1778	SqLo	Omni	8	I8GMP	01/04	DL8HCZ	QRT
144.467	OK0ED	JN99DO	290	2xDipole	Omni	0.1	OK2UWF	12/03	DL8HCZ	
144.468	F1XAW	JN26IX	561	Big wheel	Omni	10	F1RXC	01/04	F6HTJ	
144.468	LA6VHF	KP59AL	70	9 el Yagi	210	300	LA4OO	01/04	LA0BY	
144.469	GB3MCB	IO70OJ	320	3 el Yagi	45	40	G3YJX	02/04	G3YJX	
144.469	IT9A	IM67LX	825	2xBigwheel	Omni	10	IT9QPF	01/04	DL8HCZ	QRT
144.470	OH2VHH	KP20MH	75	Halo	Omni	2		01/04	OH6DD	
144.470	OK0EZ	JN79VV	350	Vertical	Omni	2/0.5	OK1DXF	12/03	DL8HCZ	
144.471	OZ7IGY	JO55VO	96	Big wheel	Omni	25	OZ7IS	01/04	OZ2TG	
144.473	OE3XAA	JN88BA	834	Halo	Omni	0.2	OE1BKW	09/99	OE1MCU	QRT
144.474	OK0EL	JO70SQ	900	Dipole	90/270	0.005	OK1AIY	12/03	DL8HCZ	
144.474,5	YU1VHF	KN04OO	200	2xQ	135/337	10	YU1AU	12/03	YT1MM	
144.475	DB0SGA	JN69KA	1024	4x4el Yagi	Omni	5 TX	DJ4YJ	01/04		ex DL0SG
144.475	LY2WN	KO25DB	95	Turnstile	Omni	5	LY2WN	01/04	LY2IC	actual QRG .473
144.476	F5XAL	JN12LL	1100	Big Wheel	Omni	10	F6HTJ	02/04	F6HTJ	
144.477	DB0ABG	JN59WI	522	Big Wheel	Omni	4 TX	DJ3TF	03/99	DJ3TF	QRT
144.477	SR2VHJ	JO93AD		10 Ele Yagi	270	60		02/04		
144.478	LA3VHF	JO38RA	30	9 el Yagi	180	120	LA3BAA	01/04	LA0BY	
144.478	OM0MVA	JN88NE	570	Dipole	90/270	0.11	OM3ID	12/03	DL8HCZ	

144.478	S55ZRS	JN76MC	1219	Dipole	Omni 1	S57C	02/04	S51KQ	
144.479	SR5VHF	KO20ON		Turnstile	Omni 0.75	SP5TAT	01/02	SP6LB	QRV?
144.479	IQ9BF/B	JM77NO	800	2xBigWheel	Omni 3	IW9AFI	01/04	DL8HCZ	QRT ex IT9S
144.480	LA8VHF	JO48XX	300	3x2 el Yagi	150 100	LA6LCA	01/04	LA0BY	
144.481	HG8BVA	KN06PW	85	5 Ele	270 5	HA8MV	02/04	HA1YA	
144.482	GB3NGI	IO65VB	518	2x4 el Yagi	45/120 25	G16ATZ	02/04	G16ATZ	
144.482	I8EMG/B	*JM89BJ	1400	Halo	Omni 3	I8EMG	02/04	I8EMG (D)	
144.484	UU9JJ	KN64RO					01/04	DL8HCZ	Proposal
144.484	EA8VHF	IL18BS		2 x agi WNW+NNE	50	DL6FAW	03/04	DL8HCZ	Proposal
144.486	DL0PR	JO44JH	75	4x6el Yagi	0/180 200TX	DL8LD	02/04		
144.488	IW1DJS/B	JN45AB			Omni	IW1DJS	02/04	IW1DJS	
144.490	DB0FAI	JN58IC	590	16 el Yagi	305 30e	DL5MCG	02/04	DL5MCG	(temp. QRP !)
144.975	RA3FGV	KO92SO	140	V-Dipole	5		02/04	RW3BP	QRV!
145.250	RB9FA	LO88DA		Vertical	Omni 2	RV9FF	09/03		
432.128	S55ZNG	JN65UU	643	Horiz. Loop	Omni 0.1	S59DKS	02/04	S51KQ	QRT
432.280	UT5EC	KN78MM	180	4 x yagi	Omni				QRV?
432.361	UT5G	KN66LS	80	Dipole	N/S 5		01/04	UR5LX	
432.3...	UA9XEA	LO..					01/04	UR5LX	Proposal
432.400	UT3BW	KN29UA	365	crossdipole	Omni 2		01/04	UT3BW	QRT
432.400	OE3XMB	JN77TX	1154	11 Ele	337 5	OE3FFC	02/04	on .806	ex 432.800
432.400	GB3BSL	IO81XJ	252	4 x 3 el Yagi	90 250	GW8AWM	02/04		ex 432.934
432.401	SK2UHF	JP94WQ	445	2 x20 el coll	0/225 300e	SK2AT	02/04	SM6CEN	ex 432.875
432.401	UU9JJ	KN64RO					01/04	DL8HCZ	Proposal
432.402	OY6BEC	IP62OA		7 dBD	135° 50e		01/04	LA0BY	
432.403	EI2WRB	IO62IJ	248	5 el Yagi	95 250	EI9GO			Proposal
432.404*	F1ZNI	JN23MM	114	BigWheel	Omni 10	F1AAM	01/04	F6HTJ	
432.405*	SK1UHF	JO97CJ	65	Slot	Omni 30	SM1IUX	01/04	SM1MUT	ex 432.950
432.405*	PI7QH	JO22KH	20	3 dB Gain	Omni 2	PA0QHN	12/03		ex 432.905
432.406	OK0EO	JN89QJ	602	Ring Dipole	Omni 0.05/0.12	OK2VLX	12/03	Planned	ex 432.966
432.407	PI7YSS	JO32CD	60	Halo	Omni 2	PA0JAZ	07/03	PE2KP	ex 432.895
432.408	F5ZPH	IN87KW	165	4 Ele	135 20	F6ETI	01/04	F6HTJ	Temp QRT
432.410	DB0ZW	JN69EQ	825	Slot	Omni 1 TX	DC9RK	03/99	DJ3TF	ex 432.810
432.412	DB0JG	JO31GT	45	Clover Leaf	Omni 1 TX	DL3QP	08/03		ex 432.925
432.412	SK6UHF	JO67HE	175	Clover Leaf	Omni 10	SM6ESG	02/04	SM6CEN	Proposal
432.413*	F5XAG	IN93WC	550	2x10 el	N/NE 40	F5HPQ	01/04	F6HTJ	
432.414	DB0JW	JO30DU	238	2 x 11 el Yagi	45 50	DL9KAS	02/04	on 432.969	ex 432.975
432.415	DB0UBI	JO42GE	125	8el Coll	30 12	DD8QA			ex 432.920
432.416	PI7CIS	JO22DC	40		Omni 100	PA0CIS	02/04		ex 432.816
432.417	DB0LB	JN48NV	367	Cornerdipole	N/S 0.2TX	DK3PS			ex 432.945
432.417	DB0OS	JO40CW	745	Slot	270 0.1	DG6YW			ex 432.945
432.417	HG3BUA	JN96CC	585	Slot	Omni 1TX		02/04	HA5SHF	ex 432.946
432.417	OH9UHF	KP36OI	307	9 dBd gain	200 70		01/04	OH6DD	ex 432.945
432.418*	F1ZQT	IN95OX	80	Big Wheel	Omni 1	F1MMR	01/04	F6HTJ	
432.420*	F5XAS	JN12BL	2400	3 el	0 50	F6HTJ	02/04	F6HTJ	
432.420	DB0VC	JO54IF	300	4xDQ	Omni 10	DL8LAO	02/04		ex 432.990
432.423	PI7HVN	JO22WW	50	Horizontal	Omni 0.5	PE1HUE	07/03	PE2KP	ex 432.873
432.425	DB0YI	JO42XC	480	Big wheel	Omni 4 TX	DL4AS	08/03		ex 432.900
432.425	EA6UHF	JM08PV	470	Halo	Omni 10	EA6FB	temp. QRT		ex 432.918
432.425	LY2WN	KO25GC	96	2 x dipole	Omni 5	LY2WN	06/01	G3UUT	Planned
432.428	HG7BUA	JN97KR	700	Slot	Omni 1	HG5ED	Proposal		ex 432.930
432.429*	HB9OK	JN45MW	1601	Halo	Omni 3.5	HB9DUR	02/04		ex 432.938,5
432.430	GB3MLY	IO93EO	600	6 el Yagi	150 40	G3PYB	11/03		ex 432.910
432.431	9A0BUH	JN85JO	489	V dipole	Omni 1	9A2MP	03/95	QRT	ex 432.847
432.432*	HB9F	JN36XN	3581	Cornerrefl.	N/S 15	HB9MHS	02/04		ex 432.984
432.432	OH6UHF	KP13GM	55	3xBig Wheel	Omni 7	OH6UH	01/04	OH6DD	ex 432.840
432.434	DB0LBV	JO61EH	234	2xSlot	Omni 2TX	DL1LWM	08/03		ex 432.845
432.435*	OH5SHF	KP30HV	145	Turnstile	Omni 1		01/04	OH6DD	ex 432.935
432.436	F5XAZ	JN06KN	144	Big wheel	Omni 50	F5EAN	01/04	F6HTJ	ex 432.886

432.437	LA1UHF	JO59MS	380	X300 (Vert)	Omni	30	LA4PE	01/04	LA0BY	Temp QRT
432.438	HG6BUA	KN07AU	1015	Slot	Omni	1	HG5ED	02/04	HA5SHF	temp. QRT
432.440*	SK7MH	JO86GP	45	Alford slot	Omni	100e		02/04	SM6CEN	ex 432.940
432.440	DB0RG	JO51GO	730	V-Dipole	Omni	1				ex 432.940
432.441	F5XBA	JN18KF	166	4xHB9CV	Omni	10	F2AI	01/04	F6HTJ	ex 432.830
432.441*	LA7UHF	JP20LG	30	4 el Yagi	0	200	LA6LU	02/04	LA0BY	ex 432.830
432.442	CT....						Proposed	06/04	CT1DDN	
432.443	OH2UHF	KP20BB	65	2x 10 dBi	20/225	60		01/04	OH6DD	ex 432.852
432.444	DB0KI	JO50WC	925	Dipole	Omni	10	DC9NL	01/04	on 432840.8	ex 432.840
432.445	G33LER	IP90JD	104	12 el Yagi	165	675	GM4IPK	06/02	G3UUT	QRT
432.446	OK0EB	JN78DU	1084	Mini Wheel	Omni	0.03/0.01	OK1APG	01/04	on 432965.7	ex 432.970
432.447	S55ZRS	JN76MC	1219	Slot Dipole	Omni	1	S50M	02/04	S51KQ	ex 432.950
432.447	DB0IH	JN39ML	630	Big wheel	Omni	1	DC8DV	08/03		ex 432.950
432.448	S55ZCE	JN76OL	554	GP (V)	Omni	0.07	S51KQ	02/04	temp QRT	ex 432.980
432.449*	OZ1UHF	JO57FJ	150	Big wheel	Omni	10	OZ9NT	01/04	OZ2TG	
432.450	DL0UB	JO62KK	120	Malteser	Omni	10 TX	DL7ACG	02/04		ex 432.850
432.450*	I5WBE/B	JN53LK	300	2x7 el Yagi	15/260	6	I5WBE	01/04	DL8HCZ	ex 432.450
432.452	OK0EC	JO60CF	778	10 el Yagi	90	1	OK1VOW	12/03		ex 432.980
432.453	GB3ANG	IO86MN	370	9 el Yagi	170	100	GM4ZUK	08/03		ex 432.980
432.455	SK3UHF	JP82XO	175	4xD-quad	Omni	10	SM3AFT	02/04	SM6CEN	ex 432.855
432.456	DB0GD	JO50AL	926	Slot	N/S	1 TX	DG6ZX			ex 432.800
432.460*	SK4BX/B	JO79LH	270	4x10el log p	Omni	43	SM4RWI	02/04	ex SK4UHF	ex 432.905
432.463	LA2UHF	JP53EG	710	10 Ele Yagi	15	300e	LA1K	02/04	LA0BY	ex 432.870
432.465	DF0ANN	JN59PL	630	Big wheel	Omni	1 TX	DL8ZX			ex 432.965
432.466*	OZ4UHF	JO75KC	130	Clover leaf	Omni	10	OZ1HTB	01/04	OZ2TG	
432.467	ON4UHF	JO20ET	180	Clover leaf	Omni	0.5	ON4LC	08/03		ex 432.994
432.468	LA6UHF	KP59AL	70	15 el Yagi	210	40	LA4OO	02/04	LA0BY	ex 432.865
432.470	OK0EZ	JN79VV	350			1	OK1DXF	01/04		ex 432.871
432.470	GB3MCB	IO70OJ	320	4 e agi 45/	12		G3YJX	02/04	G3YJX	QRT
432.471	OZ7IGY	JO55VO	93	Omni+ 9el	Omni/45	30/300	OZ7IS	01/04	OZ7IS	ex 432.930
432.473	OE3XAA	JN88BA	834			0.2	OE1BKW	09/99	OE1MCU	QRT
432.475*	ES0UHF	KO18DN		Ring	Omni	12	ES2NJ	11/03		ex 432.835
432.475	DB0SGA	JN69KA	1024	4 x 9 Yagi	Omni	5TX	DJ4YJ			ex 432.975
432.475	HG1BUA	JN87FI	725	Slot	Omni	5	HA1YA	02/04	HA1YA	ex 432.975
432.477	DB0ABG	JN59WI	522	Big Wheel	Omni	1 TX	DJ3TF			QRT
432.478	LA3UHF	JO38RA	12	13 el Yagi	180	50	LA3BAA	02/04	LA0BY	ex 432.880
432.478	OM0MUA	JN88NE	570	Dipole	90/270	0.08	OM3ID			ex 432.888
432.479	SR5UHW	KO02OF	130	Turnstile	Omni	0.25	SP5TAT			ex 432.982
432.480	LA8UHF	JO59FB	30	8 el Yagi	90/180	50	LA6LCA	02/04	LA0BY	ex 432.820
432.482	OZ2ALS	JO45UB	28	4 x dipole	Omni	40	OZ9DT	01/04	OZ2TG QRT	ex 432.983
432.482	GB3NGI	IO65VB	518	12 el Yagi	125	250	GI6ATZ	02/04	GI6ATZ	Proposal
432.485*	LA4UHF	JO29PJ	75	10 el Yagi	200	50	LA9NQ	02/04	LA0BY	ex 432.890
432.488	DB0AD	JO40AQ	690	V-Dipole	Omni	1		08/03		ex 432.918
432.489	OK0EA	JO70UP	1355	2 x 15 el Yagi	S+W	3	OK1AIY	01/04	on 432936.6	
432.489*	SK7MHL	JO65OR	100	Alford slot	Omni	40	SM7ECM	02/04	SM7ECM	ex 432.970
432.490	OH7UHF	KP32TW	215	6 dBd	225	15/1.5/15		01/04	OH6DD	ex 432.875
432.490	GB3SUT	IO92CO	270	2 x 8 el Yagi	0/135	10	G8XGG	08/03		ex 432.885
432.831	IQ0RI/B	JN62MH	1010	3 Ele. Yagi	330	40e			QRV? New QRG?	
432.886.7	OK0EP	JO80OC	1505	2 x 4 el Yagi	W+SE	2x3	OK1VPZ	02/04	will stay on old QRG!	
432.908	HG8BUA	KN06PW	85	9 el Yagi	270	5	HA8MV	02/04	HA8MV	New QRG?

This list is based on the official IARU beacon list. Many additions, updates and corrections were made.

This is a list for European Beacons. Anyhow a few other "special" beacons from other continents were included.

* on 70cm means that the beacon is already qrv on the new QRG!

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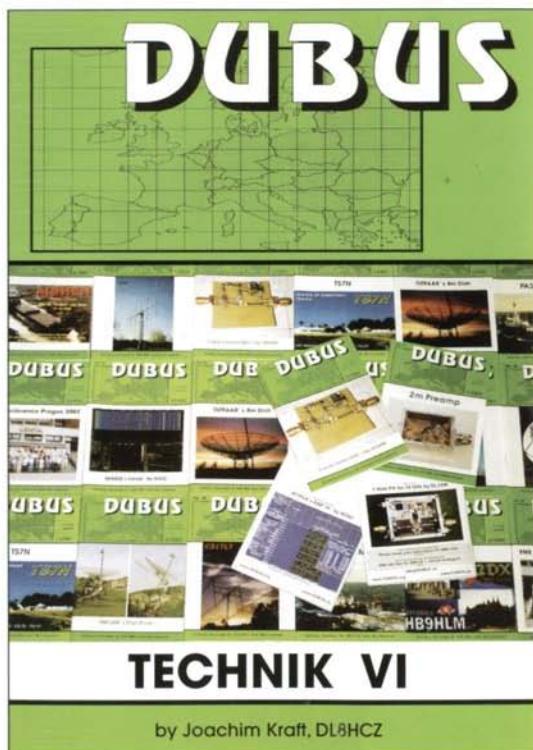
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